

# Quasi-Periodic Pulsations

## 14 May 2013

Brian Dennis  
RHESSI Workshop  
Graz, Austria  
27 – 30 July, 2016

# Long-lived X-ray Quasi-Periodic Pulsations in the X-class Solar Flare on 2013 May 14

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Andrew Inglis<sup>2</sup>, Jack Ireland<sup>3</sup>, Tongjiang Wang<sup>2</sup>, and  
Gordon D. Holman

(Solar Physics Laboratory, NASA Goddard Space Flight Center, Greenbelt, MD)

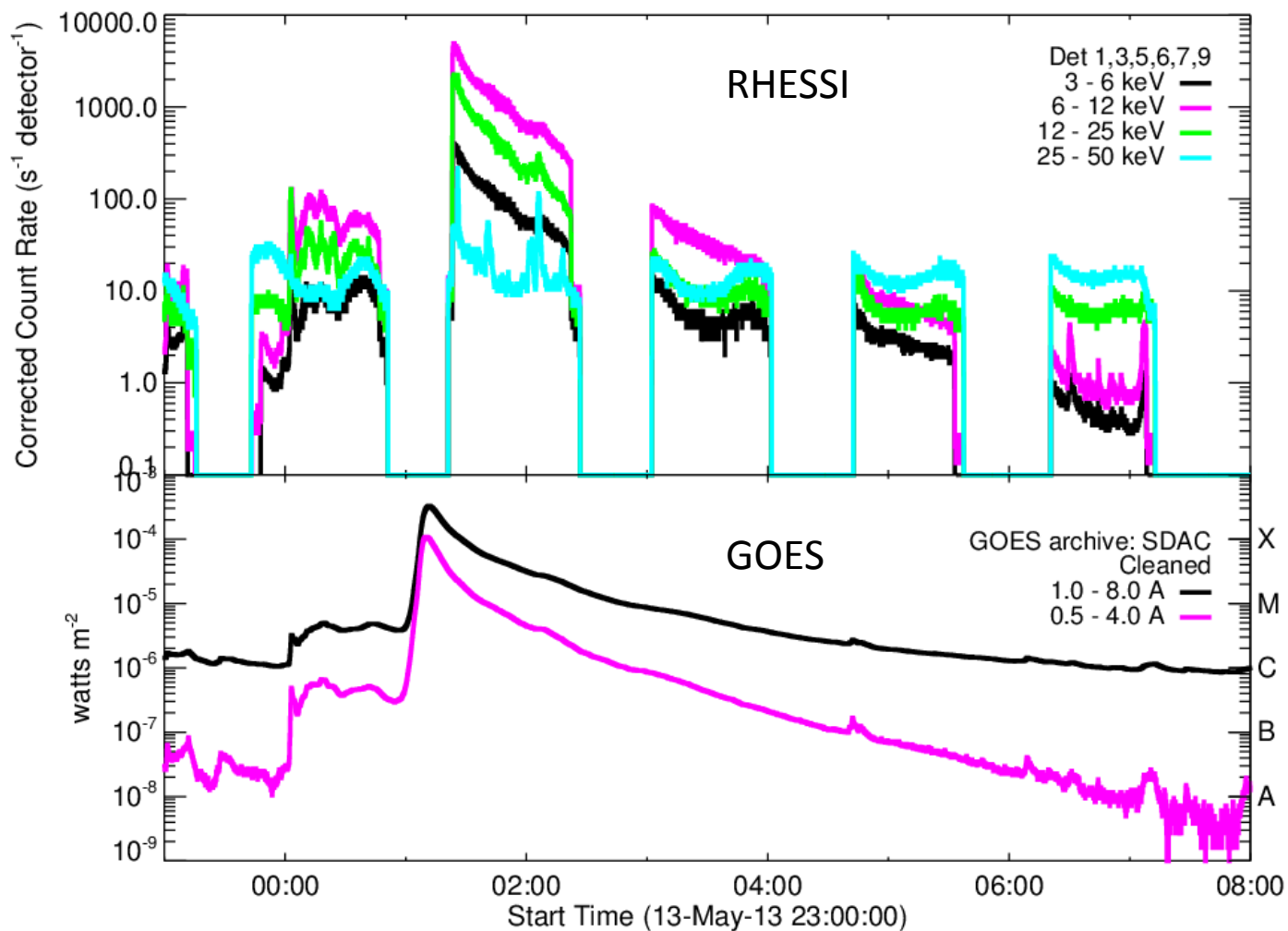
Laura A. Hayes and Peter T. Gallagher  
(School of Physics, Trinity College Dublin, Dublin 2, Ireland)

[1brian.r.dennis@nasa.gov](mailto:brian.r.dennis@nasa.gov)

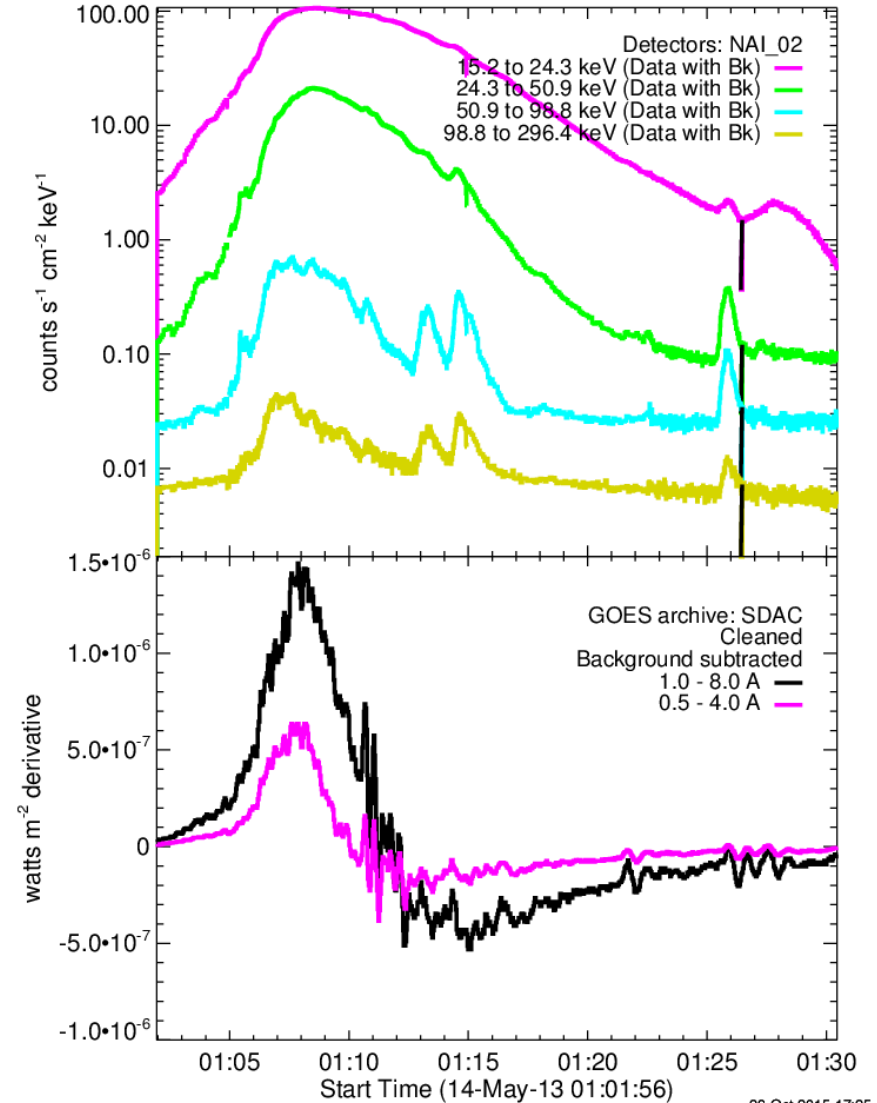
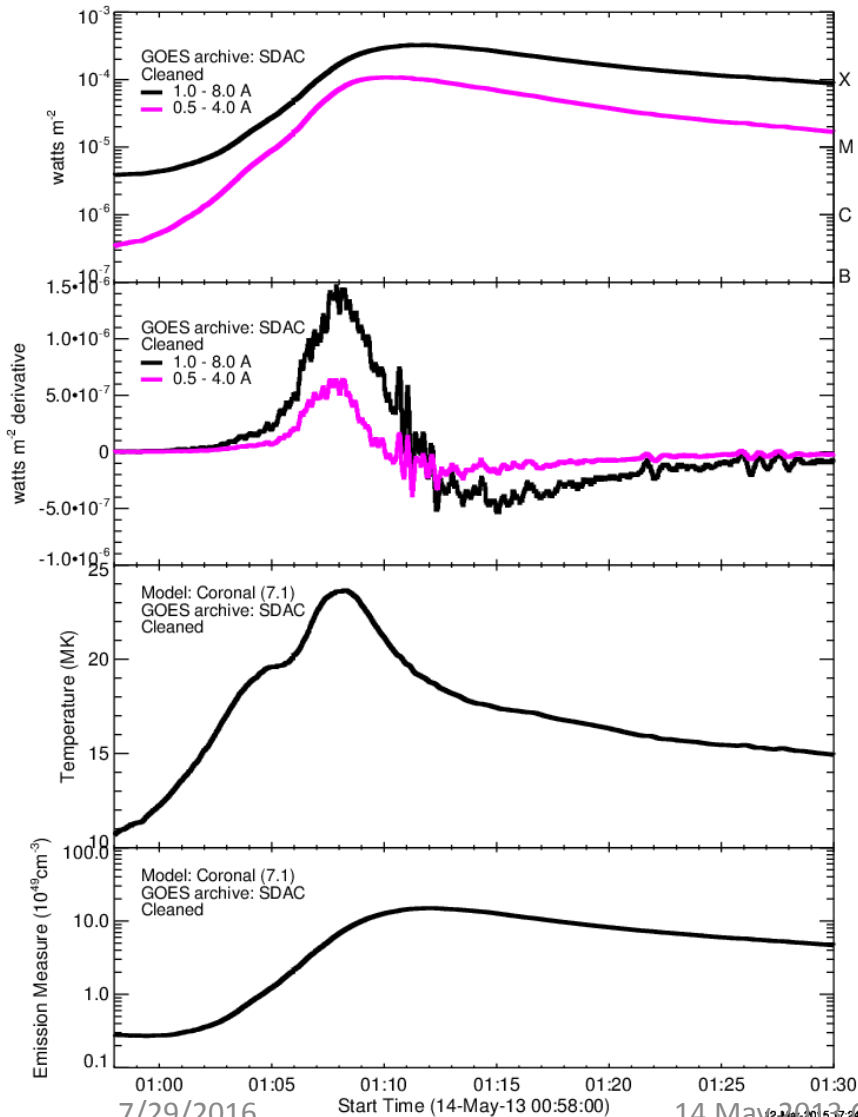
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<sup>3</sup>ADNET Systems, Inc. at NASA Goddard Space Flight Center, Greenbelt, MD 20771, USA

# 14 May 2013, RHESSI & GOES



# 14 May 2013, GOES & Fermi



7/29/2016

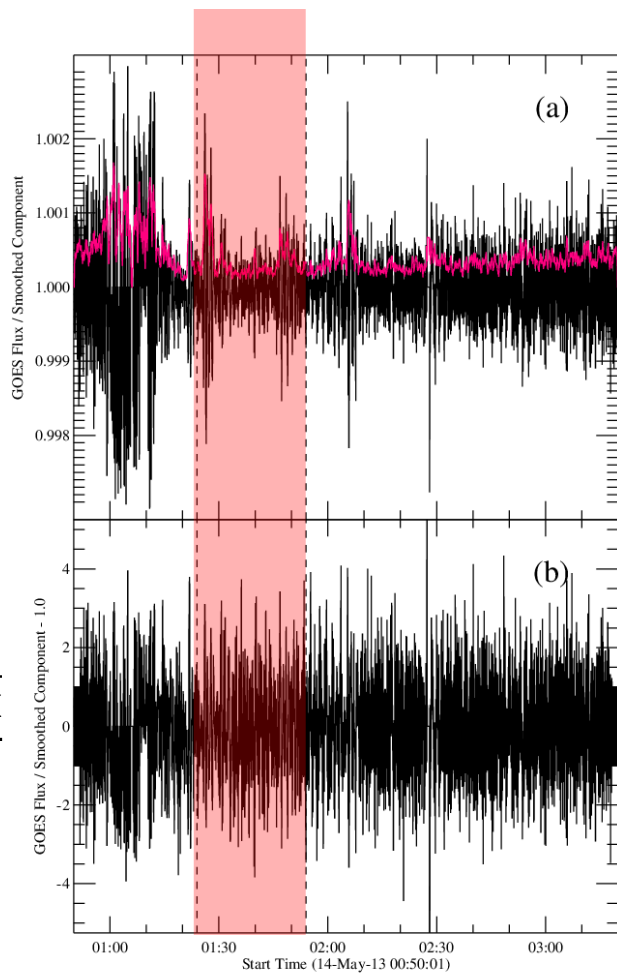
14 May 2013 Quasi-Periodic Pulsations

28-Oct-2015 17:25  
4

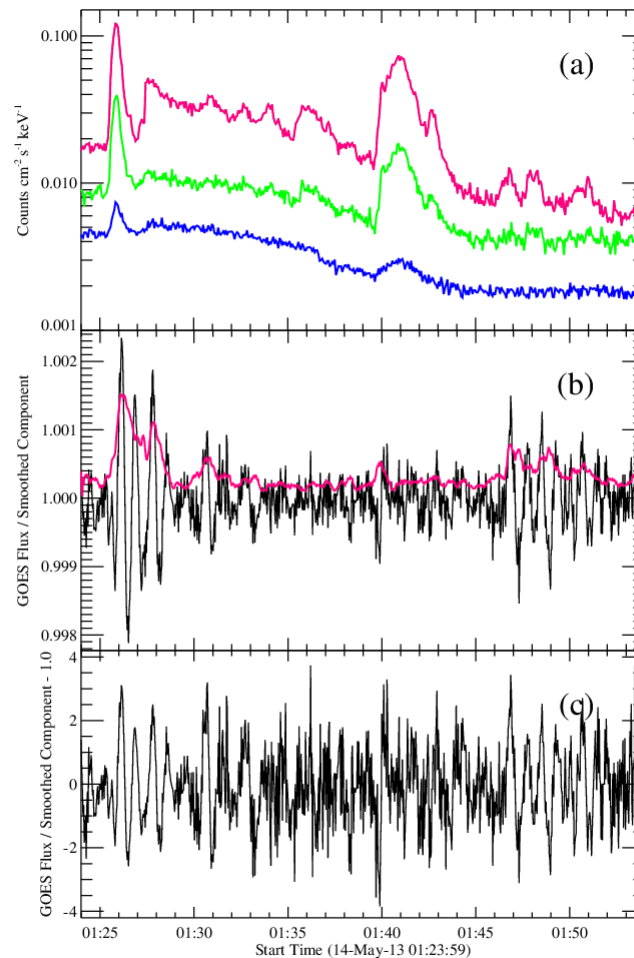
# 14 May 2013 Flux Normalization

GOES  
1 – 8 Å

$$\frac{\text{flux}}{\overline{\text{flux}}}$$



$$\frac{\text{flux} - \overline{\text{flux}}}{\overline{\text{flux}}}$$



RHESSI

6 – 12 keV

12 – 25 keV

25 – 50 keV

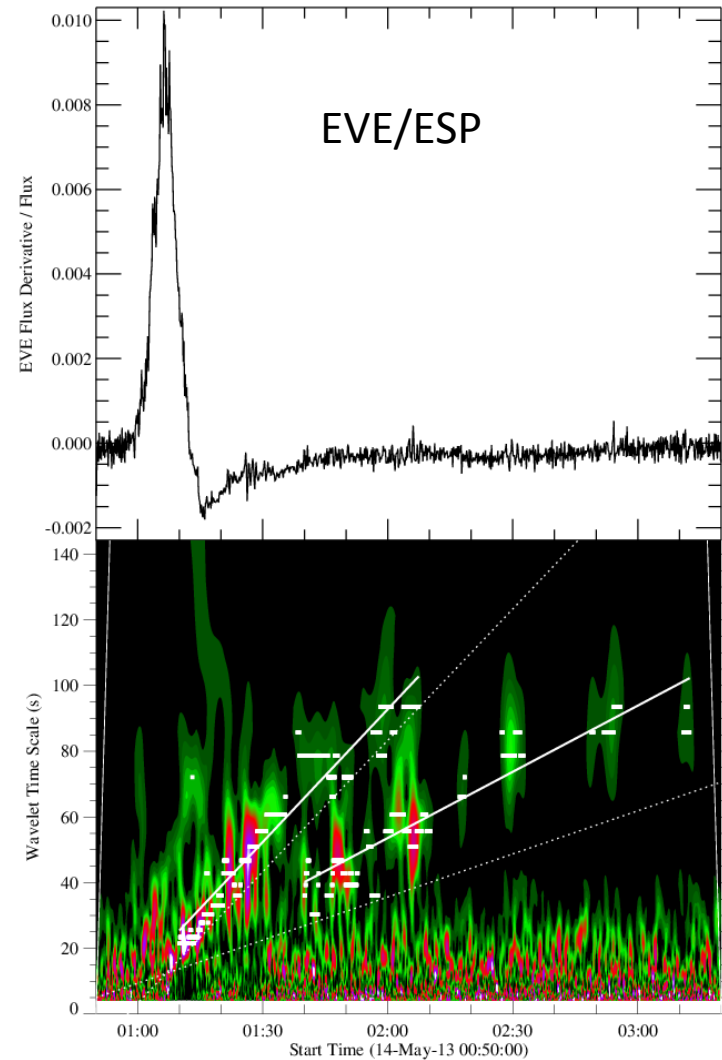
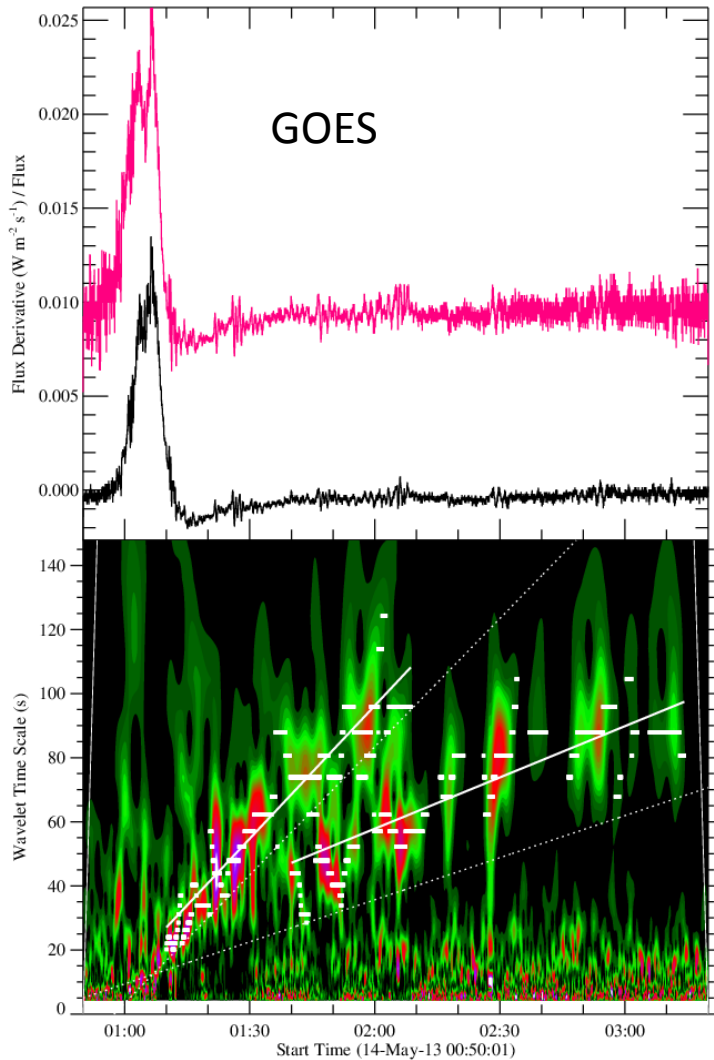
GOES

1 – 8 Å

$$\frac{\text{flux}}{\overline{\text{flux}}}$$

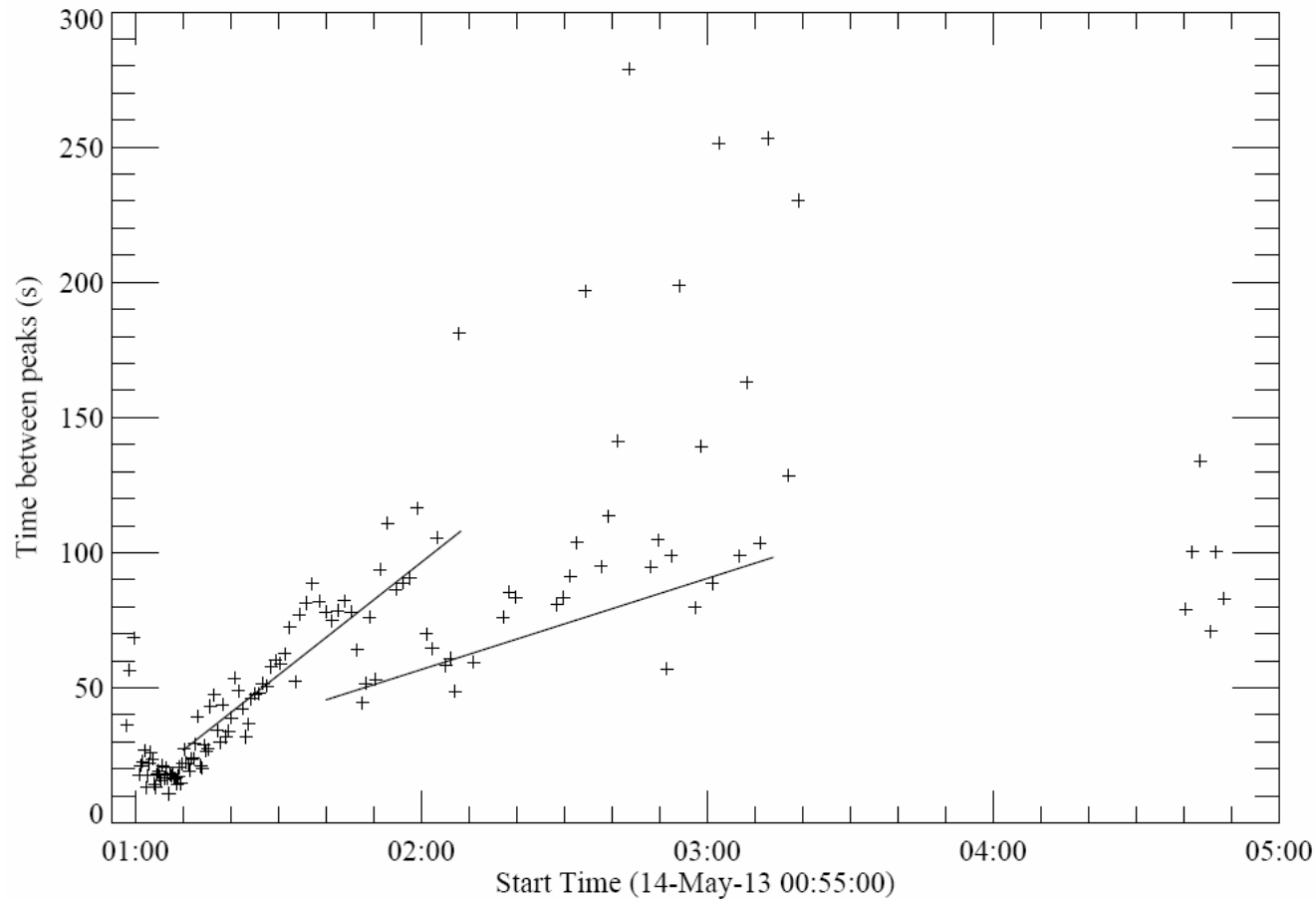
$$\frac{\text{flux} - \overline{\text{flux}}}{\overline{\text{flux}}}$$

# 14 May 2013 – Wavelet Analysis



# 14 May 2013

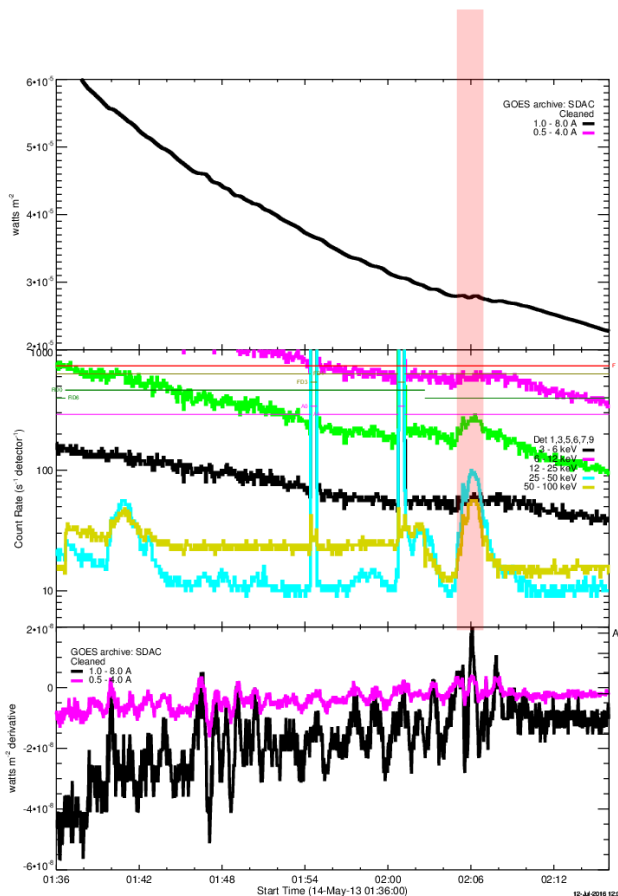
## Time between peaks



# 14 May 2013

## AIA & RHESSI Images

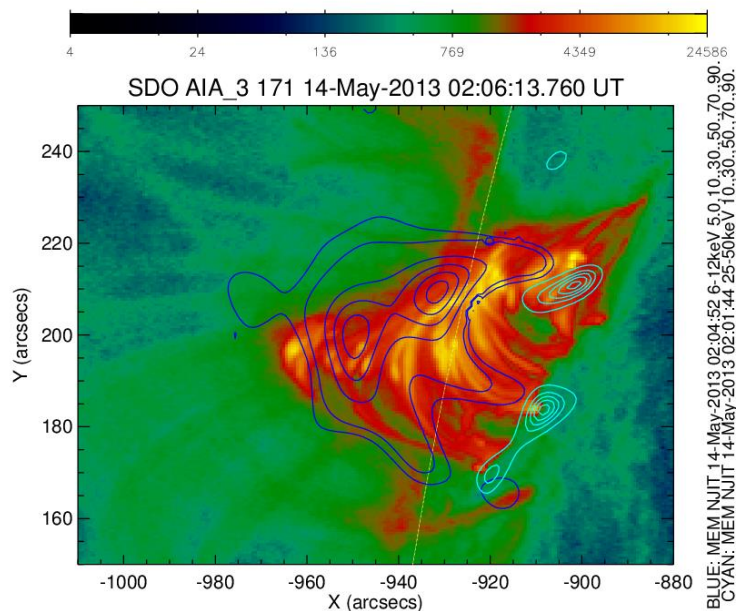
GOES Flux



RHESSI Flux

GOES Derivative

AR 1748



Color: AIA 171 Å  
Contours: RHESSI

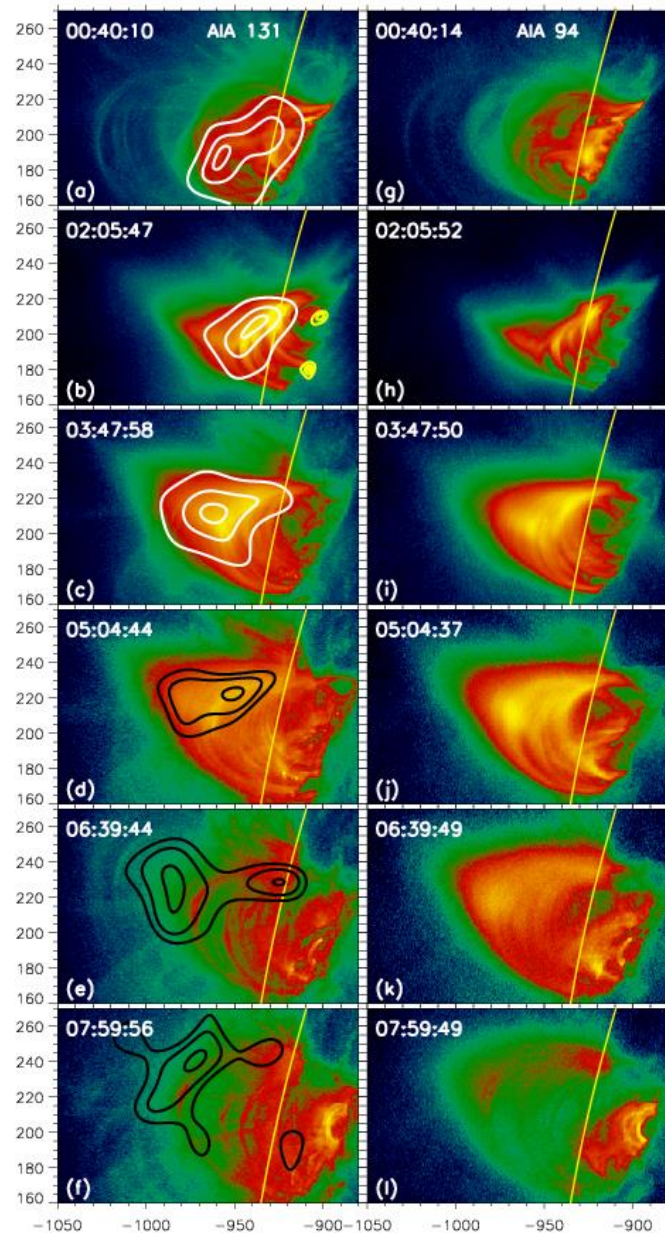
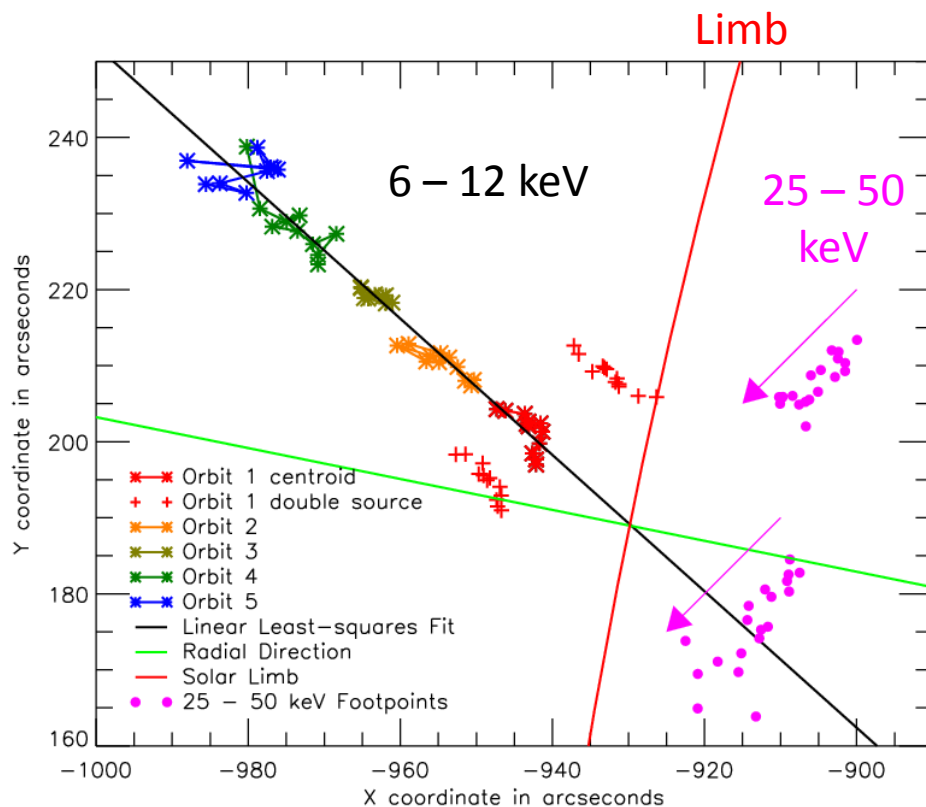
Blue 6 – 12 keV  
Cyan 25 – 50 keV

17-May-2016 17:50



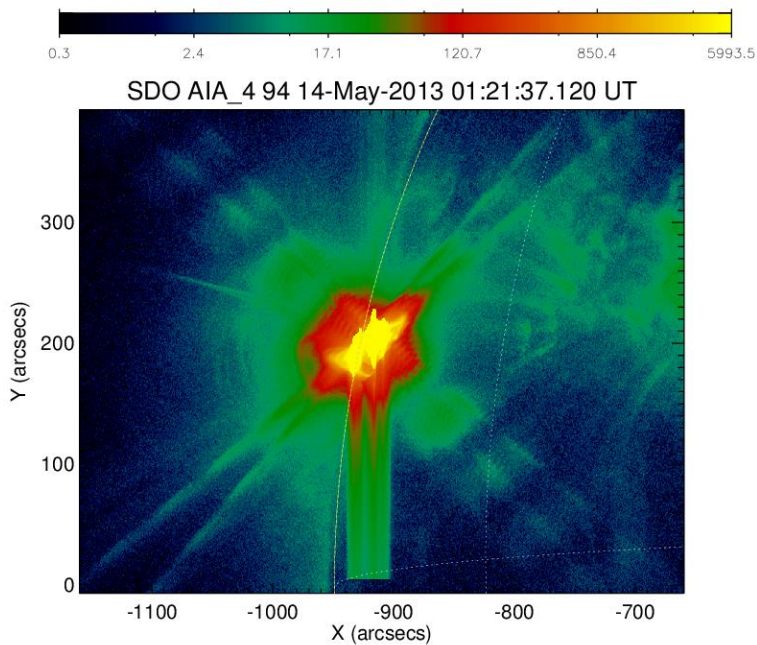
# 14 May 2013

## RHESSI Imaging

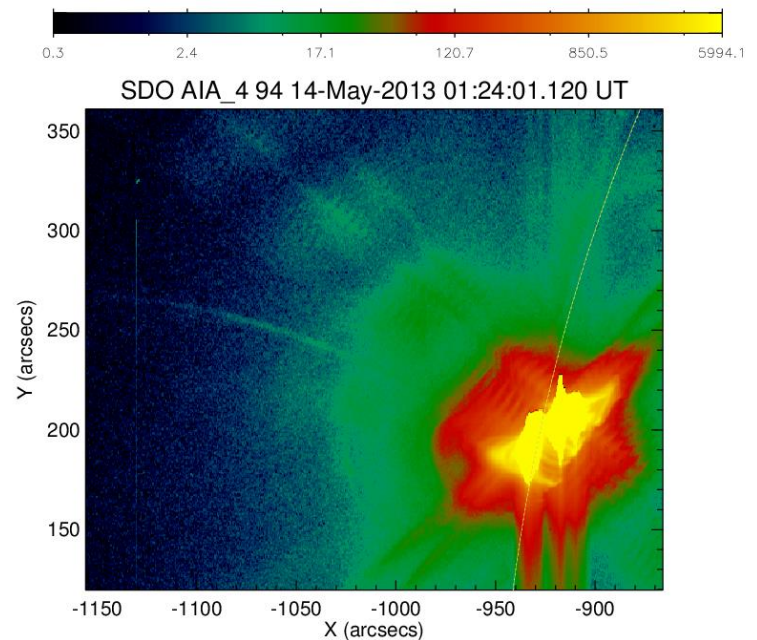


# 14 May 2013

## AIA 94 Å Current Sheet



17-May-2016 18:41

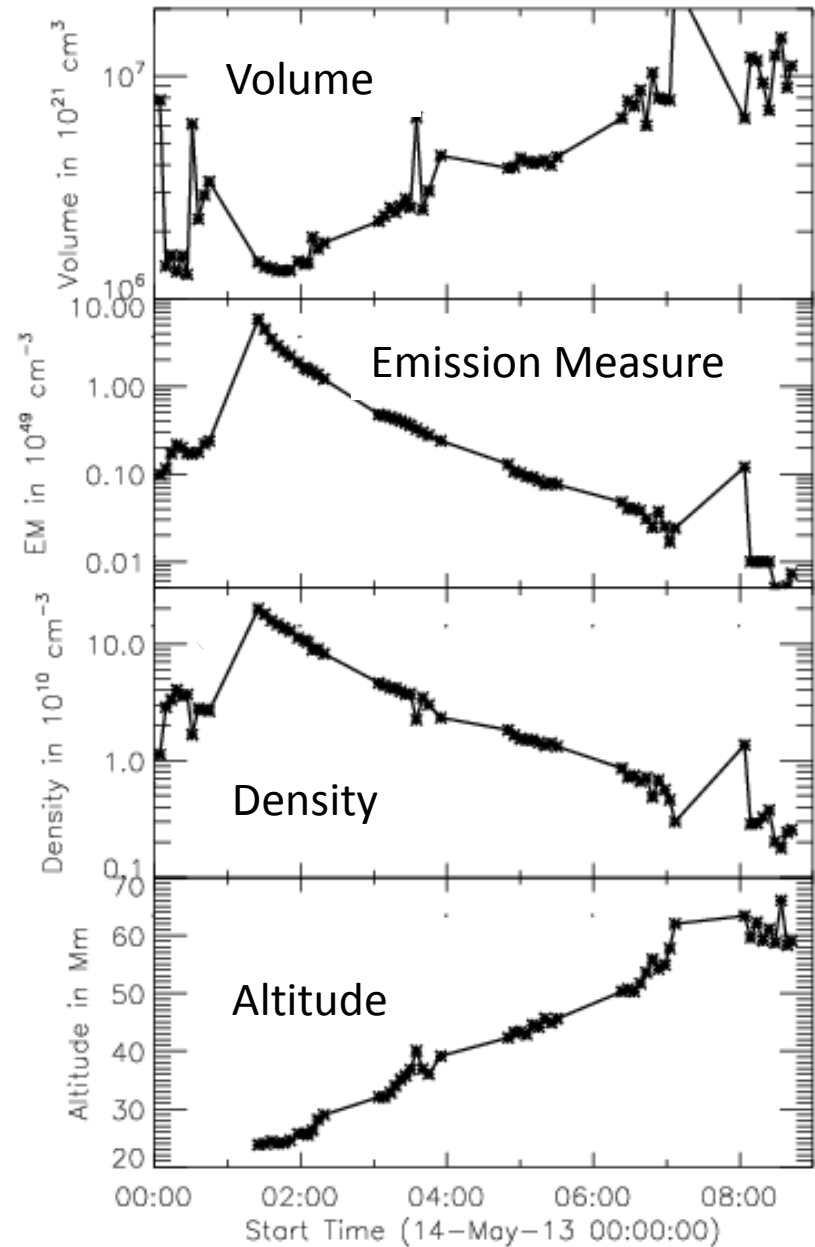


18-May-2016 09:06

14 May 2013

RHESSI

3 – 6 and 6 – 12 keV



# Coronal Magnetic Field Estimate

Assume QPP are consistent with **vertical kink mode oscillations** in the newly formed magnetic loops resulting from reconnection in the current sheet stretching up to the base of the CME.

Kink-mode period,  $P_k = 2\sqrt{2} L/v_A$

$L = \pi h/2$  cm - loop half length with height,  $h$ , obtained from RHESSI images,

$v_A = DB/\sqrt{n}$  cm s<sup>-1</sup> - Alfvén speed,

$D = 1.93 \times 10^{11}$ ,  $B$  is the magnetic field in gauss.

$n = \sqrt{\frac{EM}{fV}}$  - plasma density in loop, assumed large compared to outside loop

$EM$  is emission measure from GOES fluxes,  $f$  - filling factor assumed = 1.

$V = \pi \sigma_x \sigma_y (\sigma_x + \sigma_y)/2$  is the source volume estimated from RHESSI 6 – 12 keV images.

$\sigma_{x,y}$  - second moments of RHESSI source.

Let  $Pk(t) = \tau(t)$  – time scale at peak wavelet power at time,  $t < 03:20$  UT.

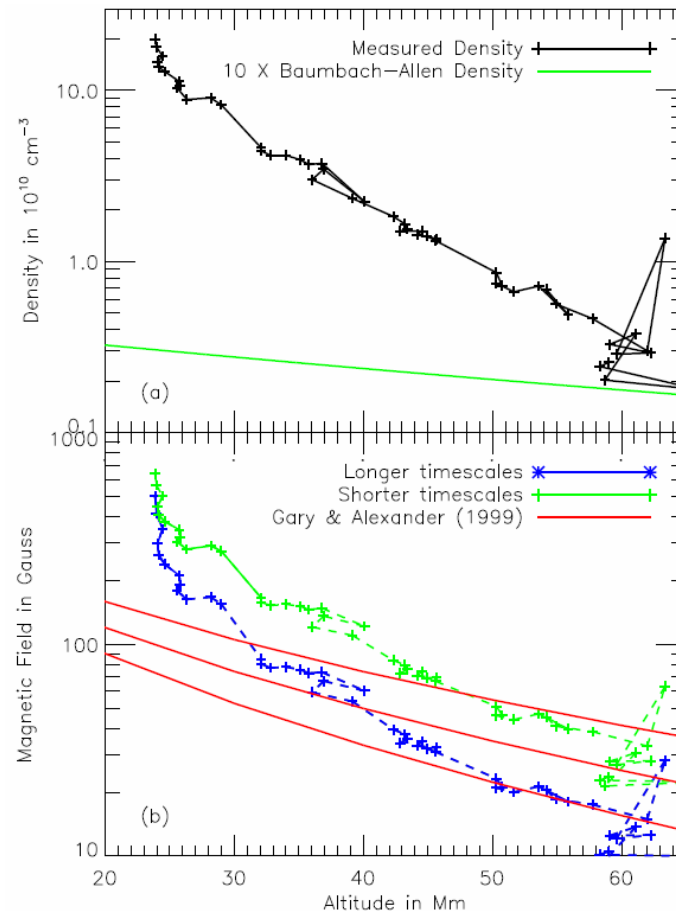
Linearly extrapolate the trend lines for  $t > 03:30$  UT.

Then,

$$B = \pi\sqrt{2n} h / (D \tau)$$

# 14 May 2013

## Density & Magnetic Field vs. Altitude



# Conclusions

- GOES X3.2 flare on 14 May 2013
- QPP seen for >2 hours – 163 peaks.
- Time scale increases from ~25 to ~100 s in ~1 hour.
- Consistent with vertical kink mode oscillations with

$$P_k = 2\sqrt{2} L/v_A$$

B = 500 G at 24 Mm

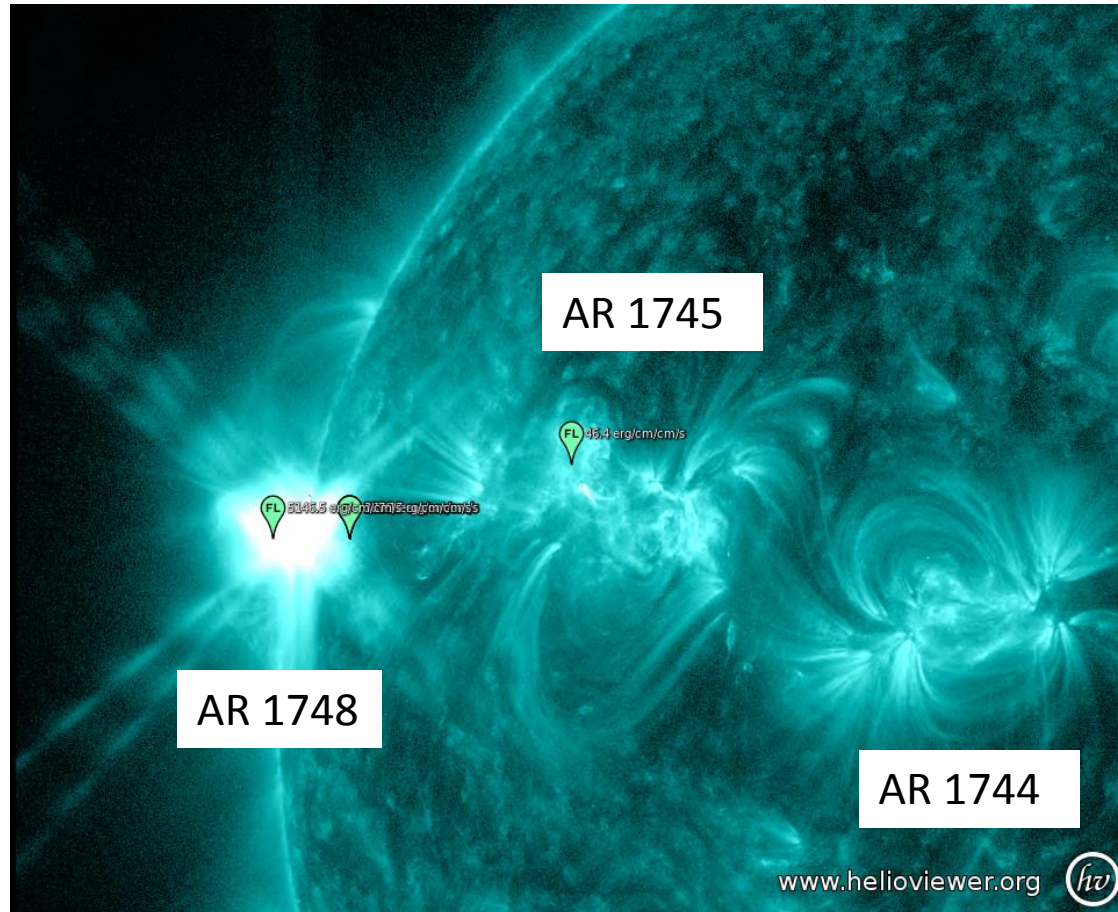


10 - 40 G at 60 Mm

- Density =  $2 \times 10^{11} \text{ cm}^{-3}$  at 24 Mm  
 $2 \times 10^{10} \text{ cm}^{-3}$  at 60 Mm

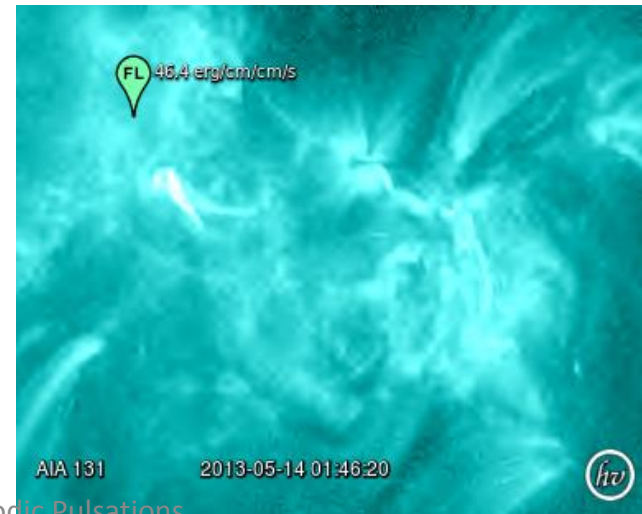
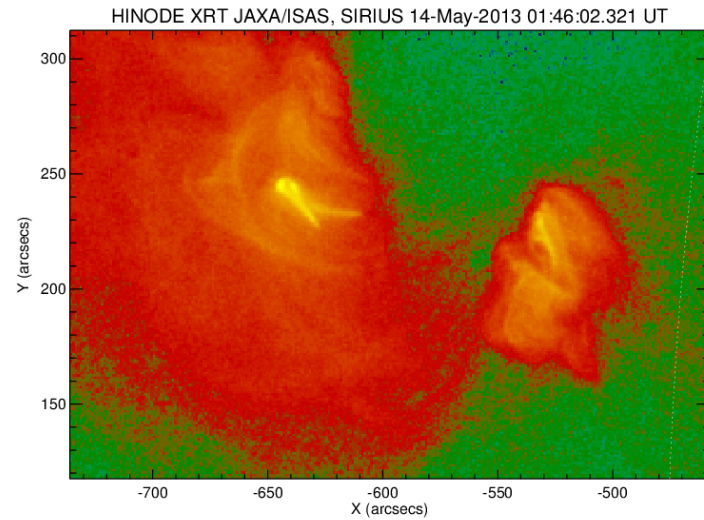
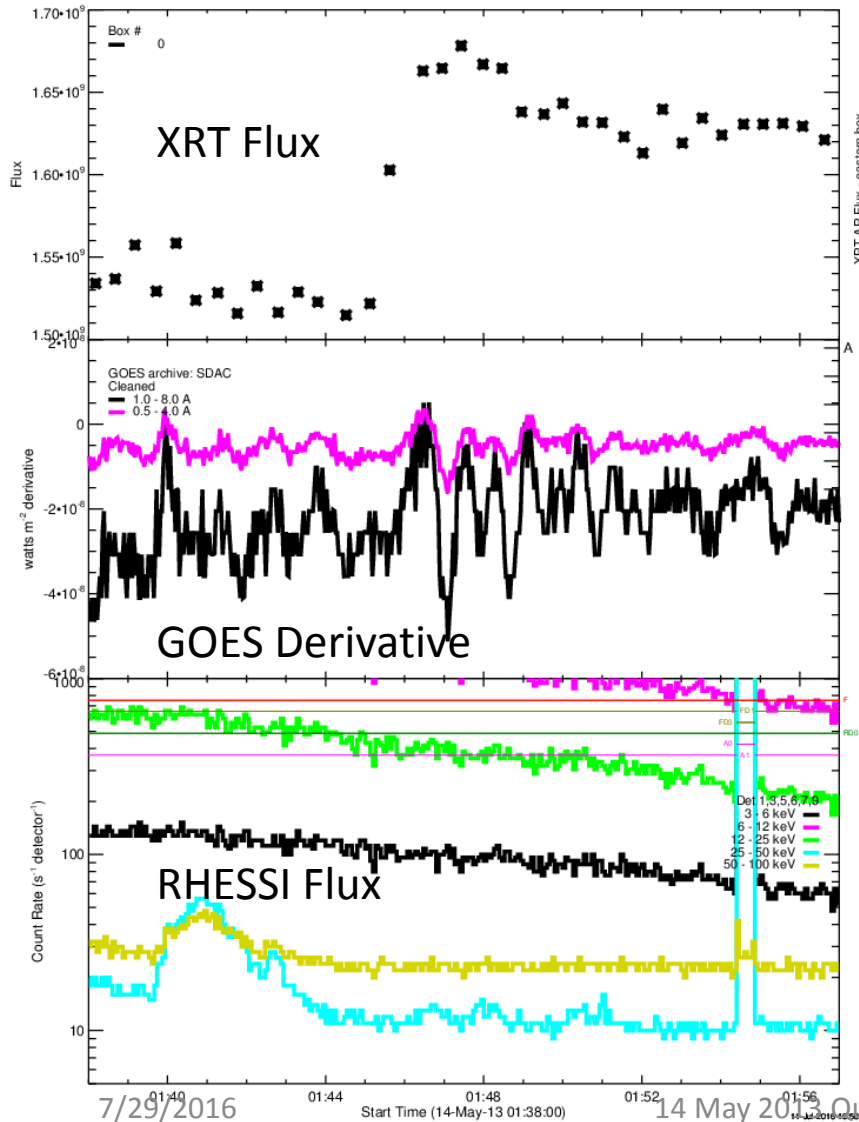


14 May 2013 01:45 UT  
AIA 131 Å



# 14 May 2013

## XRT & AIA 131Å Flare from different location AR 1745

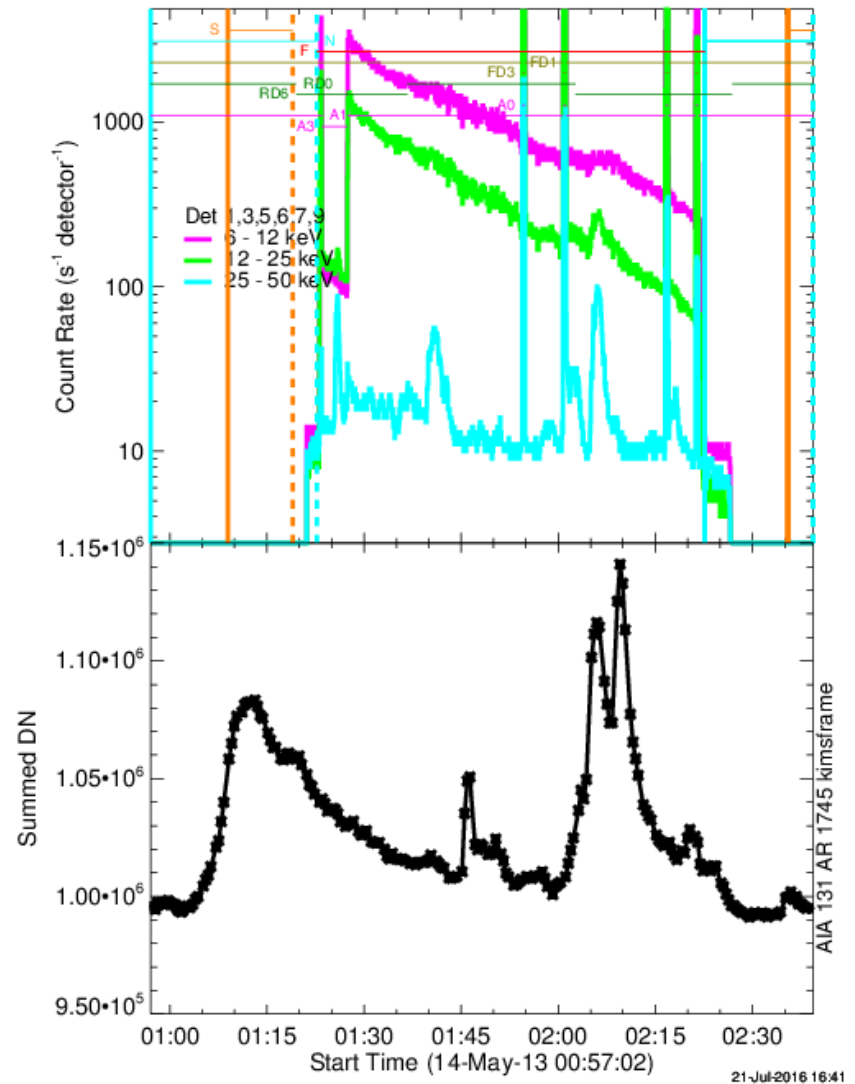




14 May 2013

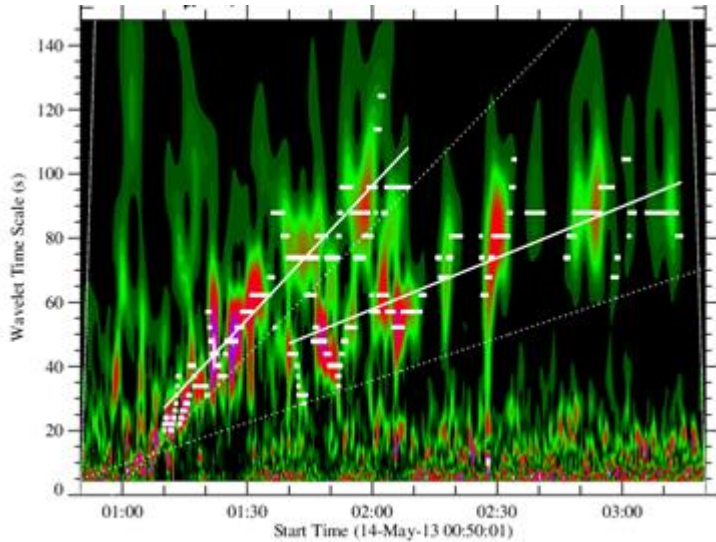
RHESSI  
6 – 50 keV

AIA 131 Å  
Flux from  
AR1745

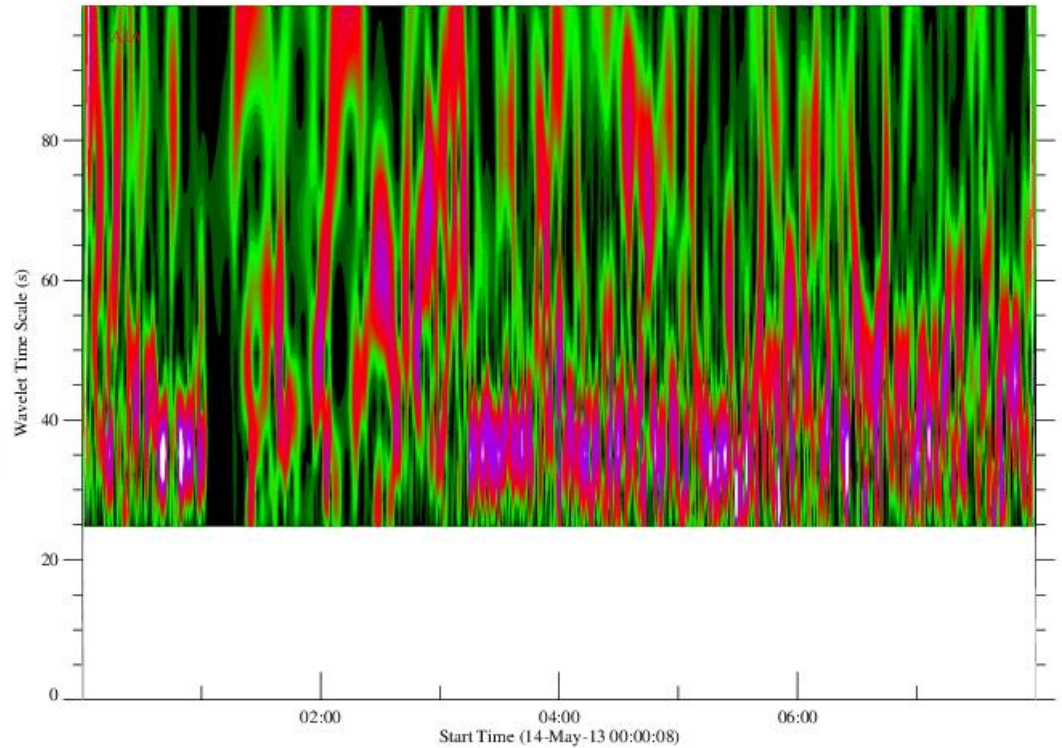


# 14 May 2013 Wavelet Power

GOES



AIA 131 Å, AR1458 fullframe



# New Conclusions

- GOES X3.2 flare on 14 May 2013
- QPP seen for >21 hours – fewer than 163 peaks.
- Time scale increases from ~25 to ~100 s in ~1 hour.
- Consistent with vertical kink mode oscillations with

$$P_k = 2\sqrt{2} L/v_A$$

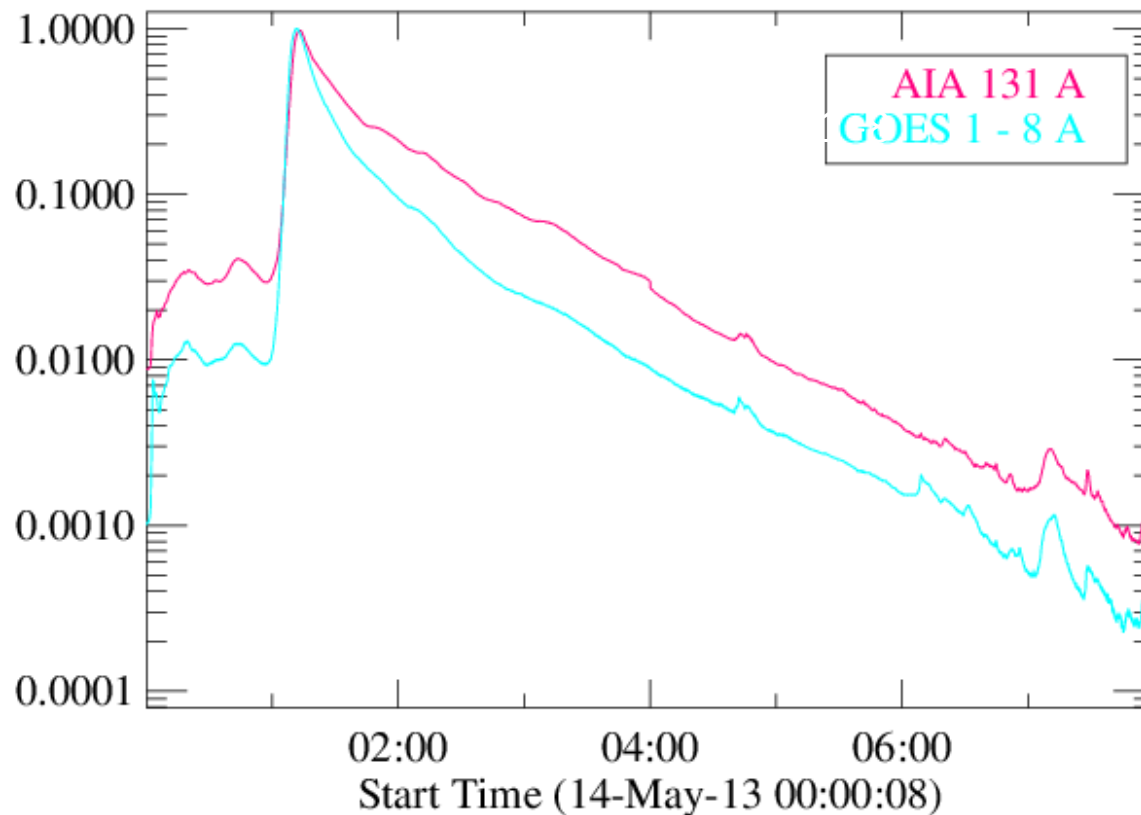
B = 500 G at 24 Mm



10 - 40 G at 60 Mm

- Density =  $2 \times 10^{11} \text{ cm}^{-3}$  at 24 Mm  
 $2 \times 10^{10} \text{ cm}^{-3}$  at 60 Mm

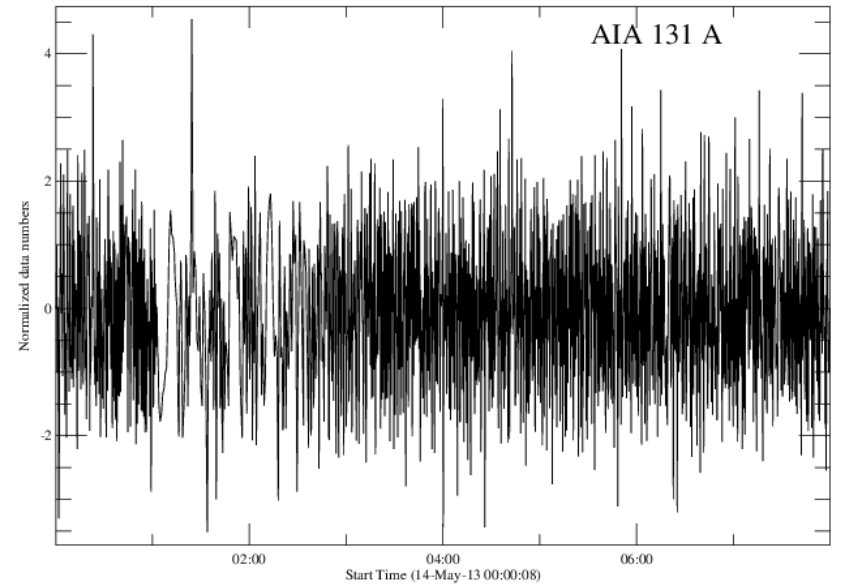
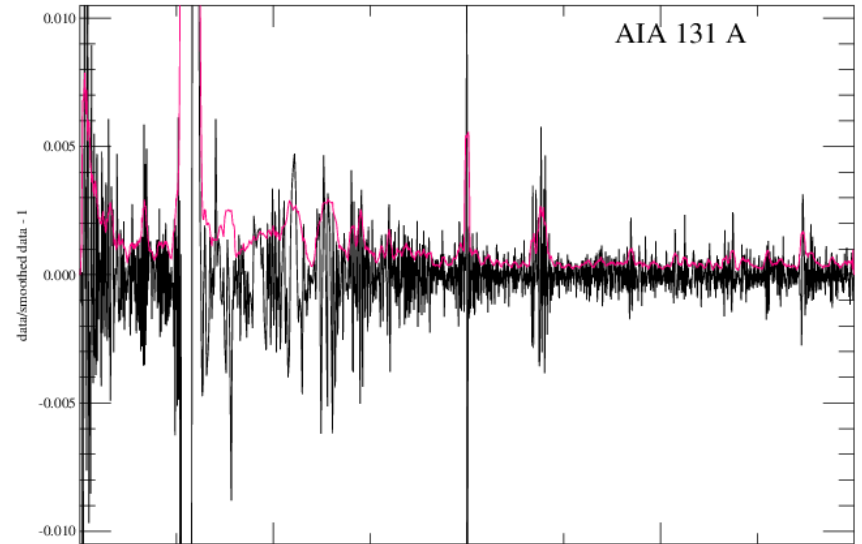
# 14 May 2013 AIA 131 Å AR1748 Fullframe Light Curve



14 May 2013

AIA 131 Å

AR 1458

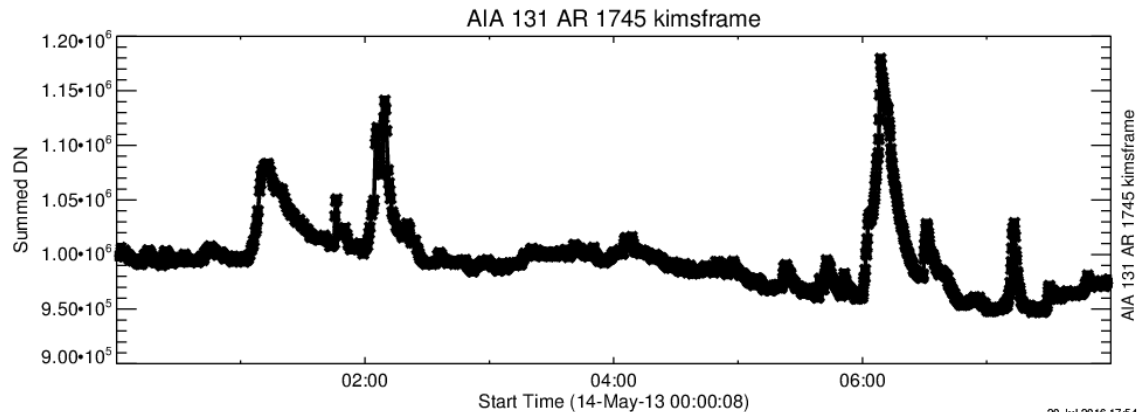
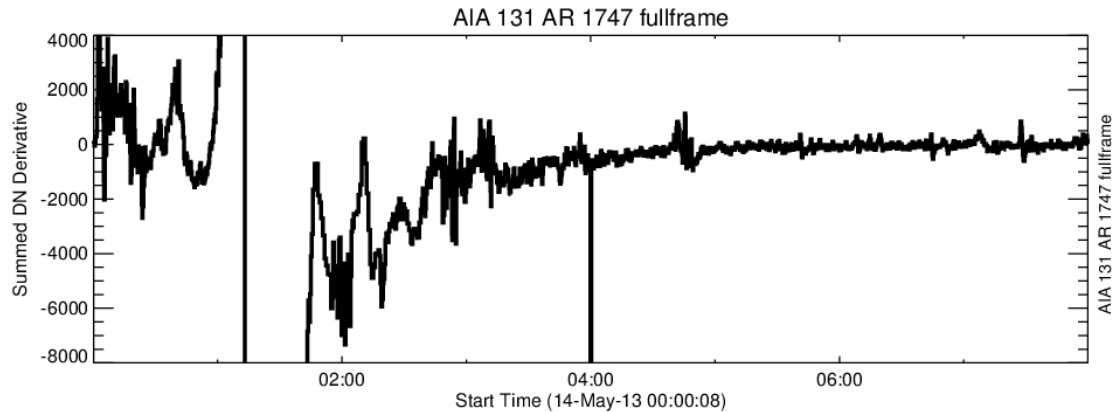


# New Conclusions

# 14 May 2013

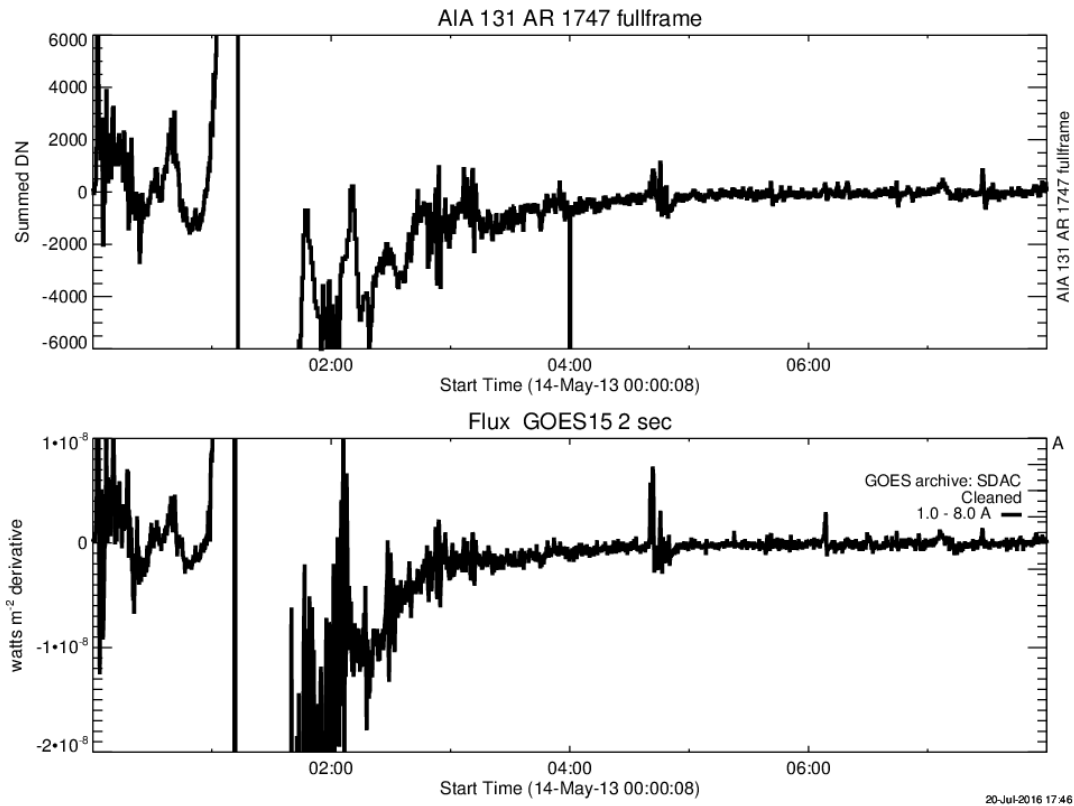
## AIA 131 Å lightcurves

### AR 1745 & 1747



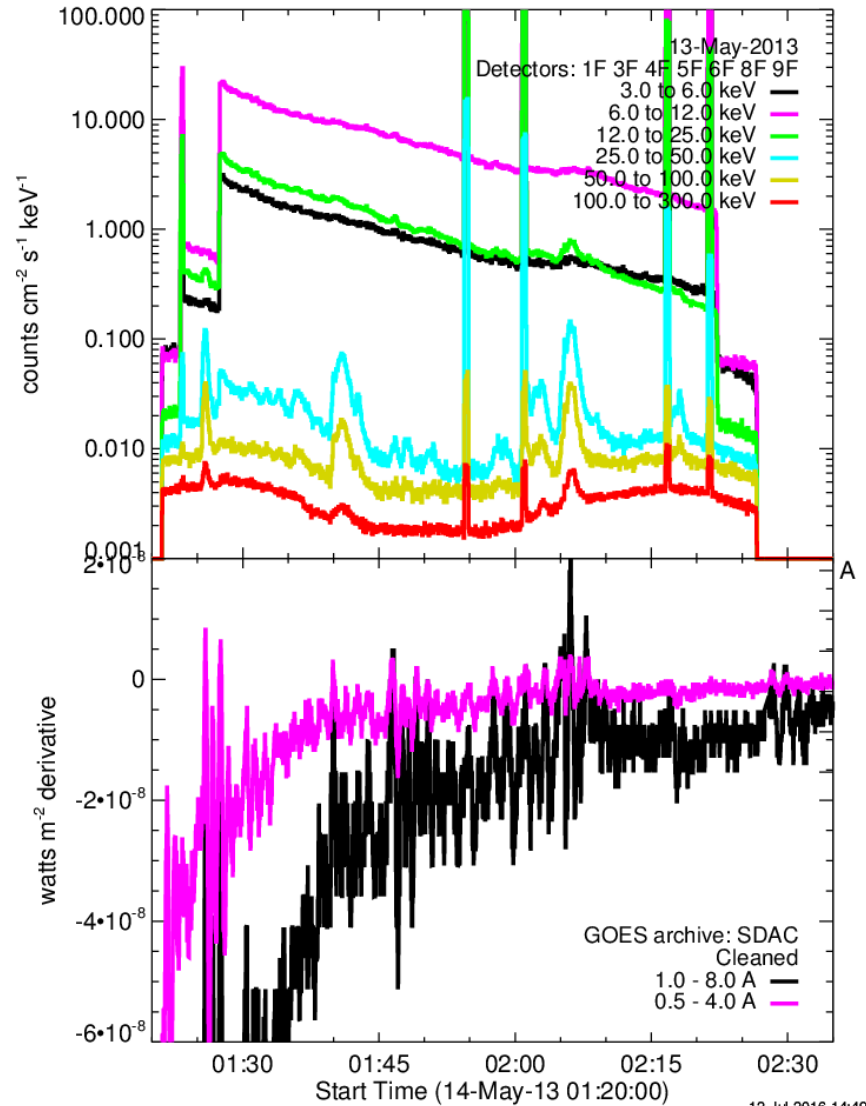
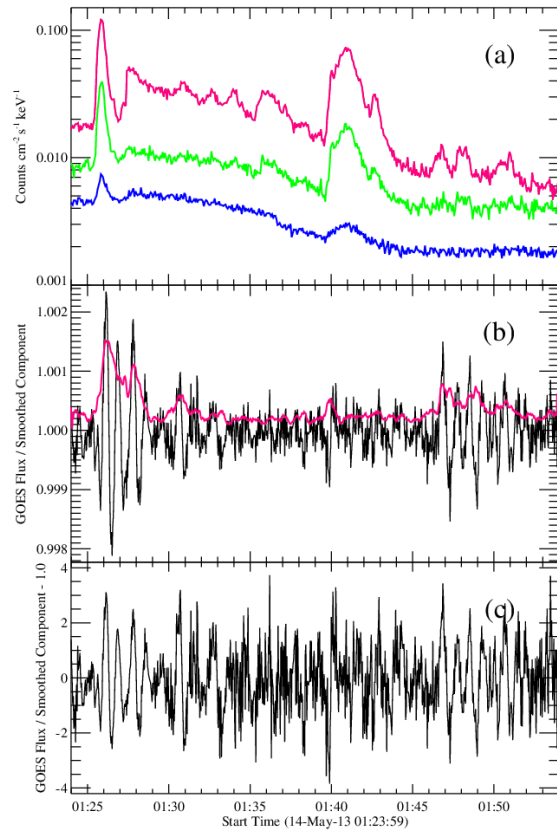
# 14 May 2013

## AIA 131 Å & GOES Time Derivatives



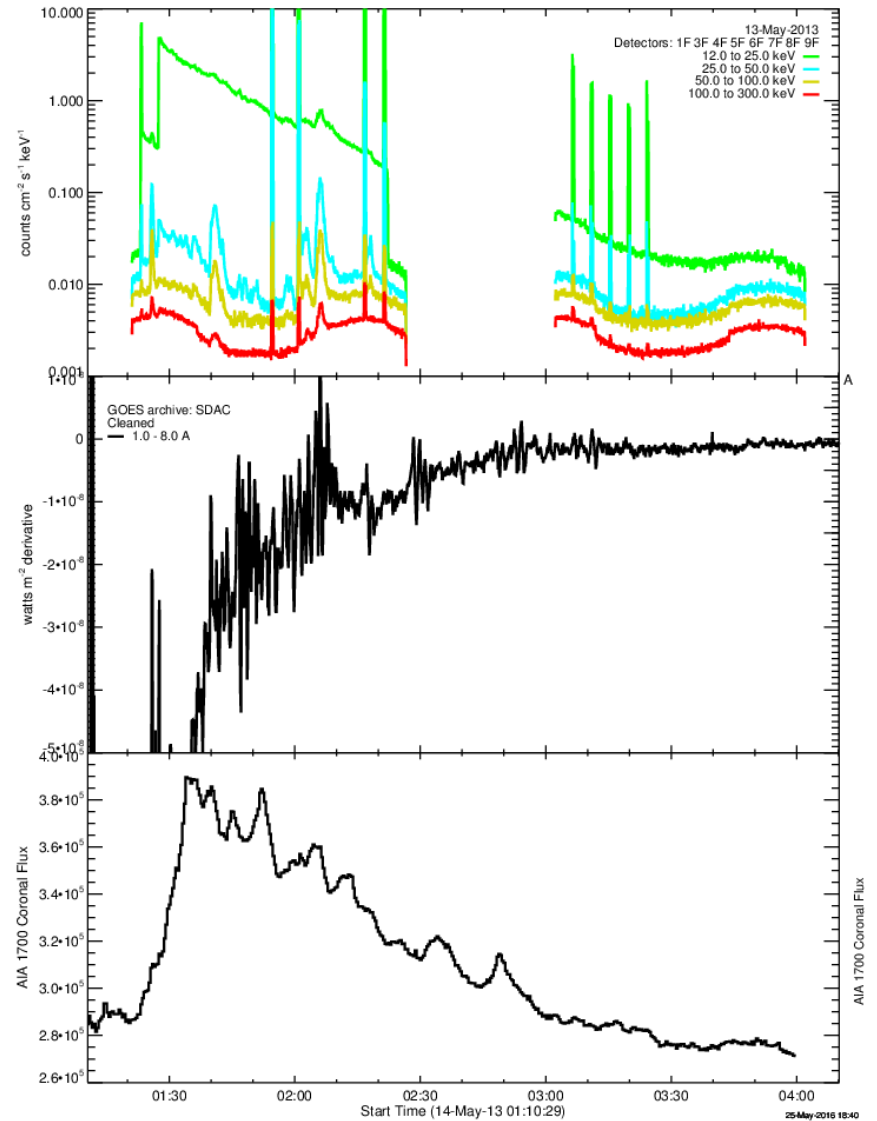
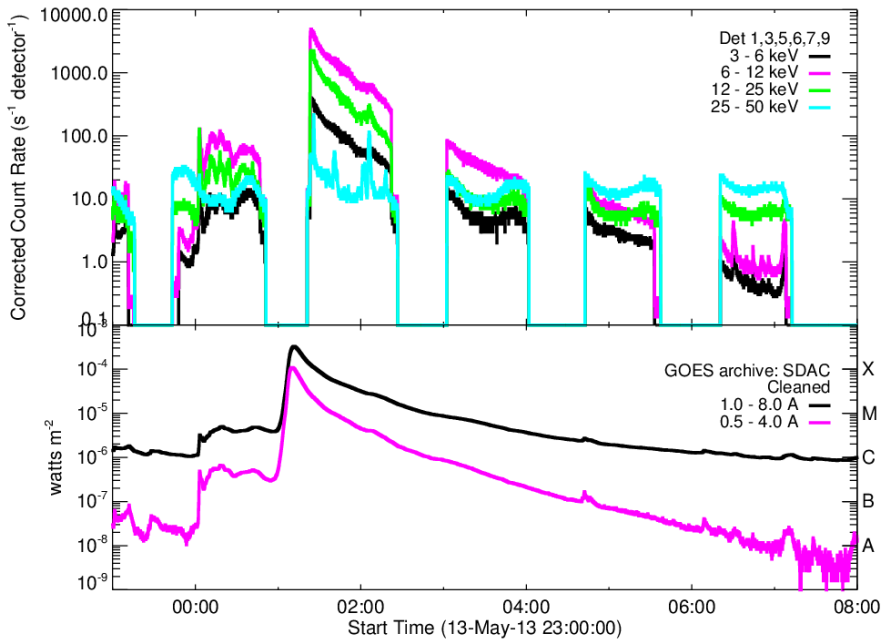


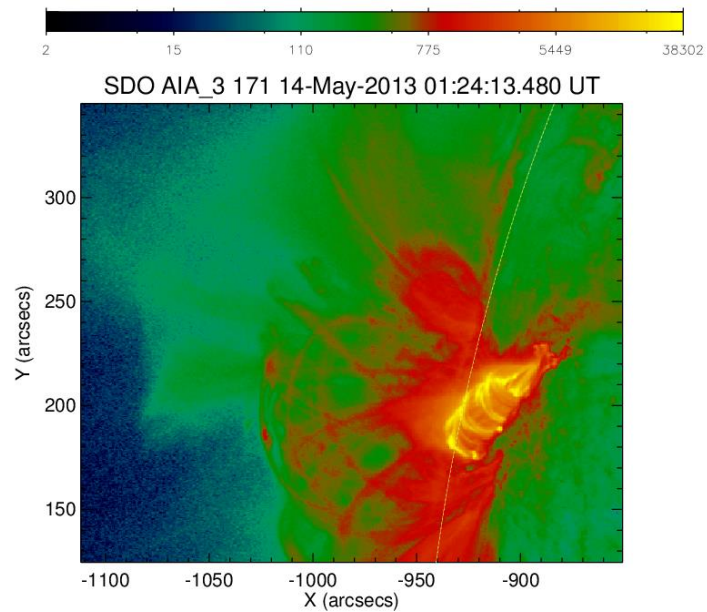
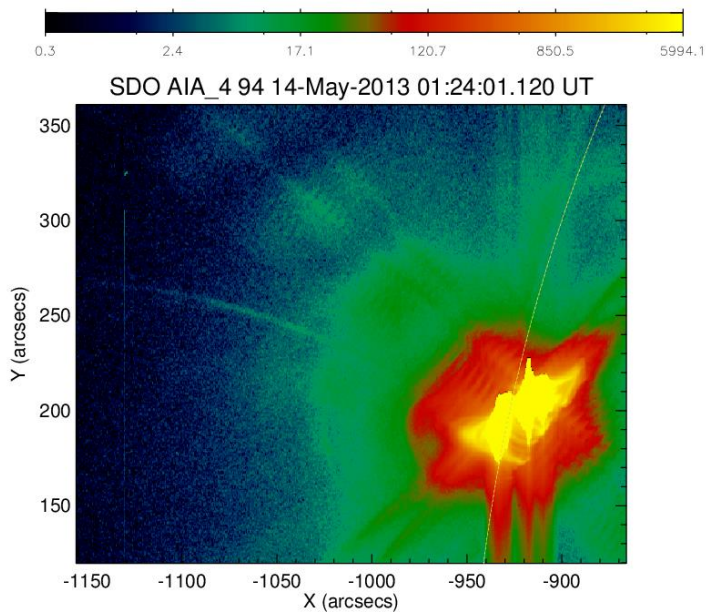
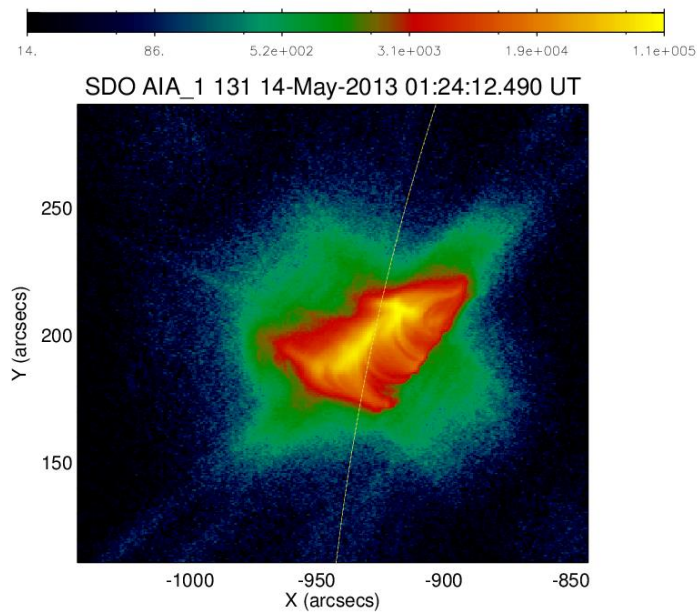
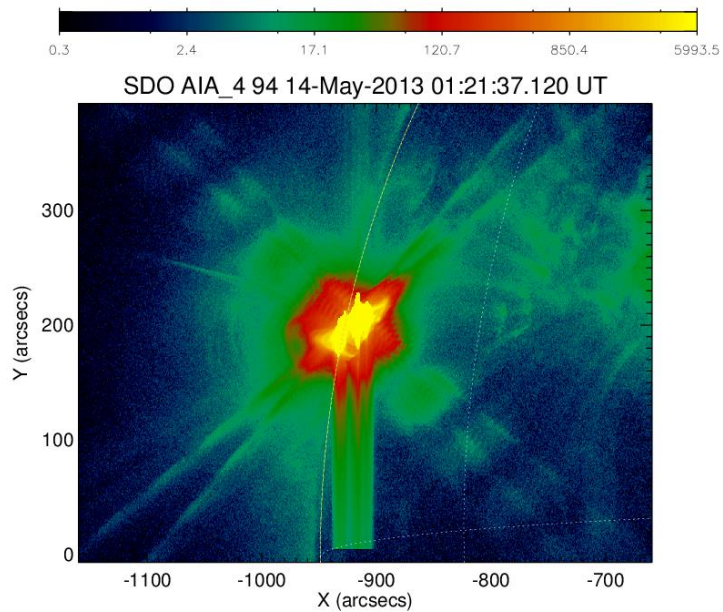
# 14 May 2013



# 14 May 2013

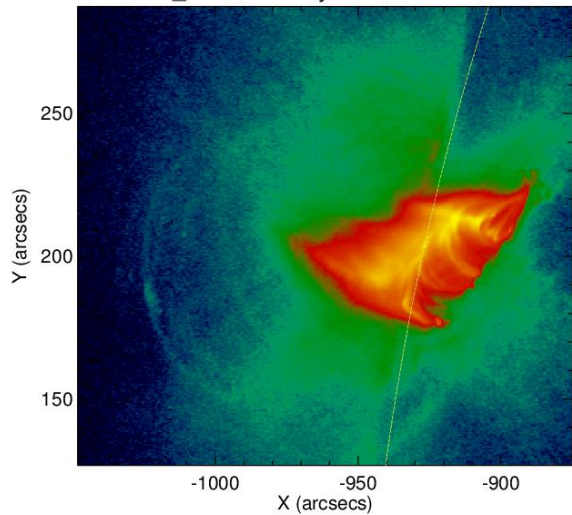
## AIA 1700 Å







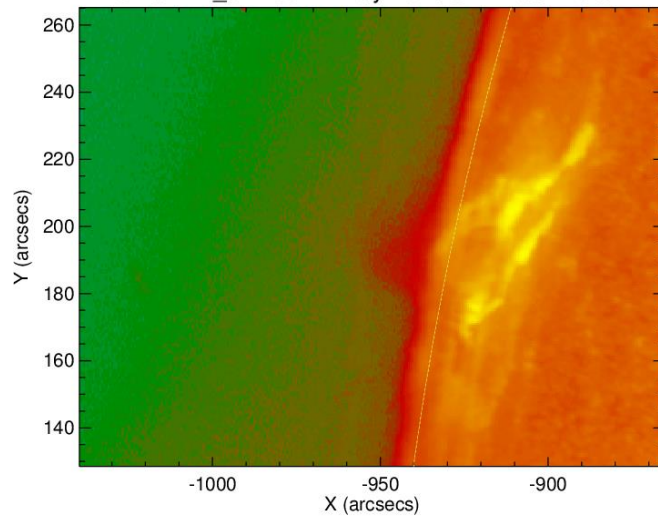
SDO AIA\_2 193 14-May-2013 01:24:08.770 UT



18-May-2016 09:15



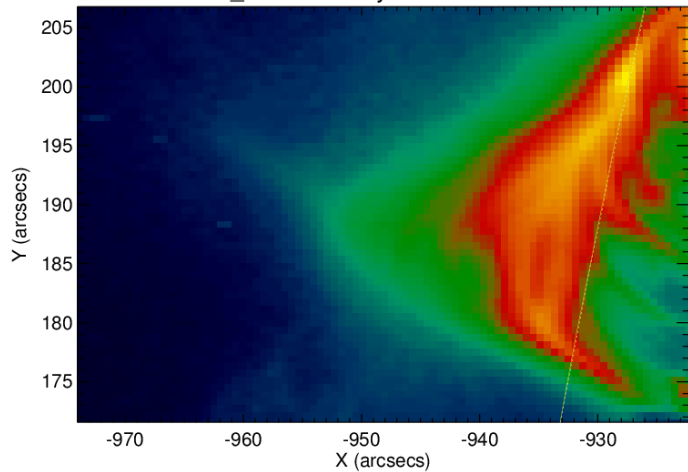
SDO AIA\_3 1700 14-May-2013 01:24:06.710 UT



18-May-2016 09:15



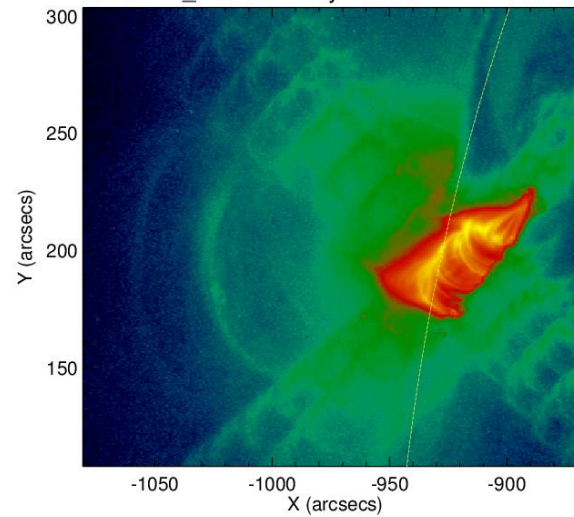
SDO AIA\_1 335 14-May-2013 01:24:02.620 UT



18-May-2016 09:10

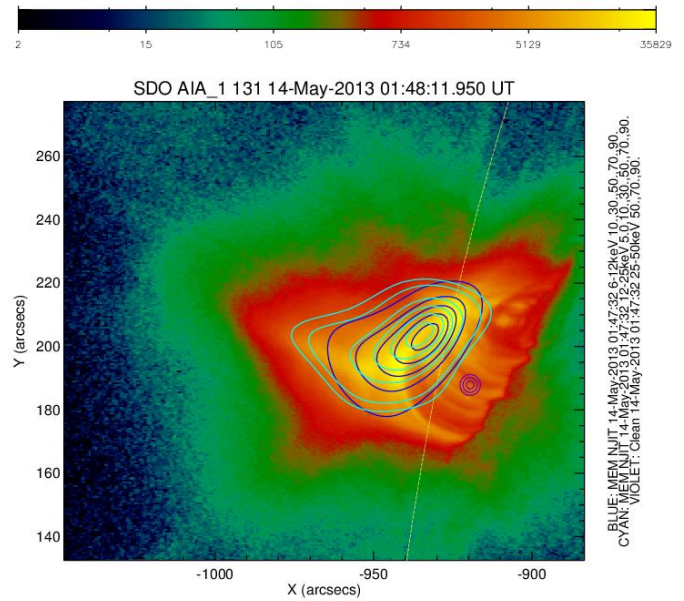
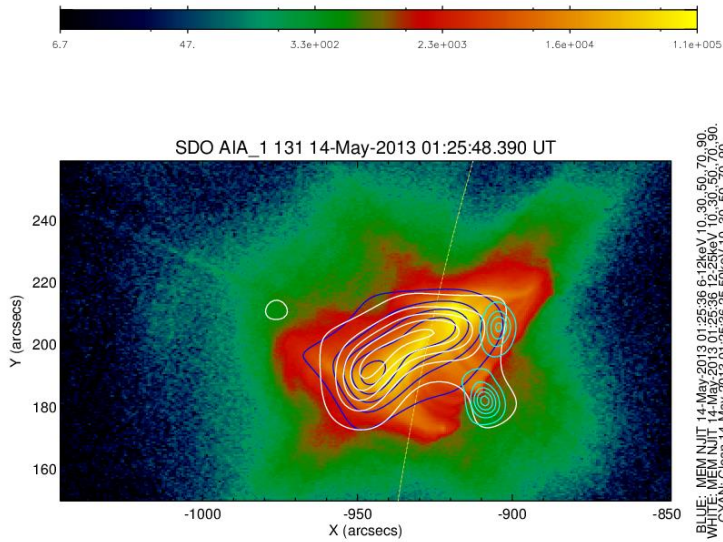


SDO AIA\_1 335 14-May-2013 01:24:14.630 UT



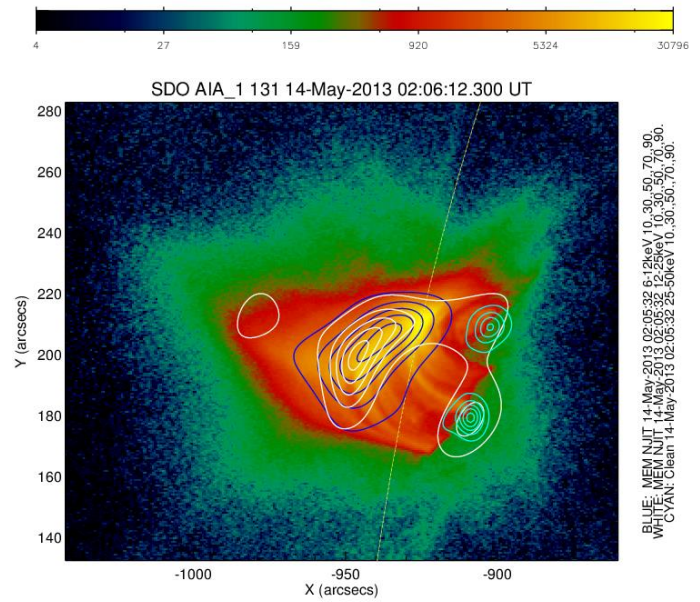
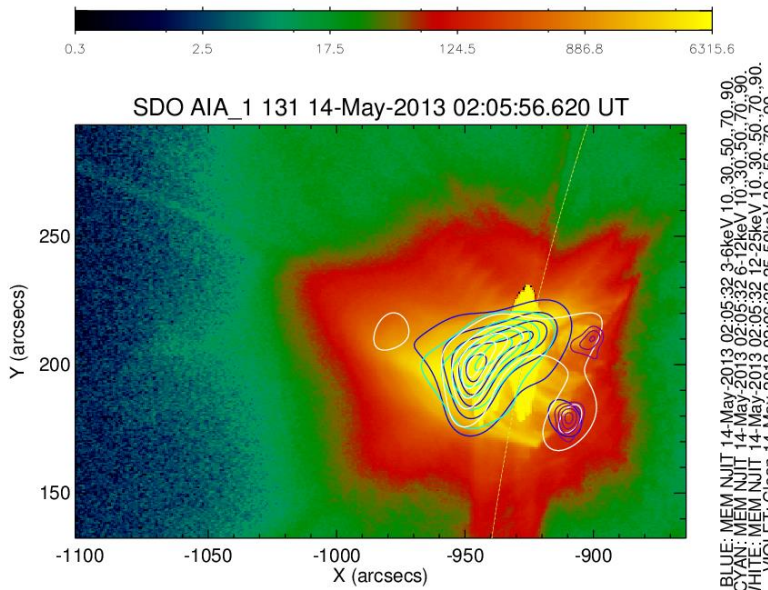
18-May-2016 09:24

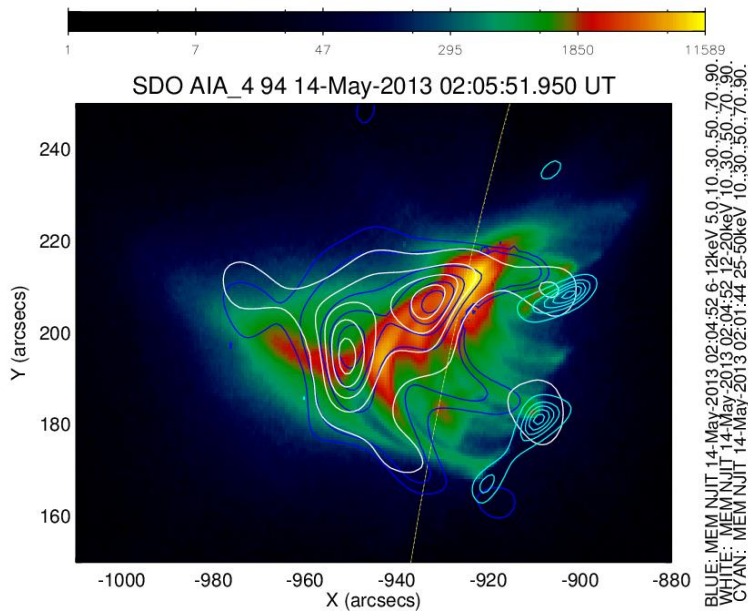




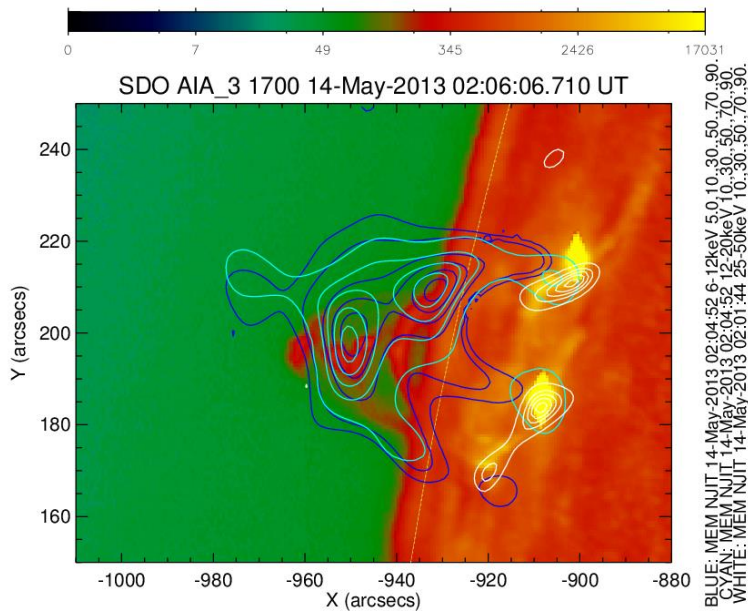
16-May-2016 17:36

16-May-2016 17:22

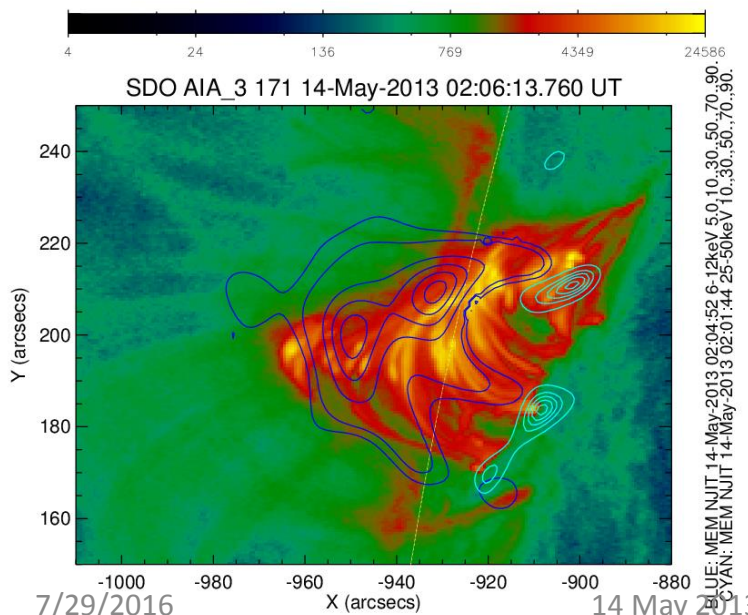




17-May-2016 18:25



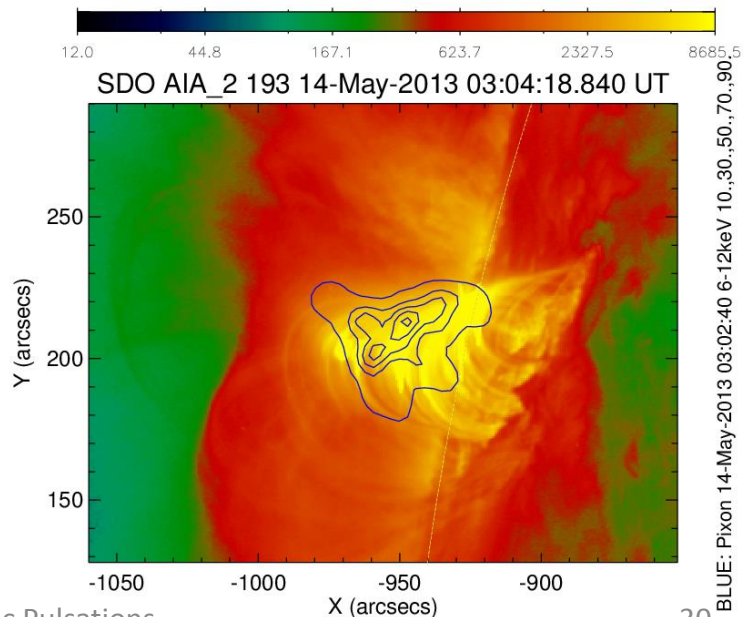
17-May-2016 18:09



7/29/2016

14 May 2013 Quasi-Periodic Pulsations

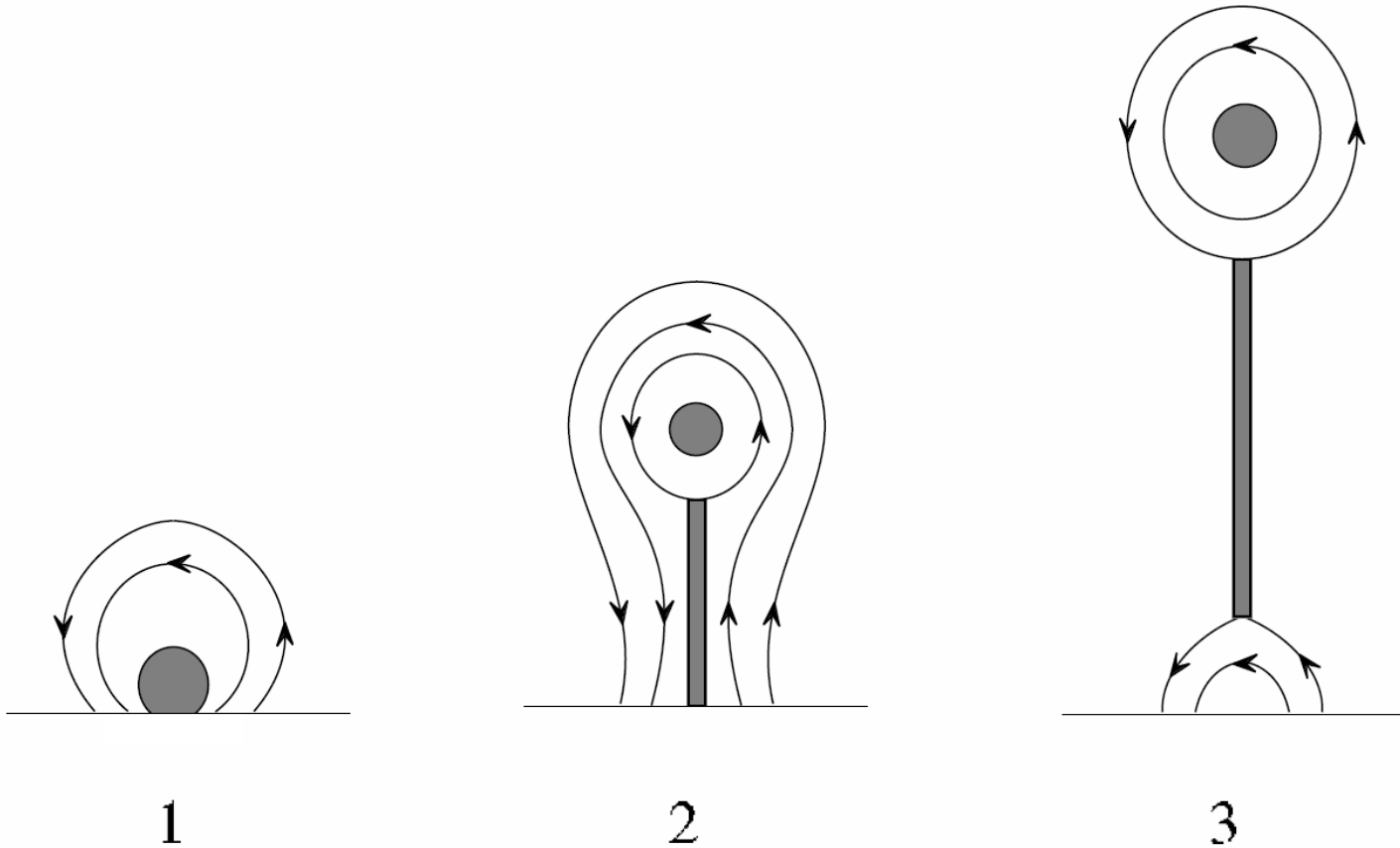
17-May-2016 17:50



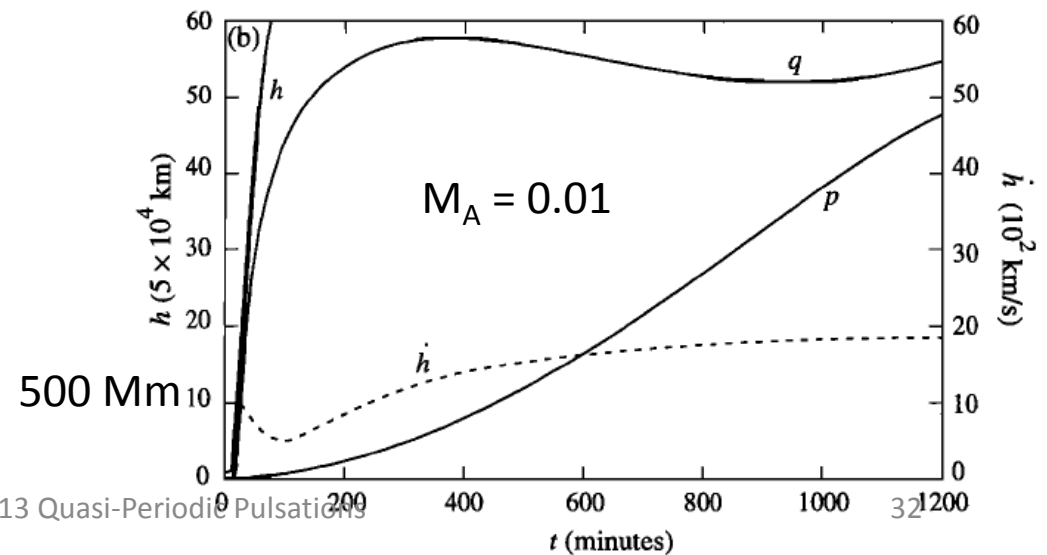
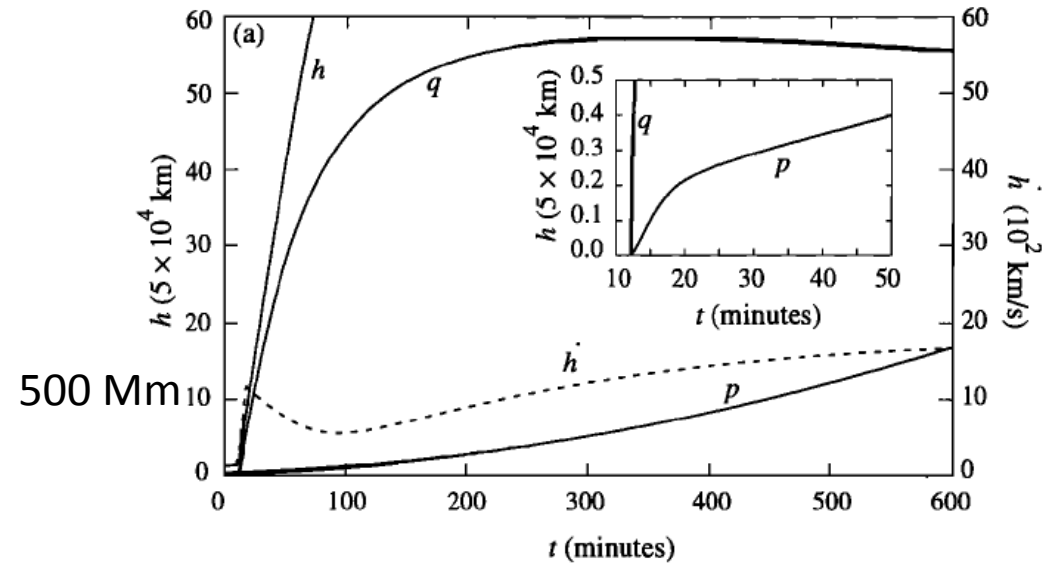
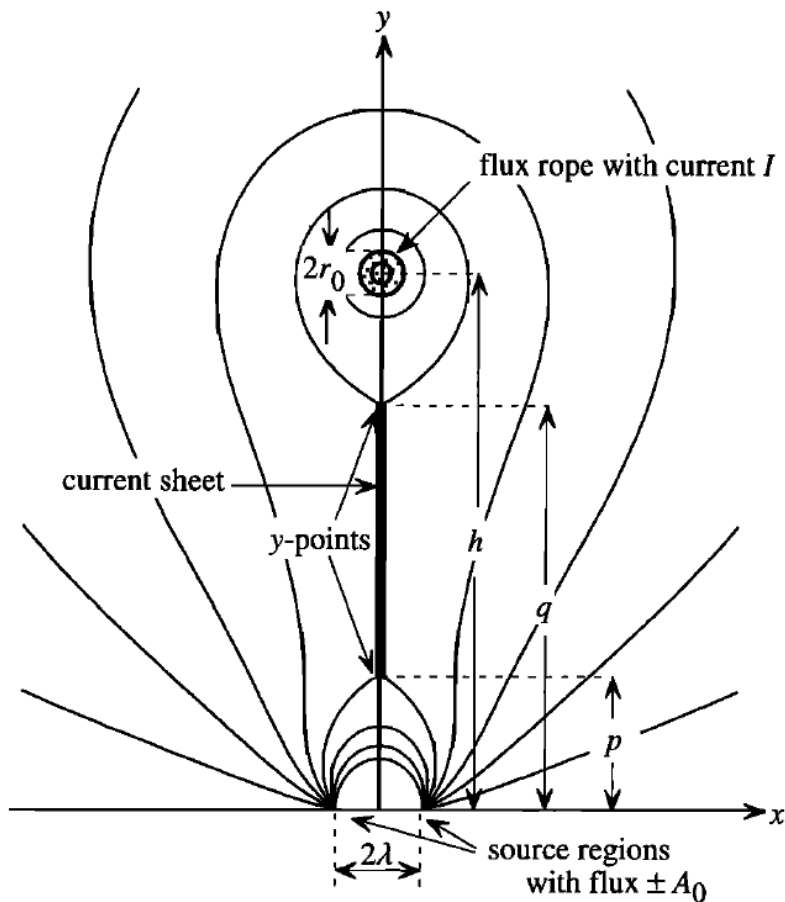
30

12-May-2016 17:09

# Priest & Forbes (2002)

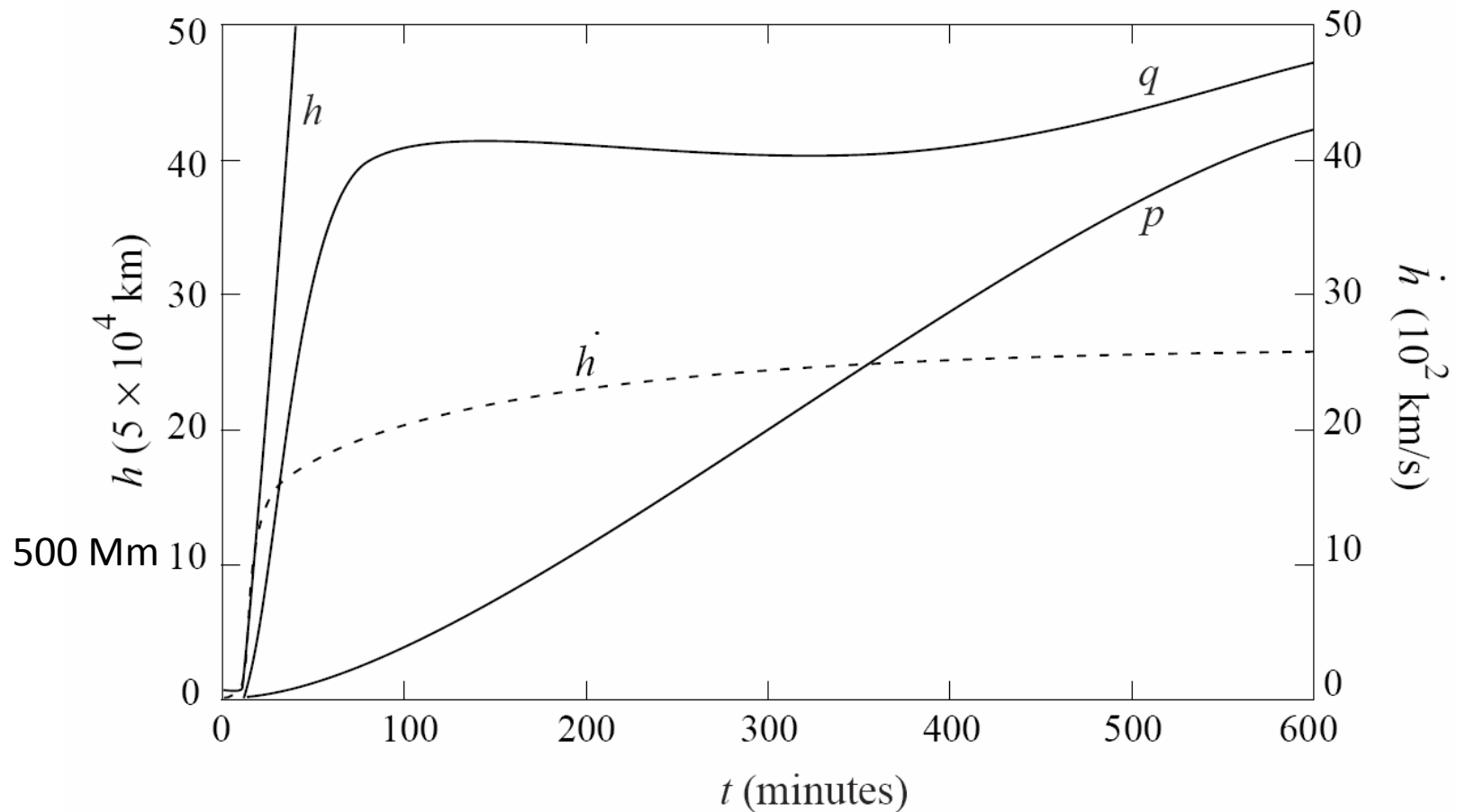


# Lin & Forbes, JGR, 105, 2375 (2000)





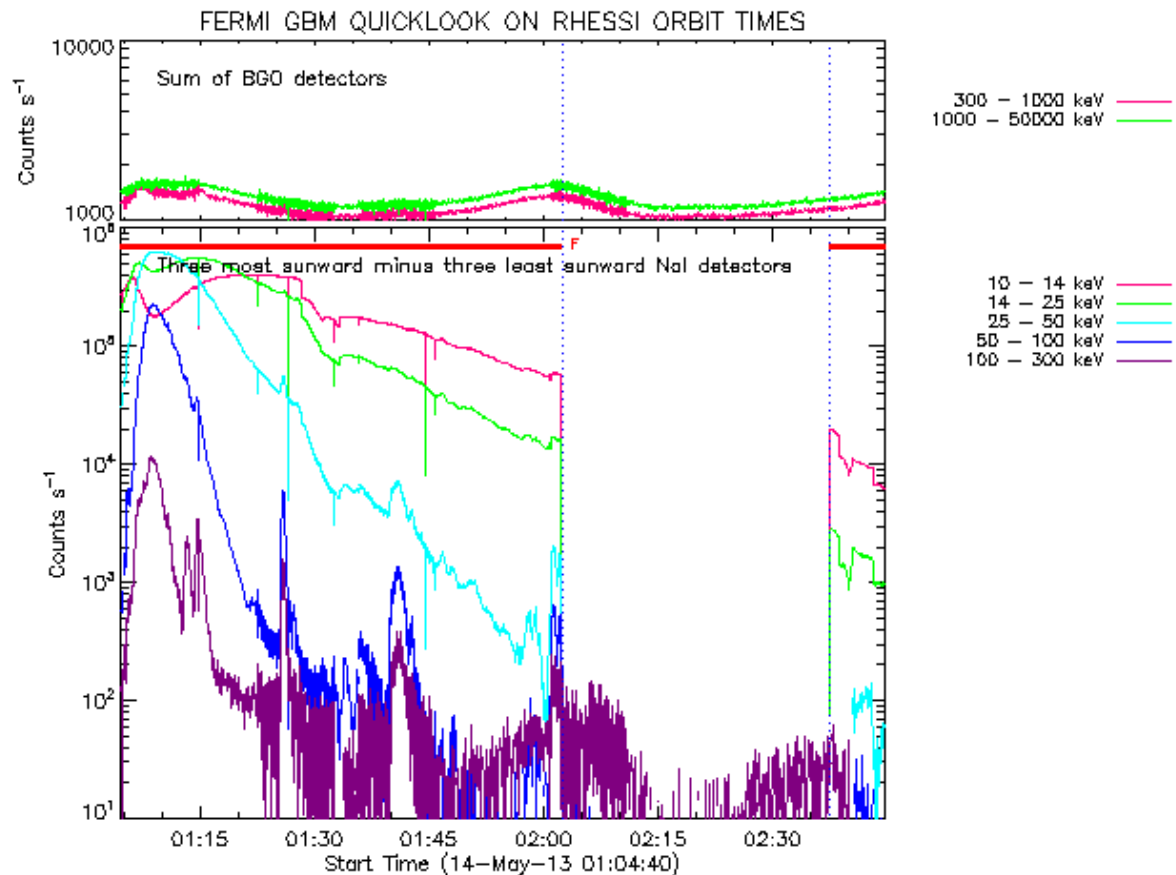
# Priest & Forbes (2002)



Trajectories for flux rope ( $h$ ), and upper ( $q$ ) and lower ( $p$ ) tips of the current sheet. Dashed line shows flux rope speed ( $\dot{h}$ ). Inflow Alfvén Mach number = 0.1 at centre of current sheet.

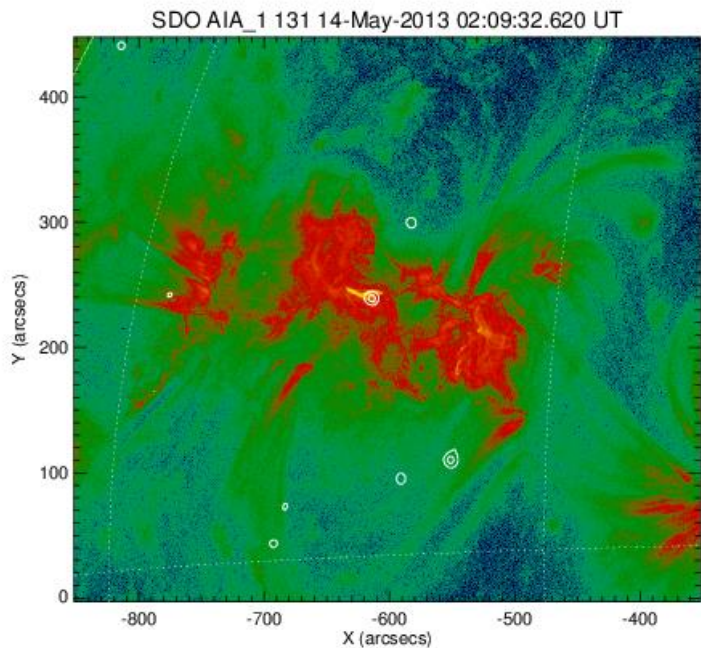
# 14 May 2013

## Fermi GBM 10 – 300 keV

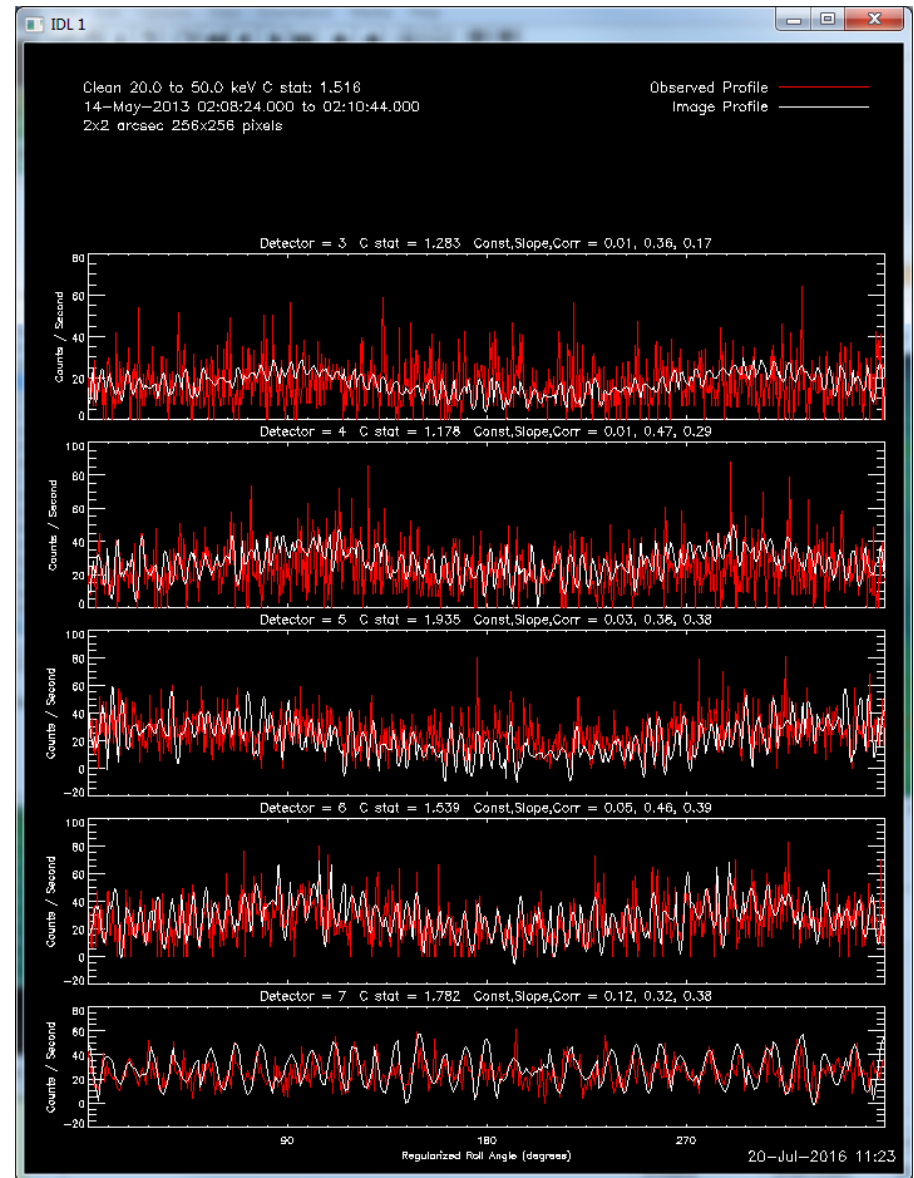


18-May-2013 10:47

# 14 May 2013 AR 1745

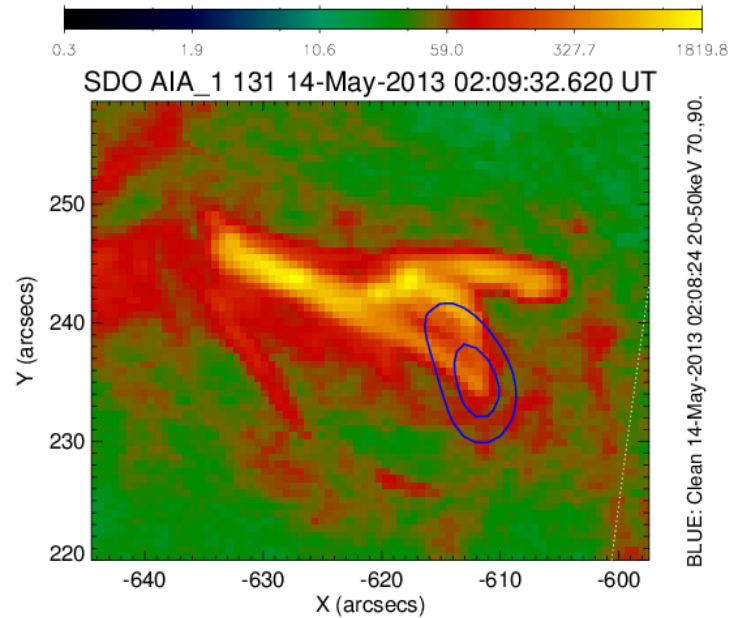
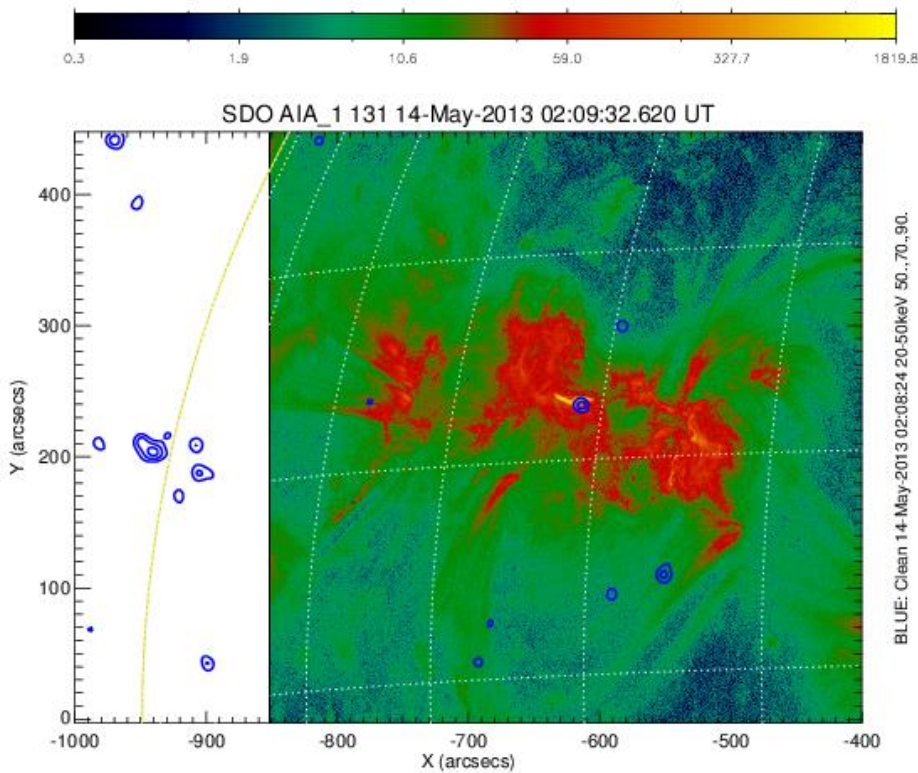


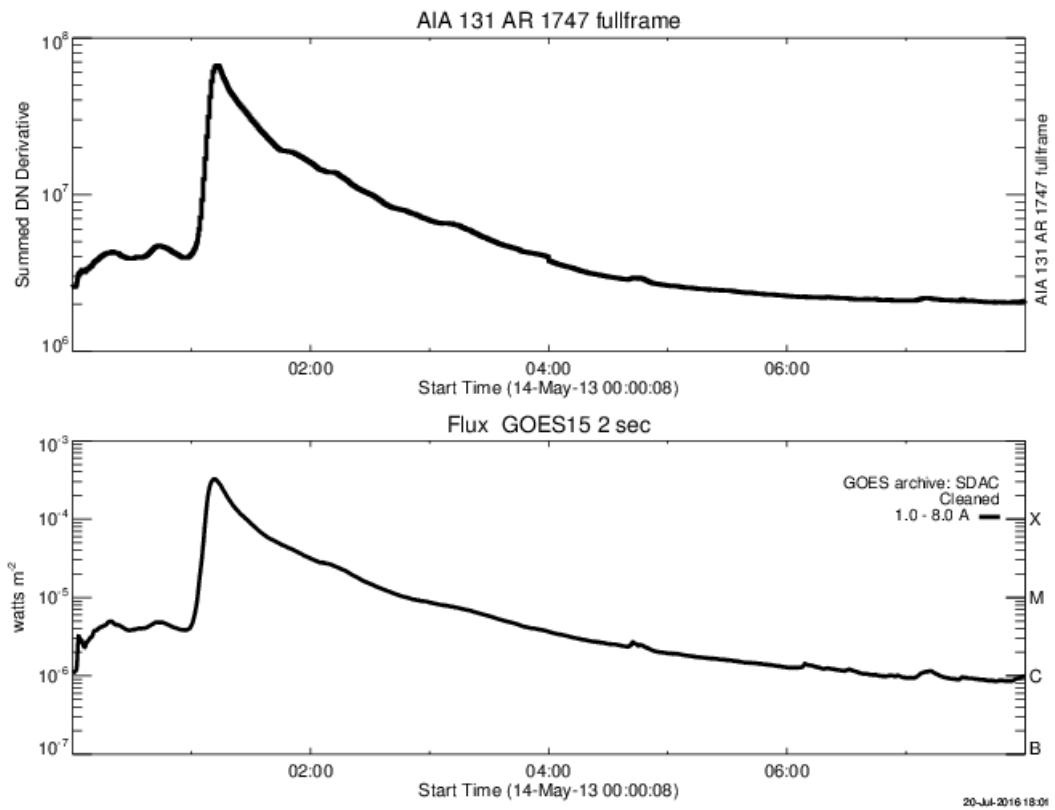
WHITE: Clean 14-May-2013 02:08:24 20-50keV 50.,70,90.



# 14 May 2013

## 20-50 keV flux from AR 1745 & 1747?





# HEK Flare

- concept: Flare
- eventtype: 9
- event\_c1error: 2
- event\_c2error: 2
- event\_coord1: -652.8
- event\_coord2: 268.8
- event\_coordsys: UTC-HPC-TOPO
- event\_coordunit: arcseconds
- Event End Time: 2013-05-14T01:53:10
- Event Peak Time: 2013-05-14T01:46:22
- event\_score: 0.00606790242831606753
- Event Start Time: 2013-05-14T01:45:10
- event\_testflag: false
- Type of Event: FL
- fl\_peakflux: 46.4379
- fl\_peakfluxunit: erg/cm/cm/s
- kb\_archivdate: 2013-05-14T02:10:21
- kb\_archivid: ivo://helio-informatics.org/FL\_FlareDetective-TriggerModule\_20130514\_020419\_2013-05-14T01:45:10.070\_1
- kb\_archivist: flare\_detective
- obs\_channelid: 131
- obs\_instrument: AIA
- obs\_meanwavel: 0.00000131
- obs\_observatory: SDO
- obs\_wavelunit: cm
- frm\_contact: Paolo C. Grigis - pgrigis@cfa.harvard.edu
- frm\_daterun: 2013-05-13T19:02:30
- frm\_humanflag: false
- frm\_identifier: Feature Finding Team
- frm\_institute: SAO
- frm\_name: Flare Detective - Trigger Module
- frm\_paramset: DerivativeThreshold= 8.00000e-02 EndFraction= 2.50000e-01
- frm\_url: <http://www.cfa.harvard.edu>
- frm\_versionnumber: 0.51