WG3: SXR and EUV Response to Flare Heating

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- Originally proposed as the RHESSI-SDO group
- Actual instruments presented included: SDO/AIA, SDO/EVE, SDO/HMI, Hinode/XRT, FOXSI, SOHO/ LASCO, Kanzelhohe, TRACE, GOES, SORCE, PROBA2/LYRA, ...

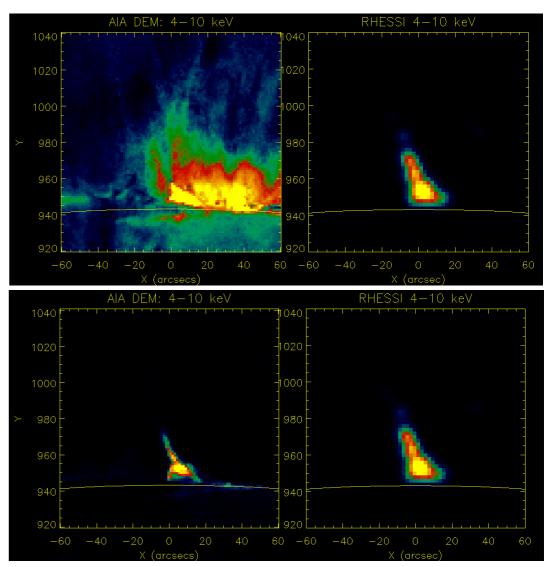
- Session 1: DEMs (Aswathi, Motorina, Su, Hernandez-Perez)
- <u>Session 2</u>: Hydrogen emission (Joint with WG1&WG6; Kleint, Kotrc, Kasparova, Heinzel)
- <u>Session 3</u>: Chromospheric Evaporation (Joint with WG1; Ning, Battaglia, Gomory, Fleishman)
- <u>Session 4</u>: Active Regions (Holman, Ishikawa, Thalmann, Veronig)
- <u>Session 5</u>: QPP & Energetics (Hayes, Dennis, Aschwanden, Warmuth)
- Session 6: Misc. (Kawate, Kuhar, Joshi, Caspi, Milligan)

Improved DEM solutions - Su

Problems: limited/incomplete knowledge of lines/filters, inaccurate temperature response and errors on data, spatial resolution

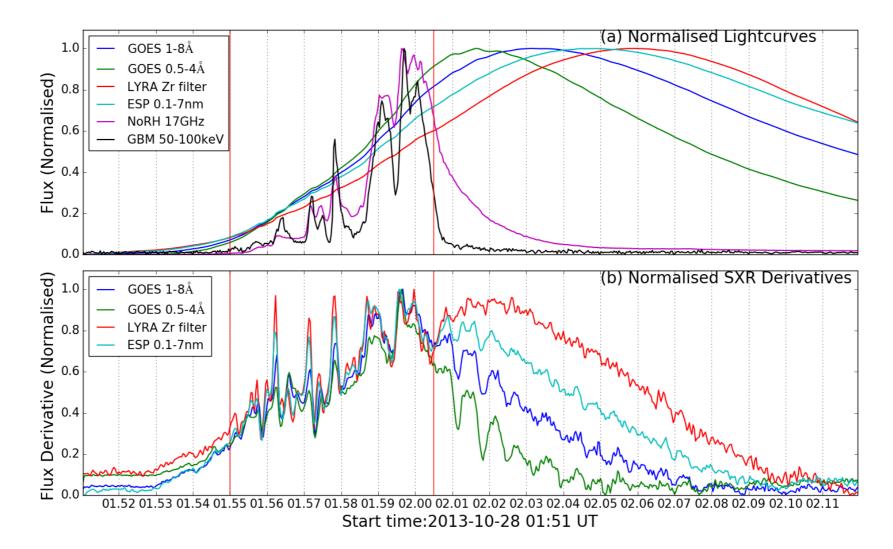
Solution: iteratively modify DEM peaks at high temperatures and change the order in which the code calculates the DEM

Old method



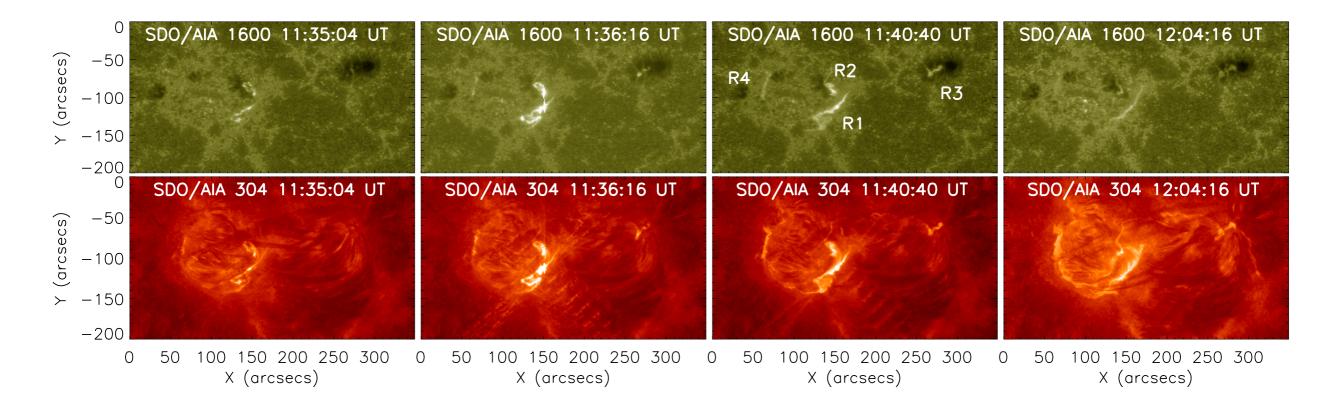
New method

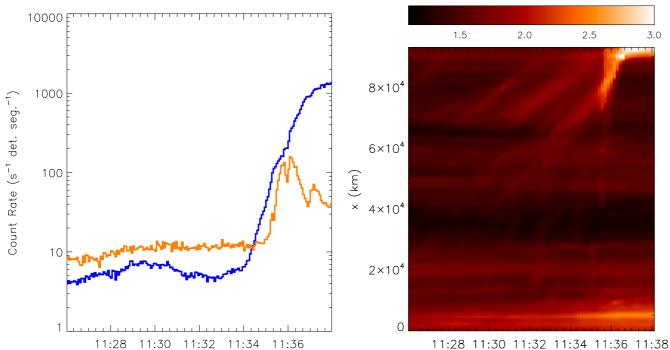
Multi-wavelength Detection of QPPs - Hayes



Fluctuations correlated across multiple instruments/ wavebands in both nonthermal (~20s) and thermal (~60s) emission. Nonthermal pulsations due to episodic reconnection. Thermal ones due to sausage/kink mode oscillations.

Studies of a Circular Ribbon - Hernandez-Perez

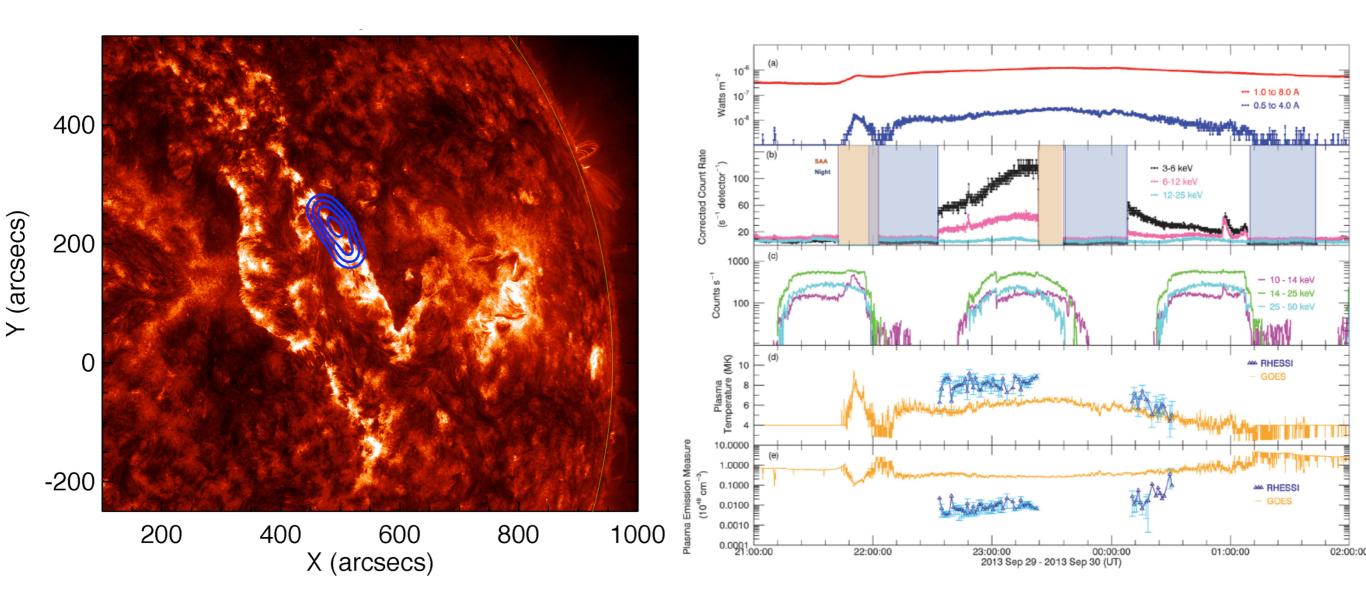




Time-slice plots showed blobs propagating towards the outer footpoints prior to flare onset - but no HXR

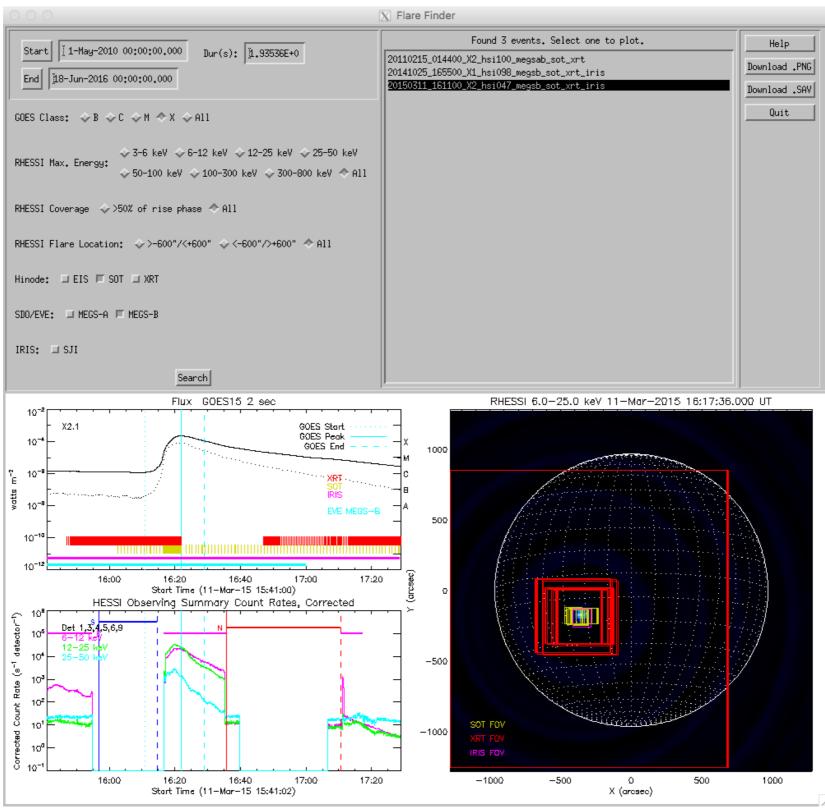
11:28 11:30 11:32 11:34 11:36 11:38 Start Time (29-Jan-15 11:26:02)

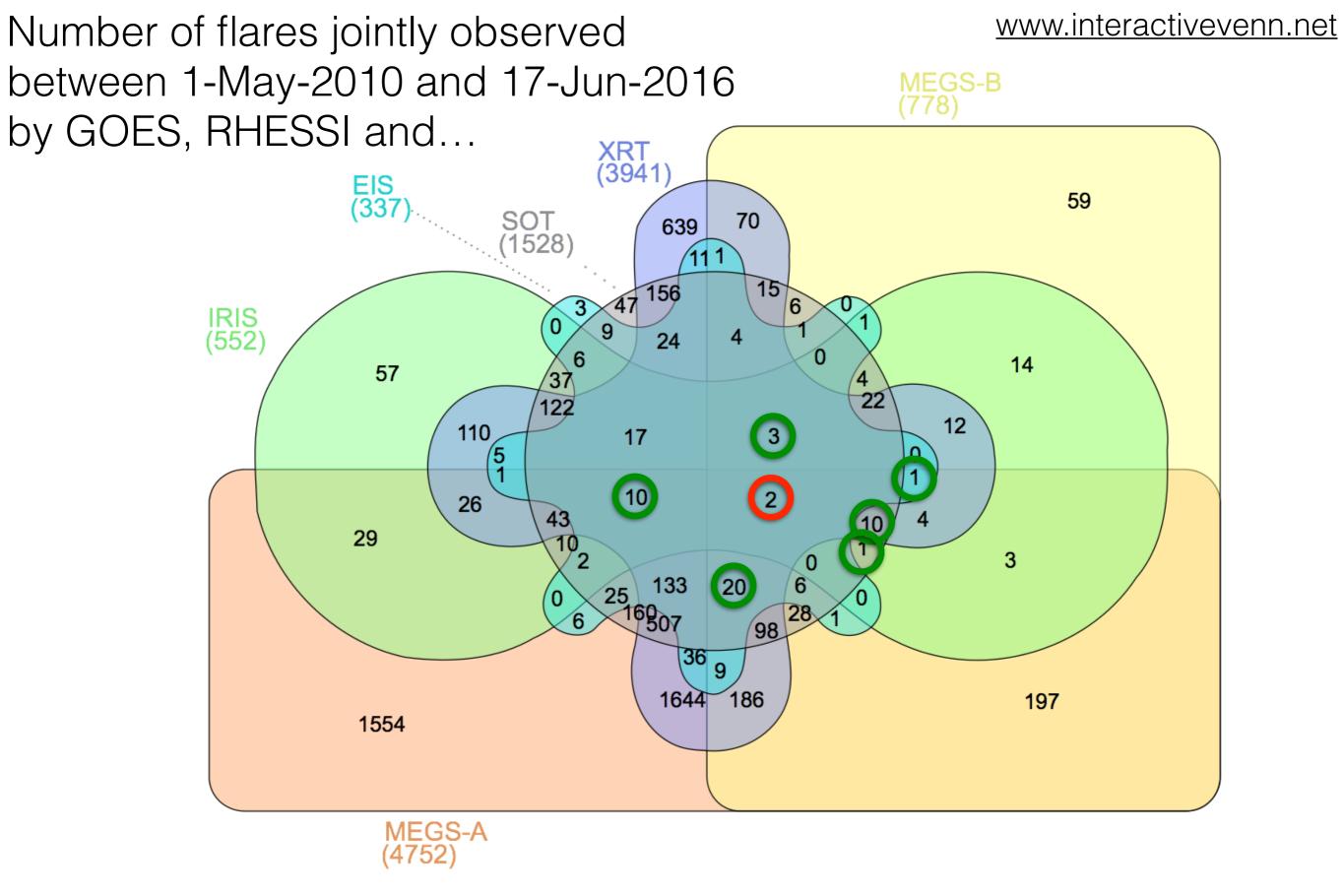
X-ray emission from Filament Eruptions Outside ARs - Holman



Filament eruption resulted in hot (6-8 MK) plasma due to secondary reconnection with strong dipole field

Widget-based search tool now in SSWIDL - Milligan (IDL>solar_flare_finder).





Total number of GOES flares = 12,254 Total number of RHESSI flares = 6,761 O=all 8 instruments
O=any 7 instruments