# Monte Carlo methods

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### 2020-2021, winter term



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# Overview

- **1** Examples: Ising model etc.
- Integration vs Sampling
- Beudo random number generators
- Generation of random numbers with a given given distribution
- 5 Percolation, Random Walks
- 6 Improtance Sampling, Markov chains, Metropolis Alg.
- 7 Statistics, error estimates: Jackknife, bootstrap
- 8 Langevin equation
- 9 fitting,  $\chi^2$  test
- **ID** Potts model, Ising model, XY model, nonlinear O(n) model

### Langevin equation

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We have an action with continous variables  $S[q_i]$ Introduce Langevin time (similar to MC time)  $q_i$  have an evolution as a function of  $\tau$  using the Langevin equation, stochastic differential equation.

$$dq_i = -\frac{\partial S}{\partial q_i} d\tau + \sqrt{2} dw_i(\tau) \tag{1}$$

 $dw_i$  is an increment of a Wiener process: random walk in the continuum

$$\langle w_i(\tau) \rangle = 0, \quad \langle (w_i(\tau') - w_i(\tau))^2 \rangle = (\tau' - \tau)$$
 (2)

Discretised with Ito calculus:

$$q_i(t + \Delta \tau) = q_i(\tau) + \left(-\frac{\partial S}{\partial q_i}\right) \Delta \tau + \eta_{i,\tau} \sqrt{2\Delta \tau}$$
(3)

with  $\langle \eta_{i,\tau} \rangle = 0$ ,  $\langle \eta_{i,\tau} \eta_{j,\tau'} \rangle = \delta_{ij} \delta_{\tau,\tau'}$ , Gaussian

### Langevin eq.

 $P(q_i, \tau)$  probability density of the variables If we have

$$P(q_i, \tau) \rightarrow_{\tau \to \infty} e^{-S(q_i)}$$
 (4)

than we can use this for simulating the thermal averages.

$$\int e^{-S(q)} F(q) d^n q = \lim_{T \to \infty} \frac{1}{T} \int_0^T F(q(\tau)) d\tau$$
(5)

Averages collected along the trajectories.  $P(q_i, \tau) \rightarrow_{\tau \rightarrow \infty} e^{-S(q_i)}$  can be shown using the **Fokker-Planck eq.** 

$$\frac{\partial P(q_i,\tau)}{\partial \tau} = \sum_i \frac{\partial}{\partial q_i} \left( \frac{\partial}{\partial q_i} + \frac{\partial S}{\partial q_i} \right) P(q_i,\tau)$$
(6)

(Stationarity of  $e^{-S(q_i)}$  is easy to show, whether *P* converges to there is a bit harder to show)

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### deriving the Fokker-Planck eq.

We need  $\partial_{\tau} P(q_i, \tau)$ .

$$P(q'_i, \tau + \Delta \tau) = \left\langle \int dq_i P(q_i, \tau) \prod_i \delta \left( q'_i - q_i + \delta \tau \frac{\partial S}{\partial q_i} - \eta_i \sqrt{2\Delta \tau} \right) \right\rangle_{\eta}$$
(7)

for a constant c we have  $\int \delta(cx)dx = \int \delta(cx)d(cx)/c = 1/c$  This implies  $\delta(cx) = (1/|c|)\delta(x)$  From that we have the rule:

$$\delta(f(x)) = \sum_{f(x_i)=0} \frac{\delta(x - x_i)}{|f'(x_i)|}$$
(8)

Now we can evaluate the formula above:

$$P(q'_{i},\tau+\Delta\tau) = \left\langle P(q'_{i}+\Delta\tau\frac{\partial S}{\partial q_{i}}-\eta_{i}\sqrt{2\Delta\tau},\tau)\frac{1}{\prod_{i}\left|-1+\Delta\tau\frac{\partial^{2}S}{\partial q_{i}^{2}}\right|}\right\rangle_{\eta}$$
(9)

We have

$$\frac{1}{\prod_{i} \left| -1 + \Delta \tau \frac{\partial^{2} S}{\partial q_{i}^{2}} \right|} = \frac{1}{\prod_{i} \left( 1 - \Delta \tau \frac{\partial^{2} S}{\partial q_{i}^{2}} \right)} \approx 1 + \sum_{i} \Delta \tau \frac{\partial^{2} S}{\partial q_{i}^{2}}$$
(10)

## deriving Fokker-Planck eq. 2

Now we taylor expand

$$P(q_i', \tau + \Delta \tau) = \left\langle P(q_i' + \Delta \tau \frac{\partial S}{\partial q_i} - \eta_i \sqrt{2\Delta \tau}, \tau) \frac{1}{\prod_i \left| -1 + \Delta \tau \frac{\partial^2 S}{\partial q_i^2} \right|} \right\rangle_{\eta}$$
(11)

to order  $\Delta \tau$  we have:

$$\begin{split} P(q'_i, \tau + \Delta \tau) &= \left\langle P(q'_i, \tau) + \sum_i \partial_i P(q'_i, \tau) \Delta \tau \partial_i S - \sum_i \partial_i P(q'_i, \tau) \eta_i \sqrt{2\Delta \tau} \right. \\ &+ \frac{1}{2} \sum_{i,j} \partial_i \partial_j P(q'_i, \tau) \eta_i \eta_j 2\Delta \tau + \sum_i P(q'_i, \tau) \Delta \tau \partial_i^2 S \right\rangle_{\eta} \end{split}$$

Now using  $\langle \eta_i \rangle = 0$ ,  $\langle \eta_i \eta_j \rangle = \delta_{ij}$ 

$$P(\tau + \Delta \tau) = P + \Delta \tau \left( \sum_{i} \partial_{i} P \partial_{i} S + \sum_{i} \partial_{i}^{2} P + \sum_{i} P \partial_{i}^{2} S \right)$$
(12)

Which finally gives the Fokker-Planck equation:

$$\frac{\partial P(q_i,\tau)}{\partial \tau} = \sum_{i} \frac{\partial}{\partial q_i} \left( \frac{\partial}{\partial q_i} + \frac{\partial S}{\partial q_i} \right) P(q_i,\tau)$$
(13)

### Convergence

Here we show that using the Fokker-Planck eq.  $\partial_{\tau} P = \sum_{i} \partial_{i} (\partial_{i} + \partial_{i} S) P$   $\implies P$  converges to  $e^{-S}$ . Let's define  $\Psi(q_{i}, \tau)$  such that

$$P(q_i,\tau) = \Psi(q_i,\tau)e^{-S(q_i)/2}$$
(14)

Writing this ansatz into Fokker Planck:

$$\begin{split} \dot{\Psi}e^{-S/2} &= \partial_i(\Psi' - \frac{1}{2}S'\Psi + S'\Psi)e^{-S/2} = \partial_i(\Psi' + \frac{1}{2}S'\Psi)e^{-S/2} = \\ &= \left(\Psi'' + \frac{1}{2}S''\Psi + \underbrace{\frac{1}{2}S'\Psi' - \frac{1}{2}S'\Psi'}_{=0} - \frac{1}{4}(S')^2\Psi\right)e^{-S/2} \end{split}$$

Now we note

$$-(-\partial_{i} + S'/2)(\partial_{i} + S'/2)\Psi = -(-\partial_{i} + S'/2)(\Psi' + S'\Psi/2)$$
(15)  
= -(-\Psi'' - S''\Psi'' - S'\Psi'' - S'

### Convergence 2

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So we can write

$$\dot{\Psi} = -H_{FP}\Psi, \quad H_{FP} = Q^+Q$$
 (16)

Where we defined the Fokker-Planck Hamiltonian  $H_{FP}$ , which we see is semi-positive definite We have one stationary solution to  $Q\Psi = 0$ , which is  $\Psi = e^{-S/2}$  (It's unique except for some special cases) We can write

$$\Psi(\tau) = c_0 e^{-S/2} + \sum_i e^{-\lambda_i \tau} \Psi_i$$
(17)

with  $\lambda_i > 0$  for all *i*.

 $\implies \Psi \to e^{-S/2}$   $P(\tau) = \Psi(\tau)e^{-S/2} \to e^{-S}$ (18)

### Langevin eq. example of a toy model

We are given the action  $S(x) = ax^2 + bx^4$ Suppose we are interested in

$$\langle x^2 \rangle = \frac{\int_{-\infty}^{\infty} dx x^2 e^{-ax^2 - bx^4}}{\int_{-\infty}^{\infty} dx e^{-ax^2 - bx^4}}$$
(19)

Calculate the drift term:

$$K(x) = -\frac{\partial S}{\partial x} = -2ax - 4bx^3$$
<sup>(20)</sup>

The discretised Langevin equation is thus:

$$x(\tau + \Delta \tau) = x(\tau) + \Delta \tau (-2ax - 4bx^3) + \sqrt{2\Delta \tau} \eta(\tau)$$
(21)

We choose some  $\Delta \tau$ , and an initial  $x(\tau = 0)$ 

We solve the Langevin eq. numerically using an independent Gaussian random number at every timestep.

### Toy model example 2

We thus have  $x(\tau)$  for  $\tau = n\Delta \tau$ .

now we use

$$\lim_{T \to \infty} \frac{1}{T - T_0} \int_{T_0}^T f(x(\tau)) d\tau = \int dx \Pi(x) f(x)$$
(22)

with  $\Pi(x) \sim e^{-S(x)}$  the normalized equilibrium distribution.  $T_0 > 0$  is used to get rid of the initial thermalization of the process, and thus get rid of the dependence on the initial conditions (similarly to Metropolis simulations). We thus use

$$\langle f(x) \rangle \to \frac{1}{N} \sum_{\tau > \tau_0} f(x(\tau))$$
 (23)

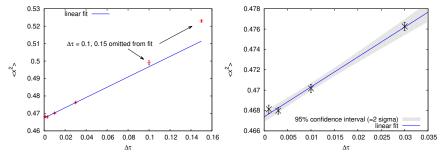
where N is the number of  $x(\tau)$  steps we have with  $\tau > \tau_0$ .

Numerically we use some  $\Delta \tau > 0$ , we have to carry out the extrapolation to  $\Delta \tau = 0$ . For the naive discretisation above the average observables are expected to have a linear dependence on  $\Delta \tau$ . (Higher order discretisations improve this)

 $\rightarrow$  measure at 3 different  $\Delta \tau$  values, fit linear to  $\langle f(x) \rangle_{\Delta \tau}$  and extrapolate to  $\Delta \tau = 0$  (i.e. read of the intercept)

### Continuum extrapolation

We need  $\langle x^2 \rangle_{\Delta \tau}$  at  $\Delta \tau \to 0$ For the naive discretisation, a linear dependence is expected, so we fit:



Large  $\Delta \tau$  values are omitted as linear behavior sets in at small  $\Delta \tau$ . How do we calculate the confidence interval of the fit?

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Let's look at fitting a linear to the data (Linear regression)

We have a dataset:  $\{x_i, y_i, \delta_i\}$  where we have data point  $y_i$  with 1  $\sigma$  error  $\delta_i$  at coordinate  $x_i$  (Assuming no errors in  $x_i$ ). We are looking for a and b such that a + b \* x best describes the dataset.

What is the error on *a* and *b*, what is their correlation?

Define the best fit by the minimal  $\chi^2$ :

$$\chi^2 = \sum_{i=1}^{N} \left( \frac{y_i - \mathbf{a} - bx_i}{\delta_i} \right)^2 \tag{24}$$

Points with smaller errors are important, points with large errors are not. Minimum:

$$0 = \frac{\partial \chi^2}{\partial a} = -2 \sum_{i=1}^{N} \frac{y_i - a - bx_i}{\delta_i^2}$$

$$0 = \frac{\partial \chi^2}{\partial b} = -2 \sum_{i=1}^{N} \frac{x_i (y_i - a - bx_i)}{\delta_i^2}$$
(25)

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### Linear fit

#### Using

$$S = \sum \frac{1}{\delta_i^2}, \quad S_x = \sum \frac{x_i}{\delta_i^2}, \quad S_{xx} = \sum \frac{x_i^2}{\delta_i^2}, \quad S_y = \sum \frac{y_i}{\delta_i^2}, \quad S_{xy} = \sum \frac{x_i y_i}{\delta_i^2},$$

and  $\Delta = SS_{xx} - S_x^2$ , we have:

$$aS + bS_x = S_y, \qquad \Longrightarrow \qquad a = \frac{S_{xx}S_y - S_xS_{xy}}{b} = \frac{S_{xy} - S_xS_y}{\Delta}$$

$$b = \frac{SS_{xy} - S_xS_y}{\Delta}$$

$$(26)$$

Using error propagation (assuming independent  $y_i$  variables):

$$\delta_{a}^{2} = \sum_{i} \left(\frac{\partial a}{\partial y_{i}}\right)^{2} \delta_{i}^{2}, \quad \delta_{b}^{2} = \sum_{i} \left(\frac{\partial b}{\partial y_{i}}\right)^{2} \delta_{i}^{2}$$
(27)

Covariance:

$$\operatorname{Cov}(\mathbf{a}, \mathbf{b}) = \langle \mathbf{a} \mathbf{b} \rangle - \langle \mathbf{a} \rangle \langle \mathbf{b} \rangle = \sum_{i} \frac{\partial \mathbf{a}}{\partial y_{i}} \frac{\partial \mathbf{b}}{\partial y_{i}} \delta_{i}^{2}$$
(28)

## Linear fit errors

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$$\left(\frac{\partial a}{\partial y_i}\right)^2 = \left(\frac{S_{xx}\frac{1}{\delta_i^2} - S_x\frac{x_i}{\delta_i^2}}{\Delta}\right)^2 = \frac{S_{xx}^2 - 2S_{xx}S_xx_i + S_x^2x_i^2}{\Delta\delta_i^4}$$
(29)  
$$\delta_a^2 = \sum_i \left(\frac{\partial a}{\partial y_i}\right)^2 \delta_i^2 = \frac{S_{xx}^2 S - 2S_{xx}S_x^2 + S_x^2S_{xx}}{\Delta^2} = \frac{S_{xx}(S_{xx}S - S_x^2)}{\Delta^2} = \frac{S_{xx}}{\Delta}$$

Similarly we can calculate:

$$\delta_b^2 = \frac{S}{\Delta}, \quad \operatorname{Cov}(a, b) = -\frac{S_x}{\Delta}$$
 (30)

Now we can get errors for any x:

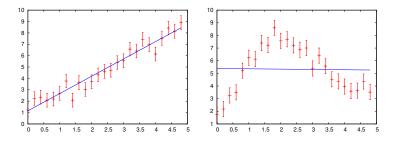
$$\delta^{2}(a+bx) = \langle (a+bx)^{2} \rangle - \langle a+bx \rangle^{2} = \delta^{2}_{a} + \delta^{2}_{b}x^{2} + 2x \operatorname{Cov}(a,b)$$
(31)

Fit quality

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### Which fit makes sense?



Besides looking at the plot (which is also important),

How can we decide whether the model (the function we fit) is probably OK or not?

 $\chi^2$  test

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We want to see if the data we get is described with the function y = f(x). We use

$$\chi^2 = \sum_{i=1}^{N} \left( \frac{y_i - f(x_i)}{\delta_i} \right)^2 \tag{32}$$

if the model is correct we get  $y_i$  from a Gaussian distribution

$$P(y_i) \sim e^{-\frac{y_i - f(x_i)^2}{2\delta_i^2}}, \quad x_i = \frac{y_i - f(x_i)}{\delta_i} \implies P(x_i) \sim e^{-x_i^2/2}$$
 (33)

So  $\chi^2$  is a sum of N Gaussian random variables with zero mean and  $\sigma = 1$ . Probability for haveing  $\chi^2 > C^2$ 

$$Q(C^{2}) = \frac{1}{(2\pi)^{N/2}} \int \prod dx_{i} e^{-\sum x_{i}^{2}/2} \Theta\left(\left(\sum x_{i}^{2}\right) - C^{2}\right)$$
(34)

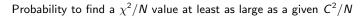
Using polar coordinates and then  $\sum x_i^2/2 = t$ , we have

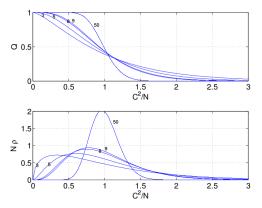
$$1 - Q(C^2) = \frac{|S_{N-1}|}{(2\pi)^{N/2}} \int_0^{C^2/2} t^{N/2-1} e^{-t} = \Gamma_{inc}(C^2/2, N/2)$$
(35)

with the incomplete Gamma function.

Q-value

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small Q means :

- model is wrong
- errors are underestimated

errors are not really Gaussian, but have long tail Smoothed Step function – for large N gets steeper Standard deviation of  $\chi^2/N$  is  $\sqrt{2/N}$ 

## Fitting procedure

Usually we have some free parameters that we want to get from the data:

- **1** *N* data points  $x_i, y_i, \delta_i$
- **2** fit function:  $f(x_i, a_j)$  with  $\nu$  free parameters

#### 3

$$\chi^2 = \sum_{1}^{N} \left( \frac{y_i - f(x_i, a_j)}{\delta_i^2} \right)^2$$
(36)

- Minimize χ<sup>2</sup> with respect to a<sub>j</sub>. (calculate errors for a<sub>j</sub> with e.g. Bootstrap, Jackknife, error propagation in simple cases)
- **5**  $\chi^2$  test: calculate  $\chi^2/N_{dof}$  with  $N_{dof} = (N \nu)$ if  $\chi^2/N_{dof} \gg 1 \implies$ : model is probably wrong (or other problems, as above) if  $\chi^2/N_{dof} \ll 1 \implies$ : This is called overfitting. Noise too large to say anything, or errors are overestimated.

### General linear Least Squares

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Sometimes a linear fit cannot suffice, we want  $f(x) = \sum a_k X_k(x)$  for e.g. a polynomial fit

$$\chi^{2} = \sum \left(\frac{y_{i} - \sum a_{k} X_{k}(x_{i})}{\sigma_{i}}\right)^{2}$$
(37)

The minimum condition leads to  $\sum \alpha_{kj} \mathbf{a}_j = \beta_k$  with

$$\alpha_{kj} = \sum_{i=1}^{N} \frac{X_j(x_i) X_k(x_i)}{\sigma_i^2}, \quad \beta_k = \sum_{i=1}^{N} \frac{y_i X_k(x_i)}{\sigma_i^2},$$
(38)

The errors and covariances are calculated from  $C = \alpha^{-1}$ 

$$\delta^{2}(a_{i}) = C_{ii}, \quad \operatorname{Cov}(a_{i}, a_{j}) = C_{ij}$$
(39)

Typically LU decomposition, QR decomposition or similar is used to calculate the inverse.

Generally we have  $f(x_i, a_j)$  nonlinear function. We might be able to convert to a linear problem

$$y_i = \exp(-\lambda x_i) \quad \rightarrow \ln y_i = -\lambda x_i$$
 (40)

Or even:

$$y_i = \frac{1}{e^{\beta x_i} + 1} \quad \rightarrow \ln\left(\frac{1}{y_i} - 1\right) = -\lambda x_i \tag{41}$$

Use error propagation to calculate errors.

Finally for the non-linear problems that remain, one uses a non-linear minimum search to get the fit parameters (e.g. gradient descent method)

## Potts model

The *q*-state Potts model is defined by the energy

$$H[S] = -\alpha \sum_{x} \sum_{\nu} \delta(S_x, S_{x+\hat{\nu}}) - h \sum_{x} \delta(S_n, 1),$$
(42)

with the spins  $S_x \in \{1, 2, ..., q\}$ , where x points to a lattice site on a d dimensional lattice and  $x + \hat{\nu}$  is a neighbor in  $+\nu$  direction. q = 2 Potts model is is equivalent to the Ising model:  $\{1, 2\} \rightarrow \{-1, 1\}$ , and the  $H_{\text{Ising}} = H_{q=2\text{Potts}} + \text{Const}$ 

The symmetryes of the system play a central role in the type of phase transitions the system can have

At h = 0 Potts model is symmetric under permutation:

$$\begin{array}{l} 1 \rightarrow p(1) \\ 2 \rightarrow p(2) \\ \vdots \\ q \rightarrow p(q) \end{array} \quad \text{with } p(i) \text{ a permutation of the numbers } 1, 2, \dots q \end{array}$$

## Symmetries

p is an element of the permutation group  $S_q$ 

$$\delta(S_x, S_{x+\hat{\nu}}) = \delta(\rho(S_x), \rho(S_{x+\hat{\nu}}))$$
(43)

if  $h \neq 0$ , the symmetry is *explicity broken* by the magnetic field

#### Other symmetries:

The Ising model (for zero magnetic field is invariant under

$$S_i \to S_i, \quad S_i \to -S_i$$
 (44)

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as  $S_1S_2 = (-S_1)(-S_2)$ . Symmetry group:  $Z_2 = \{-1, 1\}$ , (which is equivalent to  $S_2$ ) Magnetic field again gives an explicit breaking

## Continous spin systems

### Nonlinear O(n) spin model

Consider *n* component spin vectors (with real components) which satisfy  $|S_i| = 1$ 

The action is given by:

$$S = -\beta \sum_{x} \sum_{\nu} S_{x}^{T} S_{x+\hat{\nu}} - h \sum_{x} S_{1}$$

$$(45)$$

Transform all spins with  $S_x \rightarrow MS_x$  with some  $n \times n$  matrix M.

$$S_x^T S_{x+\hat{\nu}} \to S_x^T M^T M S_{x+\hat{\nu}}$$
(46)

This is a symmetry transformation if  $M^T M = 1$ , i.e. for orthogonal matrices. The magnetic field term breaks the symmetry again.

**XY-model** n = 2 of the above is sometimes called the XY-model. (n = 3 is called Heisenberg model)

Writeing the spins in terms of an angle we have  $S = (\cos(\phi), \sin(\phi))$ 

$$S = -\beta \sum_{x} \sum_{\nu} \cos(\phi_x - \phi_{x+\hat{\nu}}) - h \sum_{x} \cos(\phi_x)$$
(47)

The symmetry group is now the group of rotations in 2d, described with one angle.

### Canonical Ensemble, observables

We want to calculate the properties of these models couples to a heat bath with temperature T.

Partition function using  $\beta = 1/(k_B T)$ :

$$Z = \sum_{\{S\}} e^{-\beta H[S]}, \quad \langle O \rangle = \frac{1}{Z} \sum_{\{S\}} O(S) e^{-\beta H[S]}$$
(48)

Where we have to carry out the sum for all configurations q-state Potts model has  $q^V$ ,  $V = L^d$  configurations Continuus models: sum  $\rightarrow$  integration We use the notation  $J = \alpha \beta$ ,  $M = h\beta$ .

#### Observables:

Free energy:  $F = -\frac{1}{\beta} \ln Z$ Internal energy:

$$E = U = \langle H \rangle = \frac{1}{Z} \sum_{\{S\}} H[S] e^{-\beta H[S]} = -\frac{1}{Z} \frac{\partial}{\partial \beta} \sum_{\{S\}} e^{-\beta H[S]} = -\frac{\partial}{\partial \beta} \ln Z = \frac{\partial}{\partial \beta} (\beta F)$$

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# Observables 2

#### Magnetization

Assuming the form  $H[S] = \alpha A[S] + hM[S]$ , with the magnetization M[S] and A giving the rest of the terms.

$$\langle M \rangle = \frac{1}{Z} \sum_{\{S\}} M[S] e^{-\beta H[S]} = -\frac{1}{\beta} \frac{1}{Z} \frac{\partial}{\partial h} \sum_{\{S\}} e^{-\beta H[S]} = -\frac{1}{\beta} \frac{\partial}{\partial h} \ln Z = \frac{1}{\beta} \frac{\partial}{\partial h} (\beta F)$$

**Specific heat** is defined as  $c_V = \partial_T E$ . First we convert to  $\beta$  derivative:

$$\beta = \frac{1}{k_B T} \implies d\beta = -\frac{1}{k_B T^2} dT = -k_B \beta^2 dT$$
(49)

$$\frac{\partial}{\partial\beta}\langle H\rangle = \frac{\partial}{\partial\beta} \left( \frac{1}{Z} \sum_{\{S\}} H[S] e^{-\beta H[S]} \right) =$$

$$= \frac{1}{Z^2} \left( \sum_{\{S\}} H[S] e^{-\beta H[S]} \right)^2 - \frac{1}{Z} \sum_{\{S\}} H[S] H[S] e^{-\beta H[S]} = -\langle H^2 \rangle + \langle H \rangle^2$$
(50)

This is the energy fluctuation \*(-1)

$$\frac{1}{k_{\mathcal{B}}\beta^{2}}c_{\mathcal{V}} = -\frac{\partial}{\partial\beta}\langle H \rangle = \langle H^{2} \rangle - \langle H \rangle = -\frac{\partial}{\partial\beta}\left(-\frac{\partial}{\partial\beta}\ln Z\right) = -\frac{\partial^{2}}{\partial\beta^{2}}\left(\beta F\right) \quad (51)$$

### Observables 3

Magnetic Susceptibility Defined as:

$$\chi = \langle (M - \langle M \rangle)^2 \rangle = \langle M^2 \rangle - \langle M \rangle^2$$
(52)

Can also be calculated starting from:

$$\frac{\partial}{\partial h} \left( \frac{1}{Z} \sum_{\{S\}} M[S] e^{-\beta H[S]} \right) = -\frac{-\beta}{Z^2} \left( \sum_{\{S\}} M[S] e^{-\beta H[S]} \right)^2 + \frac{-\beta}{Z} \sum_{\{S\}} M[S]^2 e^{-\beta H[S]} = -\beta \left( \langle M^2 \rangle - \langle M \rangle^2 \right)$$

So we have

$$\chi = -\frac{1}{\beta} \frac{\partial}{\partial h} \langle M \rangle = \frac{1}{\beta^2} \frac{\partial^2}{\partial h^2} \ln Z = -\frac{1}{\beta^2} \frac{\partial^2}{\partial h^2} (\beta F)$$
(53)

**Generally:** (one usually divides with the volume  $\Omega$  for extensive quantities)

$$Z = \sum e^{-\beta(A[S]+jC[S])}$$
(54)

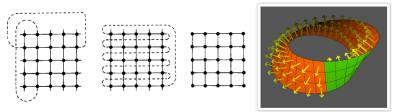
$$\langle C[S] \rangle / \Omega = -\frac{1}{\Omega \beta} \frac{\partial}{\partial j} \ln Z$$
 (55)

 $\chi_{C} = \Omega\left(\left\langle \left(C/\Omega\right)^{2}\right\rangle - \left\langle C/\Omega\right\rangle^{2}\right) = \frac{1}{\Omega\beta^{2}}\frac{\partial^{2}}{\partial j^{2}}\ln Z$ 

susceptibility:

average:

## Boundary conditions



On the computer we have a finite lattice. What happens at the edges?

- Periodic: makes the lattice a d dimensional torus. no boundary effect, but still finite size effects
- screw-periodic (aka. skew-periodic),free: easy to implement, some systematic effect always there, only gone at infinite volume limit
- antiperiodic: as periodic, but spin across the boundary gets a minus sign. This allows studying the interface between phases (can control the direction of the interface if combined with screw periodic)
- Möbius: Can be used to study topological effects

### Monte-Carlo simulations - Metropolis

We calculate averages using Markov chains with local Metropolis update.

- Choose boundary conditions
- Choose initial configuration: Cold start: all spins  $S_x = 1$ Hot Start: Independent random Spins.
- Monte Carlo "hit": a proposed update for one of the spins. One hits all spins either sequentially or chooses a random lattice point for each hit
- "Sweep": All lattice points were hit. (Or do  $\Omega$  hits)
- (if applicable) With some experimentation try to find proposals which lead to an acceptance rate of 0.5-0.8
- Before measurements one needs typically  $O(10^2) O(10^5)$  sweeps to get rid of the effect of the initial conditions (e.g. compare hot and cold starts)
- $\blacksquare$  Between two measurements one does 10-100 sweeps to decrease autocorrelations on the data
- Analysis of the data: calculate averages, errors, correlations, do fits

We can also calculate averages using the Langevin equation, in case we are looking at a system with continous variables

- Choose boundary conditions
- Choose initial configuration: Cold start: all spins S<sub>x</sub> = 1 Hot Start: Independent random Spins.
- Langevin step: One update of the fields with the discretised Langevin eq. with stepsize  $\Delta \tau$ , increase Langevin time  $\tau$  to  $\tau + \Delta \tau$
- Before measurements one needs typically O(10) O(1000) Langeevin time for thermalization to get rid of the effect of the initial conditions (e.g. compare hot and cold starts)
- Between two measurements one waits 0.01 1 Langevin time
- Analysis of the data: calculate averages, errors, correlations, do fits
- Carry out  $\Delta au 
  ightarrow 0$  extrapolation

### Lattice in the computer

Representing fields on the lattice, one has to decide on the "layout". From the coordinates  $0 \le n_i < L_i$  we calculate *n* the lattice index

$$n_1, n_2, \ldots, n_d \to n = n_1 + n_2 L_1 + n_3 L_1 L_2 + \ldots$$
 (56)

which means  $0 \le n < \Omega$  with  $\Omega = \prod_i L_i$ , with a one to one relation  $\{n_i\} \leftrightarrow n$ One can use more complex functions to improve performance on a given computer architecture

We need coordinates for neighbors:

$$\begin{array}{l} n_i^{(+\nu)} = n_i(x) \\ n_\nu^{(+\nu)} + = 1 \\ \text{if } n_\nu^{(+\nu)} = L_\nu : \quad n_\nu^{(+\nu)} = 0 \\ n_i^{(+\nu)} \to n^{(+\nu)} \end{array}$$

We either store  $n^{(\pm\nu)}$  in the memory ("neighbor map") or calculate on the fly when needed, depending on which is faster on a given architecture.