



University
of Glasgow

LOFAR – Low frequency Array

Eduard Kontar

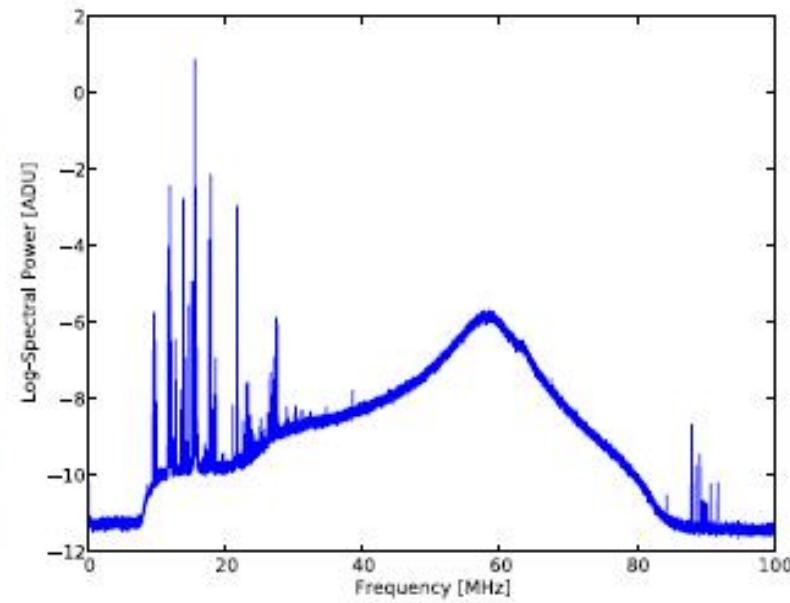
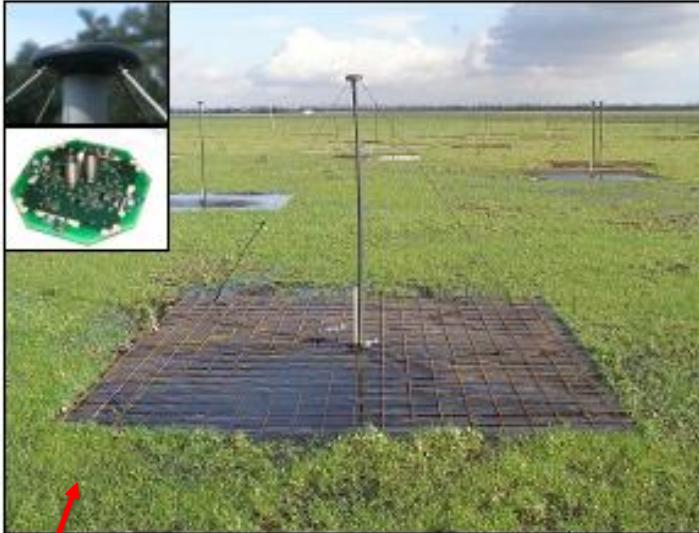
*School of Physics and Astronomy
University of Glasgow, UK*

July 30, 2016

RHESSI 15 Workshop, Graz



Low Frequency Array



| System characteristic | Options | Values | Comments |
|--|-------------------|-------------|-------------------------------------|
| Frequency range | Low-band Antenna | 10–90 MHz | |
| | | 30–90 MHz | With analog filter |
| | High-band Antenna | 110–190 MHz | 200 MHz sampling (2nd Nyquist zone) |
| | | 170–230 MHz | 160 MHz sampling (3rd Nyquist zone) |
| | | 210–250 MHz | 200 MHz sampling (3rd Nyquist zone) |
| Number of polarizations | | 2 | |
| Bandwidth | Default | 48 MHz | 16-bit mode |
| | Maximum | 96 MHz | 8-bit mode |
| Number of simultaneous beams | Minimum | 1 | |
| | Maximum | 244 | 16 bit mode, one per sub-band |
| | Maximum | 488 | 8 bit mode, one per sub-band |
| Sample bit depth | | 12 | |
| Sample rate | Mode 1 | 160 MHz | |
| | Mode 2 | 200 MHz | |
| Beamformer spectral resolution | Mode 1 | 156 kHz | |
| | Mode 2 | 195 kHz | |
| Channel width (raw correlator resolution) | Mode 1 | 610 Hz | |
| | Mode 2 | 763 Hz | |

Van Haarlem et al, A&A, 2013

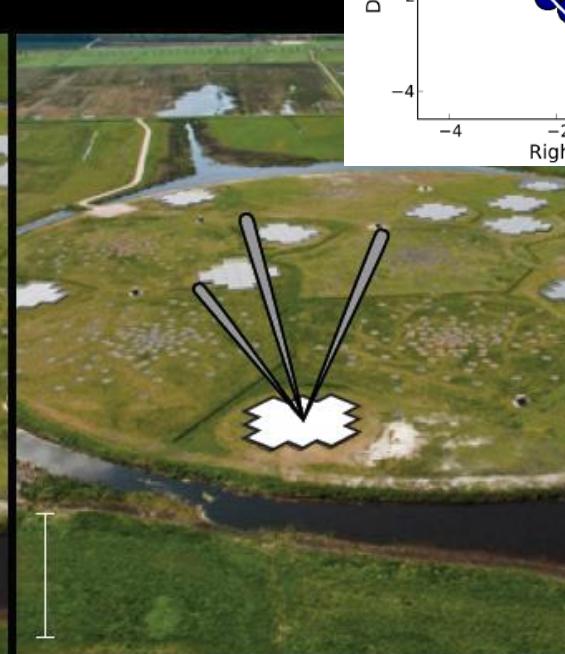
van Leeuwen - From Stappers et al. 2011



Element Beam

or

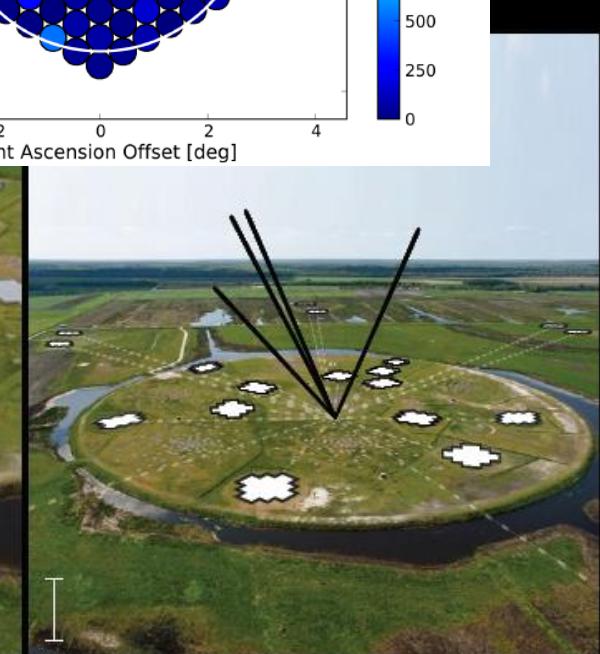
Tile Beam



Station Beam

or

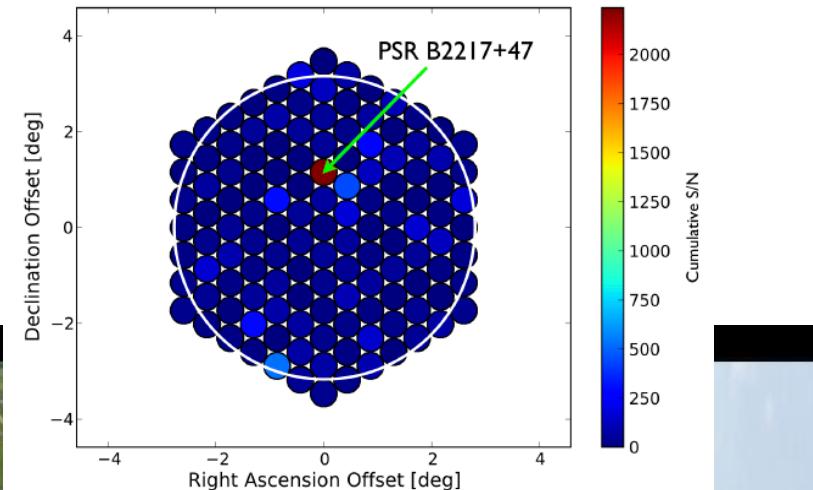
Sub-Array Pointing

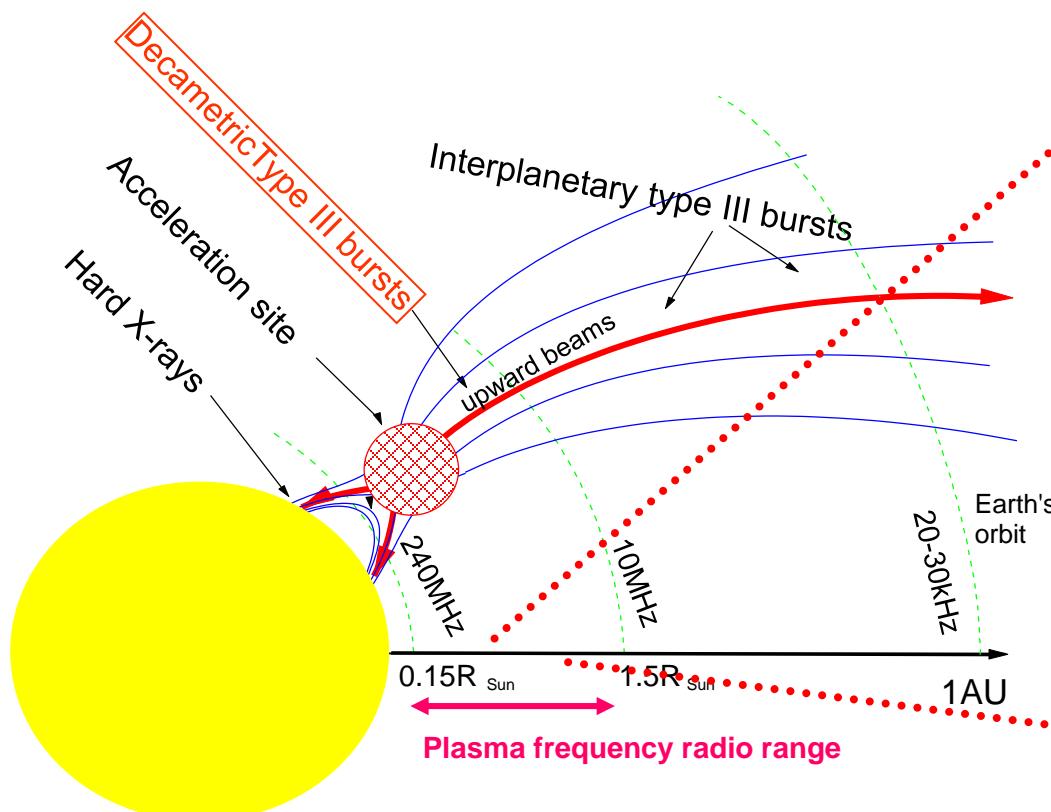


Array Beam

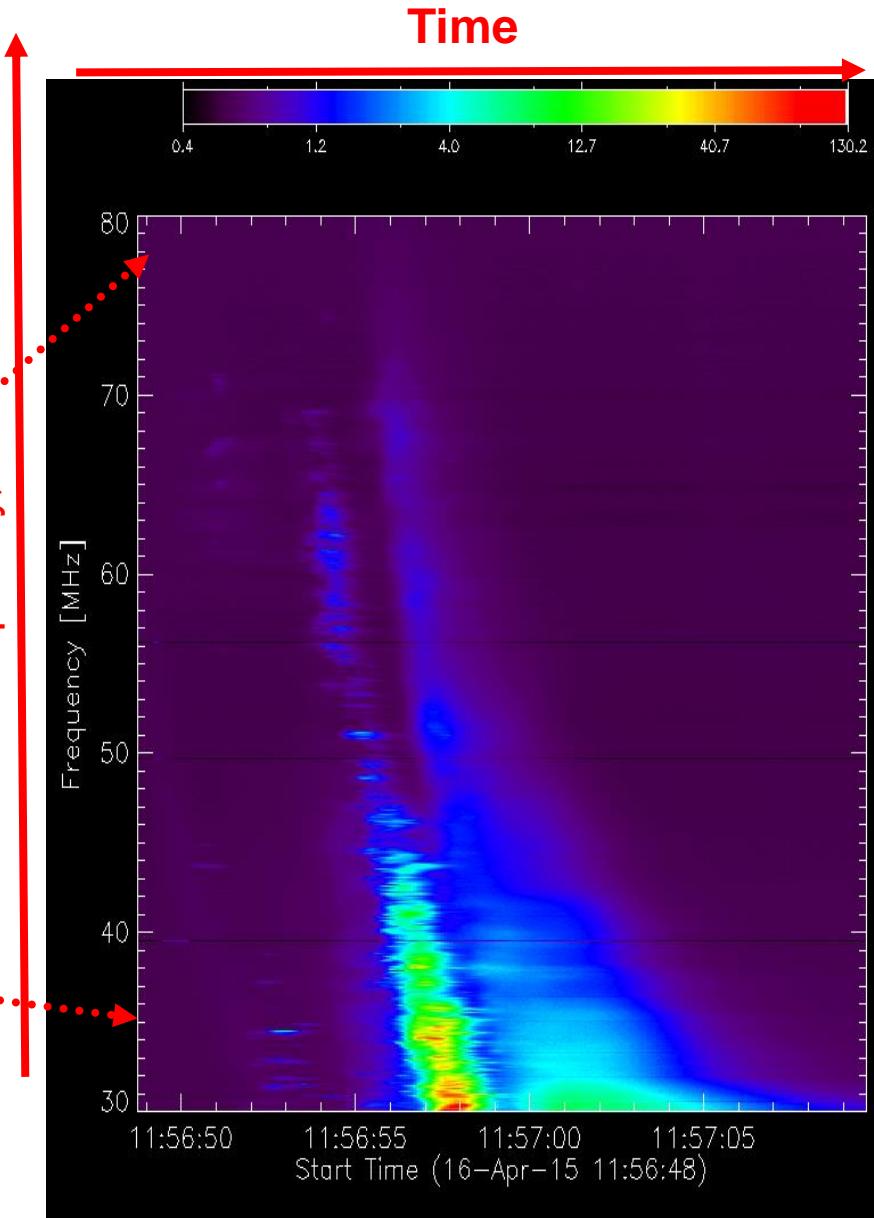
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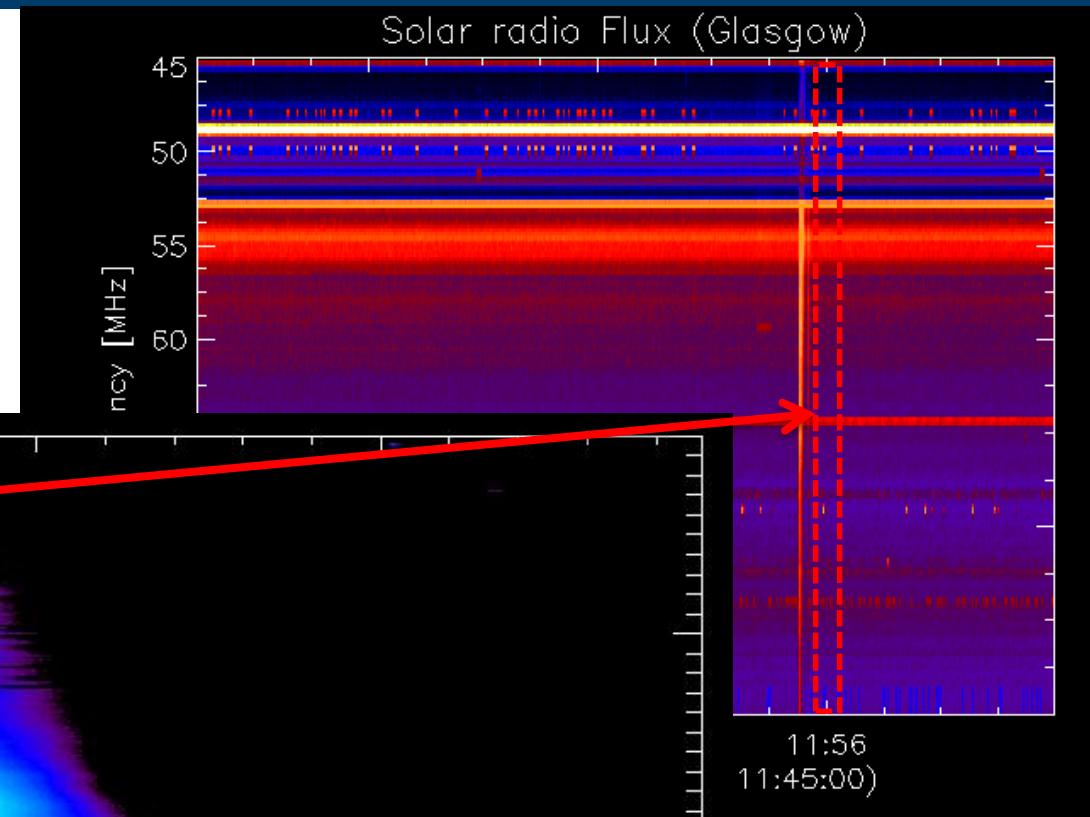
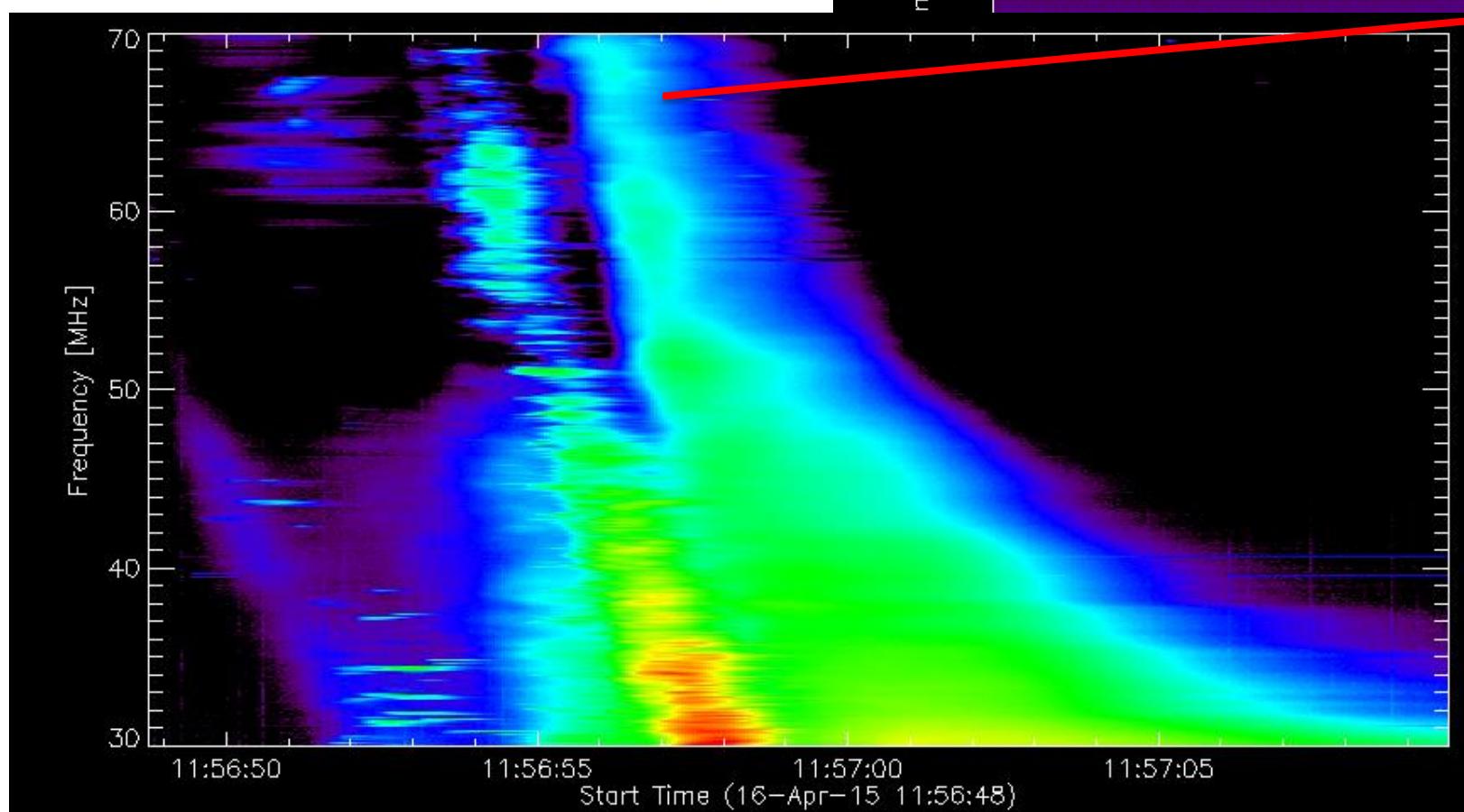
Tied-array Beam

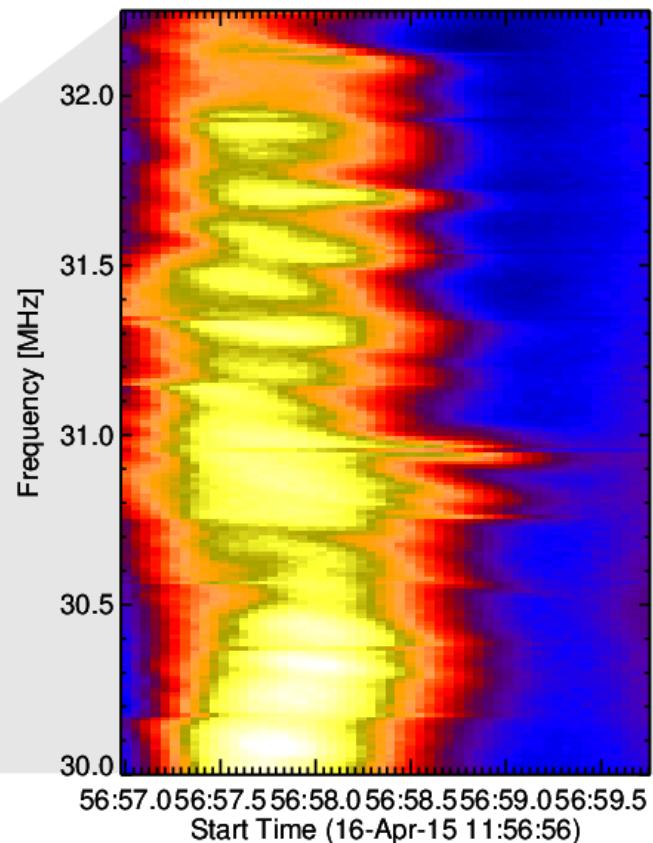
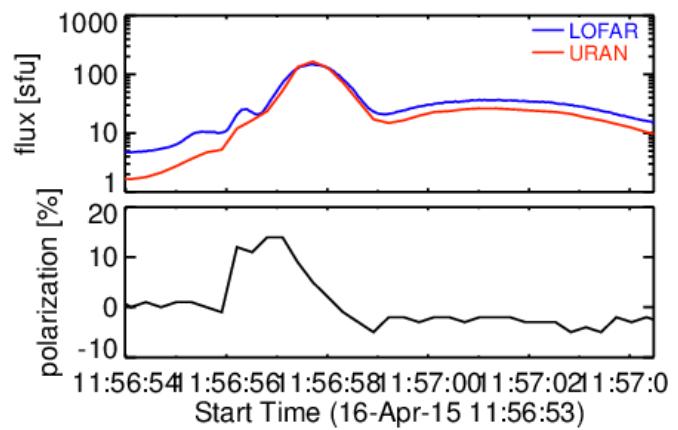
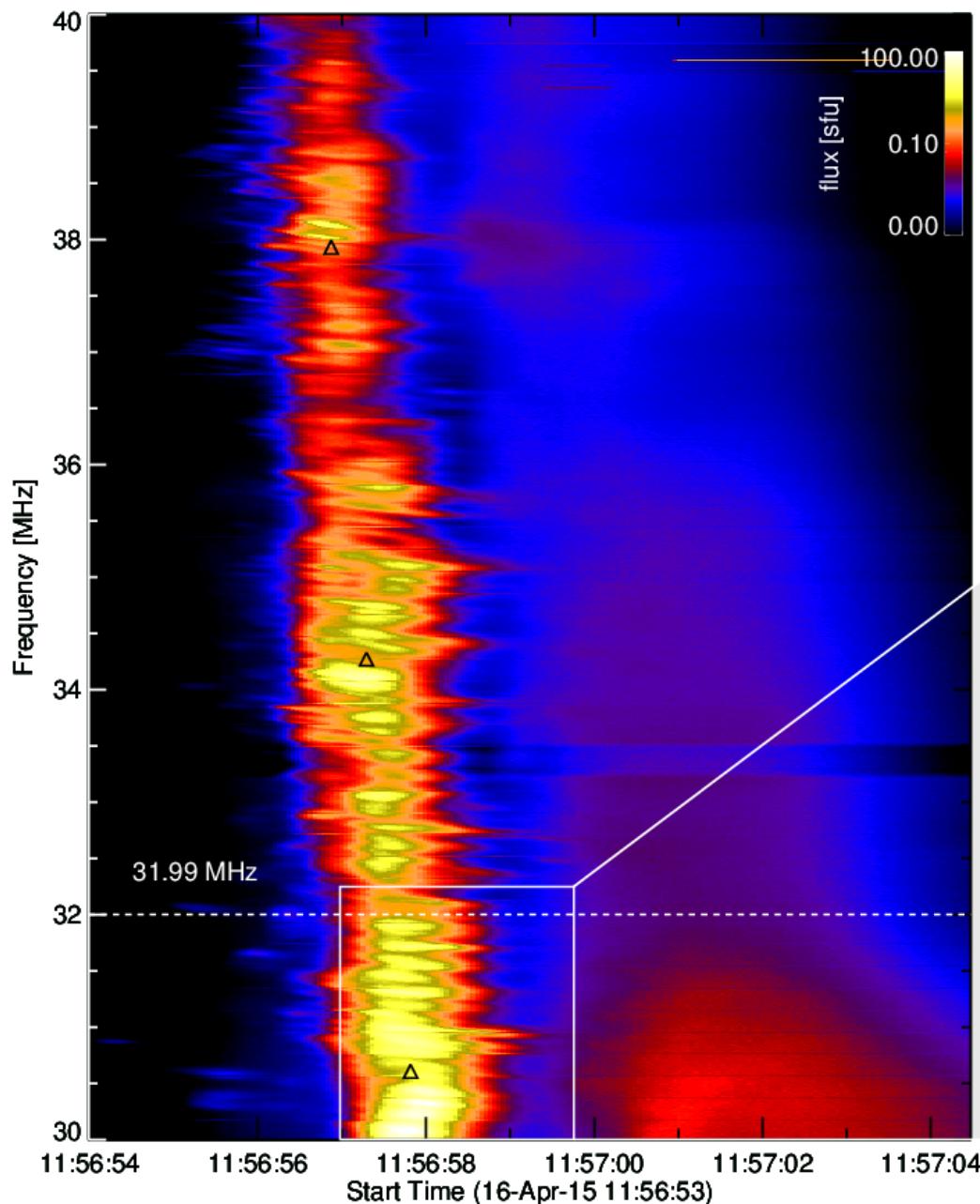


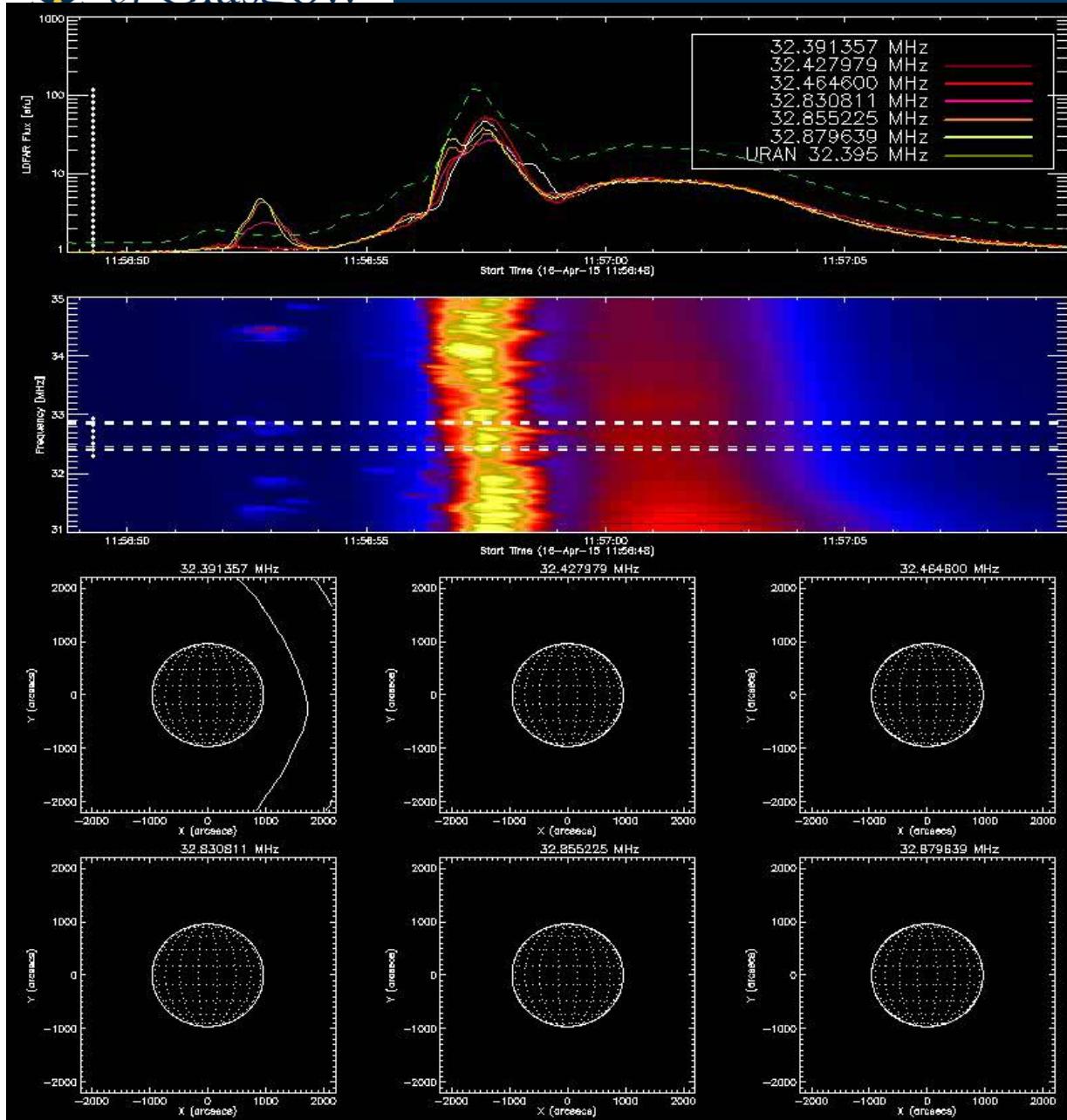


$$\nu_p = \sqrt{\frac{n_e e^2}{\pi m_e}}, \quad \leq \text{plasma frequency}$$









12 kHz bands
simultaneous
imaging between
30 and 80 MHz

C. Vocks, G. Mann, and F. Breitling

LOFAR observations of the quiet solar corona

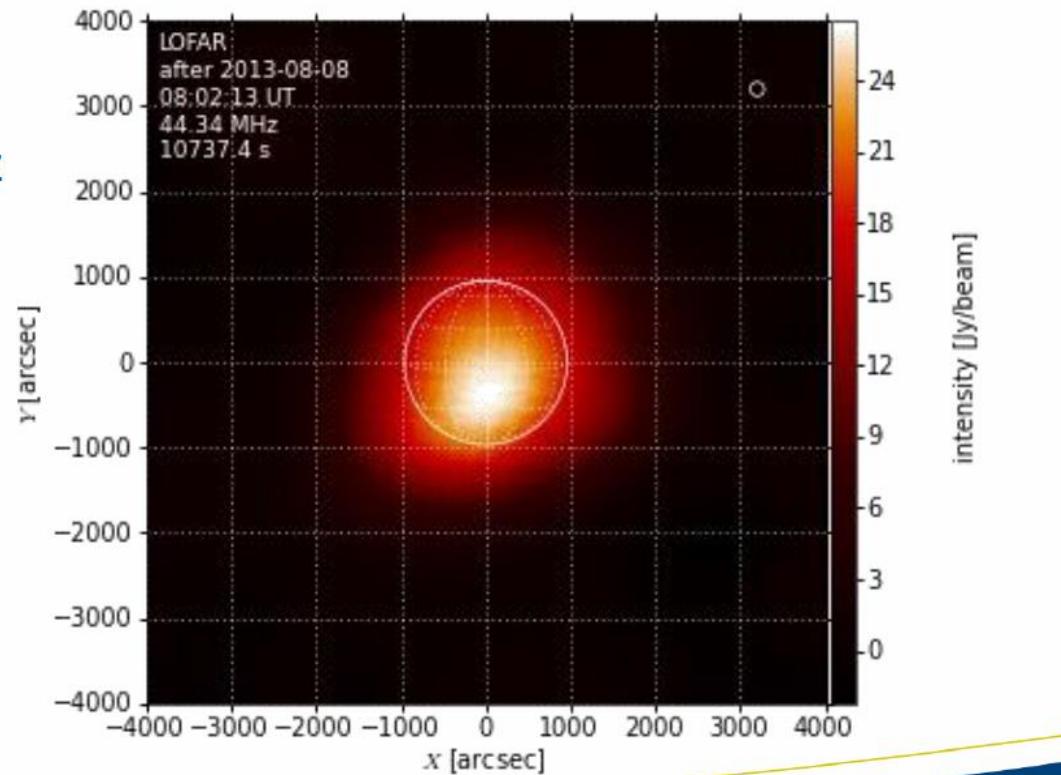


Solar corona

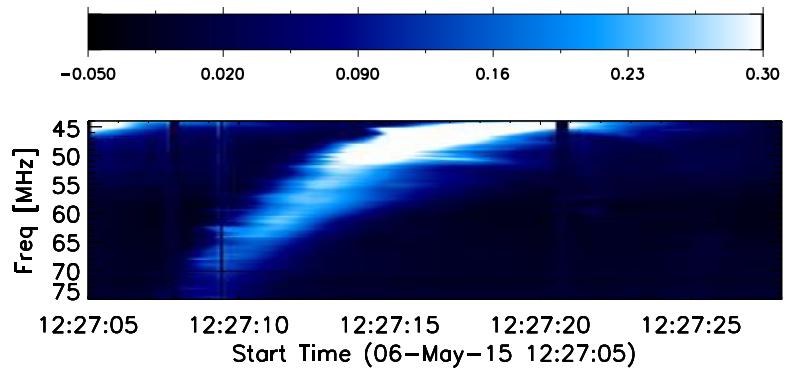


Image:

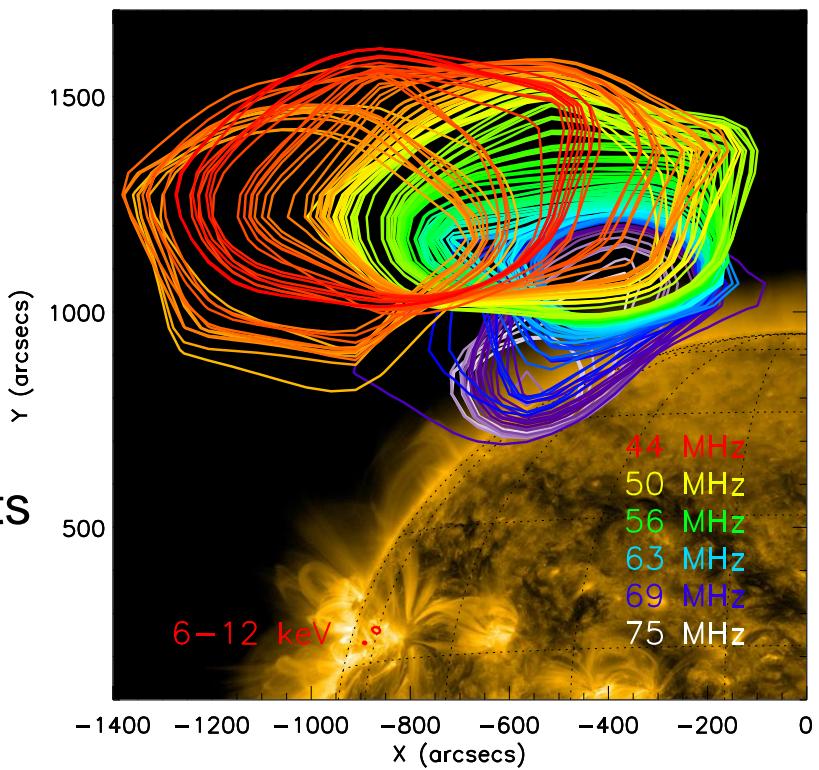
- 44 MHz
- 3 h



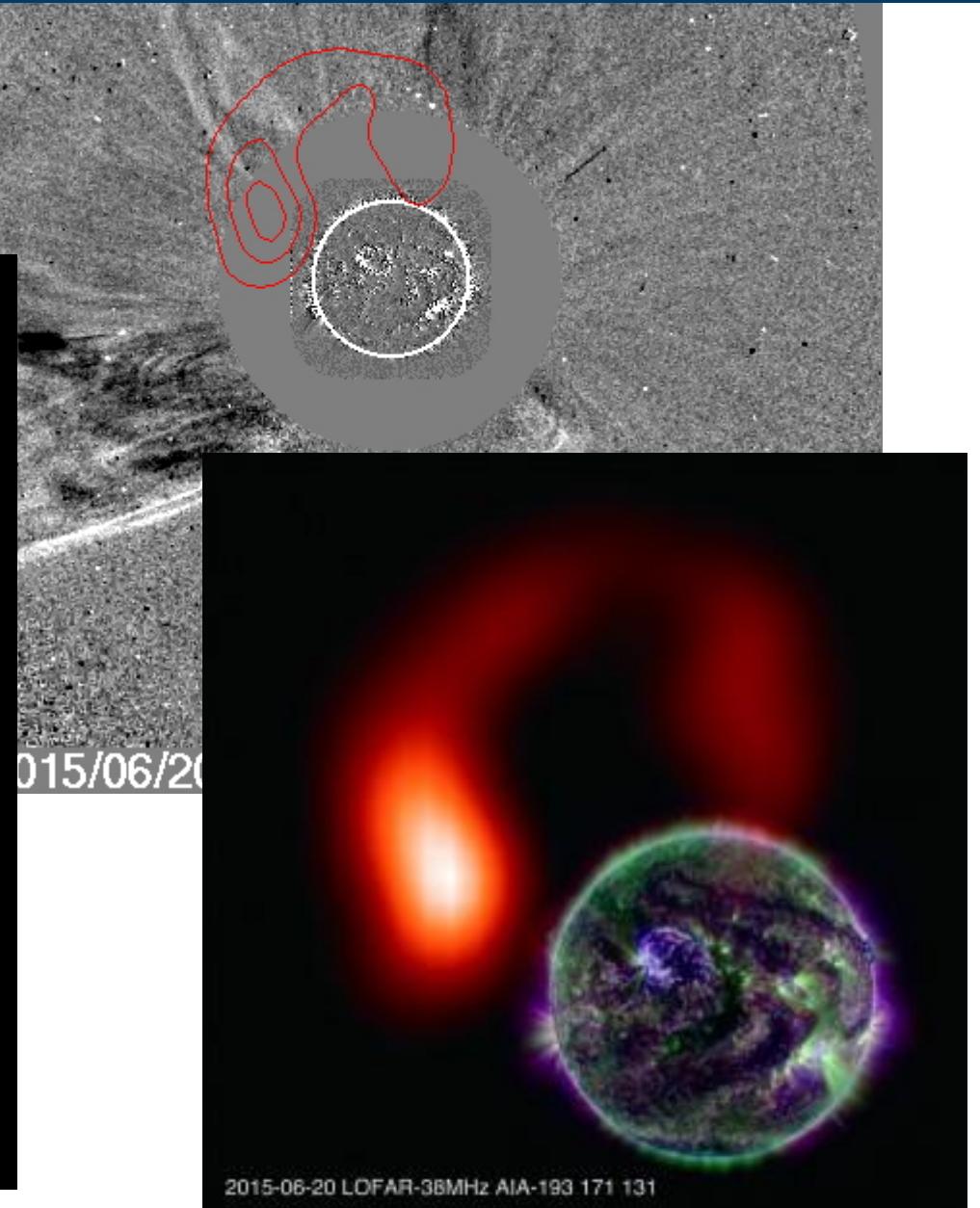
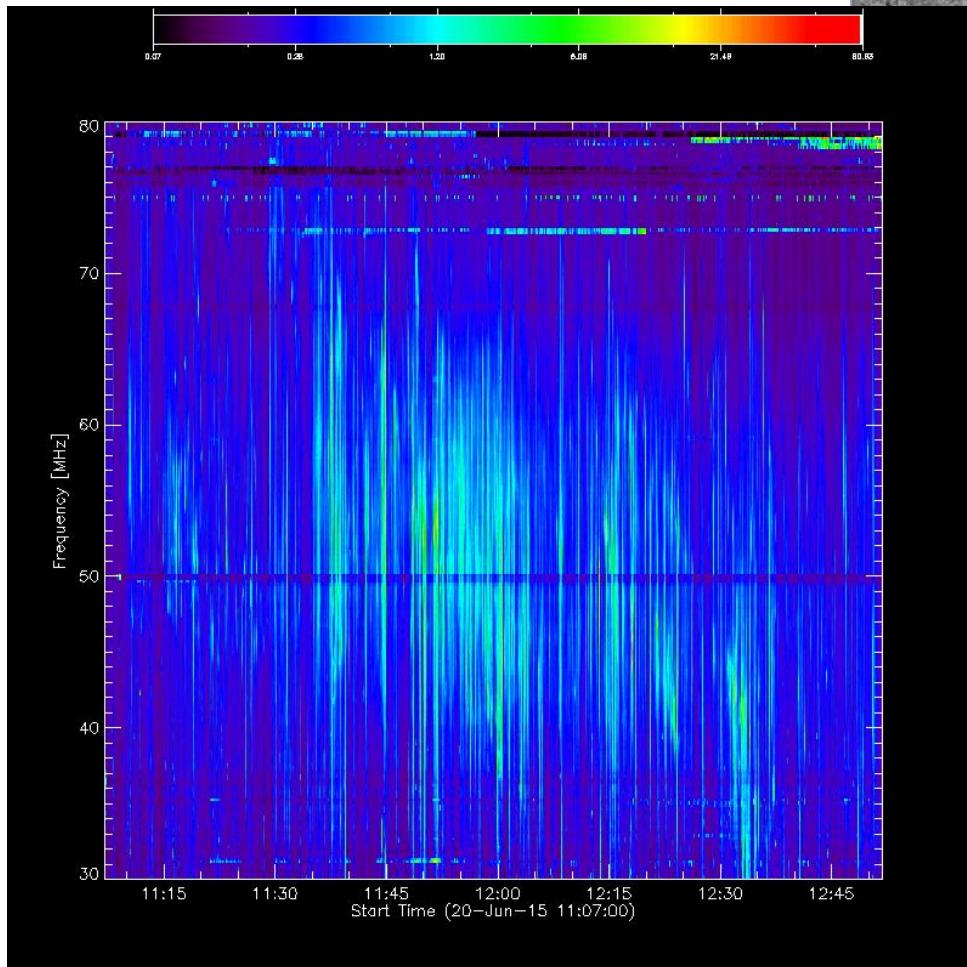
D. Morrison et al, 2014
Type III bursts



LOFAR observations of U/J bursts
H. Reid and E Kontar,

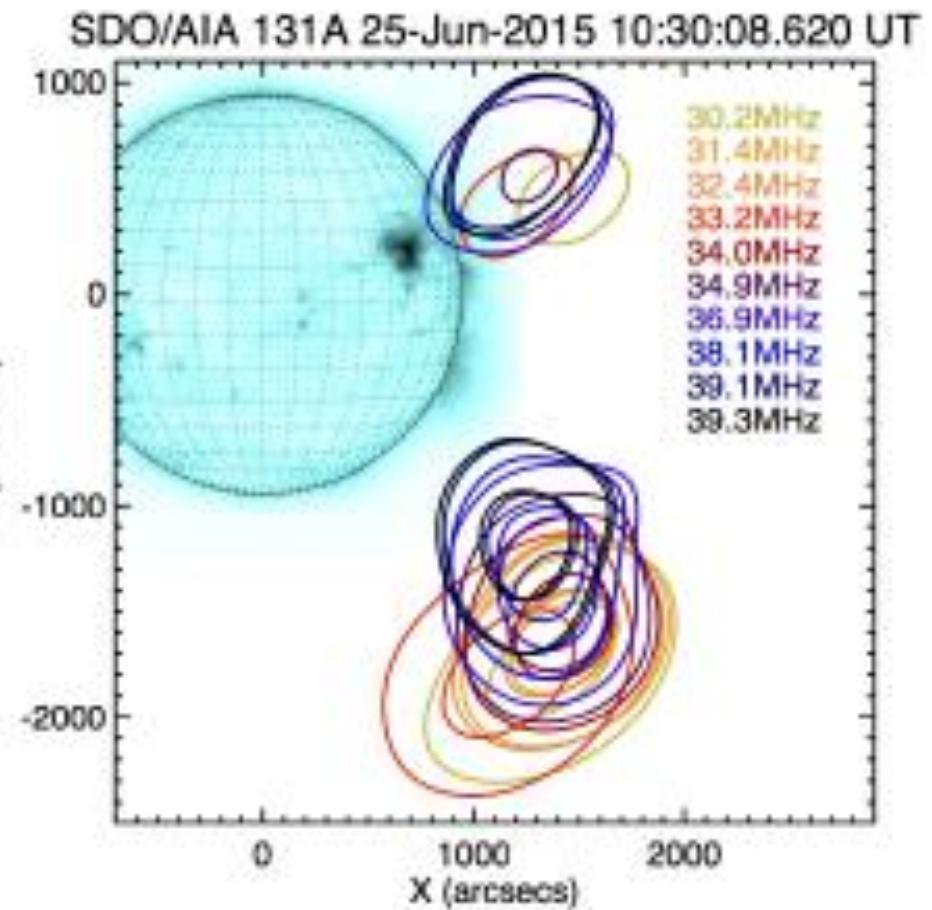
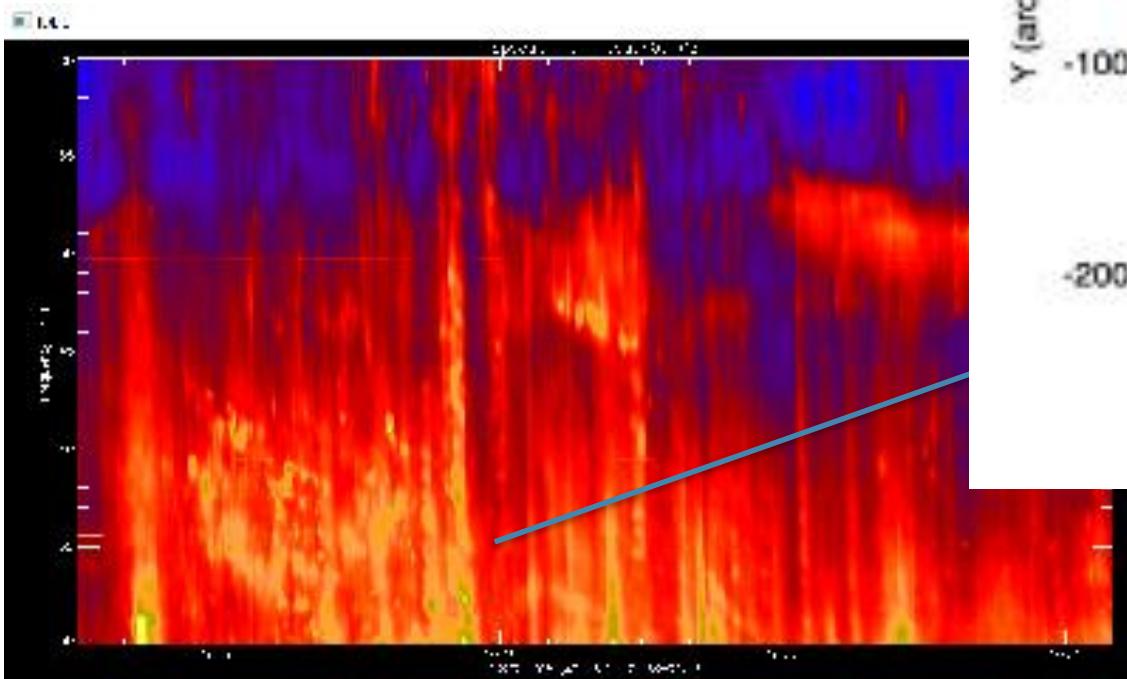


A. Kuznetsov, et al
Type IV and CMEs



L. Glesener, et al

CMEs, jets, type II and type IIIs...



- We had successful observations (Type III, (U, J), I, II, IV, V and fine structures...)
- Still requires software development.
- You can apply for LOFAR time (next call in Sept 2016)

